





The QCD phase diagrams, nuclear matter & neutron stars at large Nc

Francesco Giacosa

in collaboration with A. heinz, D. Rischke, A. Bonanno, P. Kovacs, Gy. Kovacs, G. Pagliara

UJK Kielce (Poland) & Goethe U Frankfurt (Germany)

63. Cracow School of Theoretical Physics

Nuclear Matter at Extreme Densities and High Temperatures 17-23/6/2023 Zakopane Poland

Outline



- QCD: brief review
- Large-Nc: what is that
- Basic properties
- Large-Nc at nonzero temperature
- QCD Phase-diagram at large Nc
- Nuclear matter at large-Nc
- Neutrons stars at large-Nc
- Conclusions



Symmetries of QCD



Giuseppe Lodovico Lagrangia

25 January 1736

Turin

Born

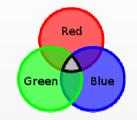
Died 10 April 1813 (aged 77)

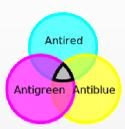
Paris

The QCD Lagrangian



Quark: u,d,s and c,b,t R,G,B



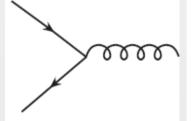


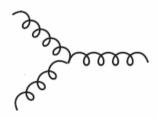
$$q_i = egin{pmatrix} q_i^R \ q_i^G \ q_i^B \end{pmatrix}; \ i = u,d,s ,...$$
 $\mathcal{L}_{QCD} = \sum_{i=1}^{N_f} \overline{q}_i (i \gamma^\mu D_\mu - m_i) q_i - rac{1}{4} G_{\mu
u}^a G^{a,\mu
u}$ e of gluons (RG,BG,...)

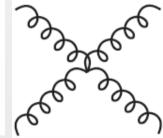
$$\mathcal{L}_{QCD} = \sum_{i=1}^{N_f} \overline{q}_i (i\gamma^{\mu} D_{\mu} - m_i) q_i - \frac{1}{4} G^a_{\mu\nu} G^{a,\mu\nu}$$

8 type of gluons (RG,BG,...)

$$A_{\mu}^{a}$$
; $a = 1,..., 8$

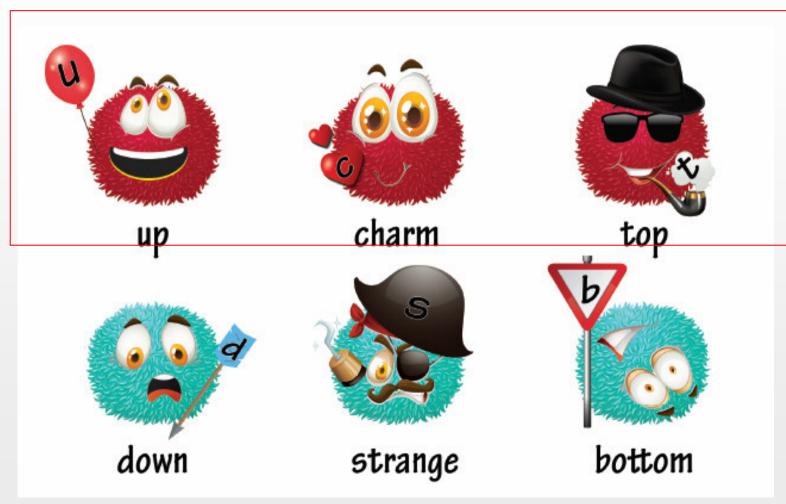






Confinement: quarks never 'seen' directly. How they might look like ©





Picture by Pawel Piotrowski

Trace anomaly: the emergence of a dimension



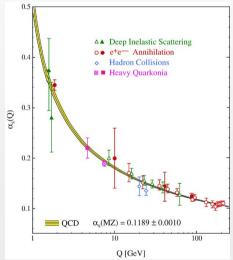
Chiral limit: $m_{\cdot} = 0$

$$x^{\mu} \rightarrow x'^{\mu} = \lambda^{-1} x^{\mu}$$

 $x^{\mu} \to x'^{\mu} = \lambda^{-1} x^{\mu}$ is a classical symmetry broken by quantum fluctuations (trace anomaly)

Dimensional transmutation
$$\Lambda_{YM} \approx 250 \text{ M eV}$$

$$\alpha_{\rm S}(\mu=Q) = \frac{g^2(Q)}{4\pi}$$

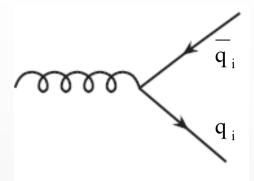


Effective gluon mass: $m_{\rm gluon} = 0 \rightarrow m_{\rm gluon}^* \approx 500 - 800 \, {\rm MeV}$

Gluon condensate: $\left\langle G_{\mu\nu}^{a}G^{a,\mu\nu}\right\rangle \neq 0$

Flavor symmetry





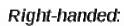
Gluon-quark-antiquark vertex

It is democratic! The gluon couples to each flavor with the same strength

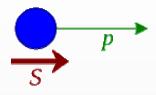
$$q_{\scriptscriptstyle i} \to U_{\scriptscriptstyle ij} q_{\scriptscriptstyle j}$$

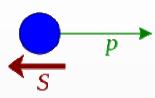
$$U \in U(3)_V \rightarrow U^+U = 1$$

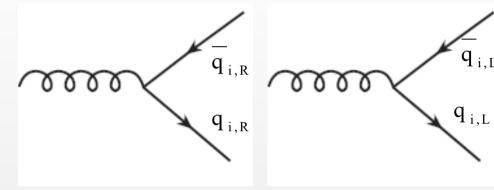
Chiral symmetry













$$q_{i} = q_{i,R} + q_{i,L}$$

$$q_{i,R} = \frac{1}{2} (1 + \gamma^{5}) q_{i}$$

$$q_{i,L} = \frac{1}{2} (1 - \gamma^{5}) q_{i}$$

$$q_{i} = q_{i,R} + q_{i,L} \rightarrow U_{ij}^{R} q_{j,R} + U_{ij}^{L} q_{j,L}$$

$$U(3)_{R} \times U(3)_{L} = U(1)_{R+L} \times U(1)_{R-L} \times SU(3)_{R} \times SU(3)_{L}$$

baryon number

anomaly U(1)A

SSB into SU(3)V

Chiral (or axial) anomaly: explicitely broken by quantum fluctuations

$$\partial^{\mu}(\bar{q}^{i}\gamma_{\mu}\gamma_{5}q^{i}) = \frac{3g^{2}}{16\pi^{2}} \varepsilon^{\mu\nu\rho\sigma} \operatorname{tr}(G_{\mu\nu}G_{\rho\sigma})$$

In the chiral limit (mi=0) chiral symmetry is exact, but is spontaneously broken by the QCD vacuum

Spontaneous breaking of chiral symmetry: chiral condensate and constituent mass



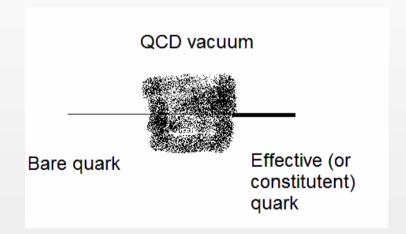
$$U(3)_R \times U(3)_L = U(1)_{R+L} \times U(1)_{R-L} \times SU(3)_R \times SU(3)_L$$

SSB:
$$SU(3)_R \times SU(3)_L \rightarrow SU(3)_{V=R+L}$$

Chiral symmetry → Flavor symmetry

$$\left\langle \overrightarrow{q}_{i}q_{i}\right\rangle = \left\langle \overrightarrow{q}_{i,R}q_{i,L} + \overrightarrow{q}_{i,L}q_{i,R}\right\rangle \neq 0$$

$$m \approx m_{u} \approx m_{d} \approx 5 \text{ MeV} \rightarrow m^{*} \approx 300 \text{ MeV}$$



$$m_{\rho-meson} \approx 2m^*$$
 $m_{proton} \approx 3m^*$

At nonzero T the chiral condensate decreases

Symmetries of QCD and breakings



SU(3)color: exact. Confinement: you never see color, but only white states.

Dilatation invariance: holds only at a classical level and in the chiral limit.

Broken by quantum fluctuations (**scale anomaly**)

and by quark masses.

SU(3)R**xSU(3)**L: holds in the chiral limit, but is broken by nonzero quark

masses. Moreover, it is **spontaneously** broken to U(3)v=R+L

U(1)_{A=R-L}: holds at a classical level, but is also broken by quantum

fluctuations (chiral anomaly)

Hadrons



The QCD Lagrangian contains 'colored' quarks and gluons. However, no ,colored' state has been seen.

Confinement: physical states are "white" and are called hadrons.

Hadrons can be:

Mesons: bosonic hadrons

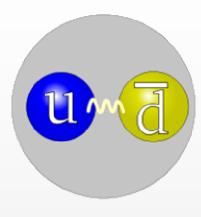
Baryons: fermionic hadrons

A meson is **not necessarily** a quark-antiquark state.

A quark-antiquark state is a conventional meson.

Example of conventional quark-antiquark states: the ρ and the π mesons





Rho-meson

$$m_{\rho^{+}} = 775 \text{ MeV}$$

$$\left|
ho^+
ight> \propto \left| u ar{d}
ight> + rac{1}{N_c} \left(\left| \pi^+ \pi^0
ight> + \ldots
ight)$$

where

$$\left|u\bar{d}\right\rangle = \left|\text{valence }u + \text{valence }\bar{d} + \text{gluons}\right\rangle$$

Pion

$$m_{\pi^{+}} = 139 \text{ MeV}$$

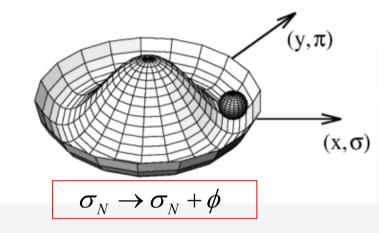
$$m_u + m_d \approx 7 \text{ MeV}$$

Mass generation in QCD is a nonpert. penomenon based on SSB

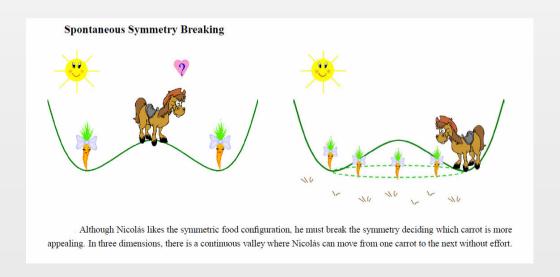
(mentioned previusly).

SSB and the donkey of Buridan: hadronic approaches





Jean Buridan (in Latin, Johannes Buridanus) (ca. 1300 – after 1358)



Quark-antiquark mesons (PDG 2018)

Uniwersytet

$n^{2s+1}\ell_J$	J^{PC}	$I = 1$ $u\overline{d}, \overline{u}d, \frac{1}{\sqrt{2}}(d\overline{d} - u\overline{u})$	$ \begin{aligned} \mathbf{I} &= \frac{1}{2} \\ u\overline{s}, \ d\overline{s}; \ \overline{ds}, \ -\overline{u}s \end{aligned} $	I = 0 f'	I = 0 f	$\theta_{ ext{quad}}$	$ heta_{ m lin}$ [°]
1 ¹ S ₀	0-+	π	K	η	$\eta'(958)$	-11.3	-24.5
1 ³ S ₁	1	$\rho(770)$	$K^*(892)$	$\phi(1020)$	$\omega(782)$	39.2	36.5
1 ¹ P ₁	1+-	$b_1(1235)$	K_{1B}^{\dagger}	$h_1(1380)$	$h_1(1170)$		
$1 {}^{3}P_{0}$	0++	$a_0(1450)$	$K_0^*(1430)$	$f_0(1710)$	$f_0(1370)$		
$1\ ^{3}P_{1}$	1++	$a_1(1260)$	K_{1A}^{\dagger}	$f_1(1420)$	$f_1(1285)$		
1 ³ P ₂	2++	$a_2(1320)$	$K_2^*(1430)$	$f_2^\prime(1525)$	$f_2(1270)$	29.6	28.0
$1\ ^{1}D_{2}$	2-+	$\pi_2(1670)$	$K_2(1770)^\dagger$	$\eta_2(1870)$	$\eta_2(1645)$		
$1\ ^{3}D_{1}$	1	ho(1700)	$K^*(1680)$		$\omega(1650)$		
$1~^3D_2$	2		$K_2(1820)$				
1 ³ D ₃	3	$ ho_3(1690)$	$K_3^*(1780)$	$\phi_{3}(1850)$	$\omega_3(1670)$	31.8	30.8
$1~^3F_4$	4++	$a_4(2040)$	$K_4^*(2045)$		$f_4(2050)$		
$1~^3G_5$	5	$\rho_5(2350)$	$K_5^*(2380)$				
$1\ ^{3}H_{6}$	6++	$a_6(2450)$			$f_6(2510)$		
2 ¹ S ₀	0-+	$\pi(1300)$	K(1460)	$\eta(1475)$	$\eta(1295)$		
$2\ ^{3}S_{1}$	1	ho(1450)	K*(1410)	$\phi(1680)$	$\omega(1420)$		
3 ¹ S ₀	0-+	$\pi(1800)$			$\eta(1760)$		

Some selected nonets



$n^{2S+1}L_J$	J^{PC}	$I=1$ $u\overline{d}, d\overline{u}$ $\frac{d\overline{d}-u\overline{u}}{\sqrt{2}}$	$I=1/2$ $u\overline{s}, d\overline{s}$ $s\overline{d}, s\overline{u}$	$I=0$ $\approx \frac{u\overline{u}+d\overline{d}}{\sqrt{2}}$	$ I=0 \approx s\overline{s}$	Meson names	Chiral Partners
$1^{1}S_{0}$	0-+	π	K	$\eta(547)$	$\eta'(958)$	Pseudoscalar	J=0
$1^{3}P_{0}$	0++	$a_0(1450)$	$K_0^{\star}(1430)$	$f_0(1370)$	$f_0(1500)/f_0(1710)$	Scalar	J=0
$1^{3}S_{1}$	1	$\rho(770)$	$K^{\star}(892)$	$\omega(782)$	$\phi(1020)$	Vector	J=1
$1^{3}P_{1}$	1++	$a_1(1260)$	K_{1A}	$f_1(1285)$	$f_1'(1420)$	Axial-vector	J=1
$1^{1}P_{1}$	1+-	$b_1(1235)$	K_{1B}	$h_1(1170)$	$h_1(1415)$	Pseudovector	$J=1^{\star}$
$1^{3}D_{1}$	1	$\rho(1700)$	$K^{\star}(1680)$	$\omega(1650)$	$\phi(???)$	Excited-vector	J=1
$1^{3}P_{2}$	2++	$a_2(1320)$	$K_2^{\star}(1430)$	$f_2(1270)$	$f_2'(1525)$	Tensor	J=2
$1^{3}D_{2}$	2	$\rho_2(???)$	$K_2(1820)$	$\omega_2(???)$	$\phi_2(???)$	Axial-tensor	J=Z
$1^{1}D_{2}$	2-+	$\pi_2(1670)$	$K_2(1770)$	$\eta_2(1645)$	$\eta_2(1870)$	Pseudotensor	
$1^{3}D_{3}$	3	$\rho_3(1690)$	$K_3^{\star}(1780)$	$\omega_3(1670)$	$\phi_3(1850)$	J=3 - Tensor	

Chiral partners



$n^{2S+1}L_J$	J^{PC}	$I=1$ $u\overline{d}, d\overline{u}$ $\frac{d\overline{d}-u\overline{u}}{\sqrt{2}}$	$I=1/2$ $u\overline{s}, d\overline{s}$ $s\overline{d}, s\overline{u}$	$I=0$ $\approx \frac{u\overline{u}+d\overline{d}}{\sqrt{2}}$	$I=0$ $\approx s\overline{s}$	Meson names	Chiral Partners
$1^{1}S_{0}$	0-+	π	K	$\eta(547)$	$\eta'(958)$	Pseudoscalar	J=0
$1^{3}P_{0}$	0++	$a_0(1450)$	$K_0^{\star}(1430)$	$f_0(1370)$	$f_0(1500)/f_0(1710)$	Scalar	J = 0
$1^{3}S_{1}$	1	$\rho(770)$	<i>K</i> *(892)	$\omega(782)$	$\phi(1020)$	Vector	J=1
$1^{3}P_{1}$	1++	$a_1(1260)$	K_{1A}	$f_1(1285)$	$f_1'(1420)$	Axial-vector	J = 1
$1^{1}P_{1}$	1+-	$b_1(1235)$	K_{1B}	$h_1(1170)$	$h_1(1415)$	Pseudovector	$J=1^{\star}$
$1^{3}D_{1}$	1	$\rho(1700)$	$K^{\star}(1680)$	$\omega(1650)$	$\phi(???)$	Excited-vector	J=1
$1^{3}P_{2}$	2++	$a_2(1320)$	$K_2^*(1430)$	$f_2(1270)$	$f_2'(1525)$	Tensor	J=2
$1^{3}D_{2}$	2	$\rho_2(???)$	$K_2(1820)$	$\omega_2(???)$	$\phi_2(???)$	Axial-tensor	J = Z
$1^{1}D_{2}$	2-+	$\pi_2(1670)$	$K_2(1770)$	$\eta_2(1645)$	$\eta_2(1870)$	Pseudotensor	
$1^{3}D_{3}$	3	$\rho_3(1690)$	$K_3^{\star}(1780)$	$\omega_3(1670)$	$\phi_3(1850)$	J=3 - Tensor	

Tensor and (axial-)tensors



$n^{2S+1}L_J$	J^{PC}	$I=1$ $u\overline{d}, d\overline{u}$ $\frac{d\overline{d}-u\overline{u}}{\sqrt{2}}$	$I=1/2$ $u\overline{s}, d\overline{s}$ $s\overline{d}, s\overline{u}$	$I=0$ $\approx \frac{u\overline{u}+d\overline{d}}{\sqrt{2}}$	$I=0$ $\approx s\overline{s}$	Meson names	Chiral Partners
$1^{1}S_{0}$	0-+	π	K	$\eta(547)$	$\eta'(958)$	Pseudoscalar	J=0
$1^{3}P_{0}$	0++	$a_0(1450)$	$K_0^{\star}(1430)$	$f_0(1370)$	$f_0(1500)/f_0(1710)$	Scalar	J = 0
$1^{3}S_{1}$	1	$\rho(770)$	<i>K</i> *(892)	$\omega(782)$	$\phi(1020)$	Vector	J=1
$1^{3}P_{1}$	1++	$a_1(1260)$	K_{1A}	$f_1(1285)$	$f_1'(1420)$	Axial-vector	J = 1
$1^{1}P_{1}$	1+-	$b_1(1235)$	K_{1B}	$h_1(1170)$	$h_1(1415)$	Pseudovector	$J=1^*$
$1^{3}D_{1}$	1	$ \rho(1700) $	$K^{\star}(1680)$	$\omega(1650)$	$\phi(???)$	Excited-vector	J = 1
$1^{3}P_{2}$	2++	$a_2(1320)$	$K_2^{\star}(1430)$	$f_2(1270)$	$f_2'(1525)$	Tensor	J=2
$1^{3}D_{2}$	2	$\rho_2(???)$	$K_2(1820)$	$\omega_2(???)$	$\phi_2(???)$	Axial-tensor	J — Z
$1^{1}D_{2}$	2-+	$\pi_2(1670)$	$K_2(1770)$	$\eta_2(1645)$	$\eta_2(1870)$	Pseudotensor	
$1^{3}D_{3}$	3	$\rho_3(1690)$	$K_3^{\star}(1780)$	$\omega_3(1670)$	$\phi_3(1850)$	J=3 - Tensor	

PHYSICAL REVIEW D 106, 036008 (2022)

From well-known tensor mesons to yet unknown axial-tensor mesons

Shahriyar Jafarzade[®], Arthur Vereijken, Milena Piotrowska, and Francesco Giacosa^{1,2} ¹Institute of Physics, Jan Kochanowski University, ulica Uniwersytecka 7, 25-406 Kielce, Poland ²Institute for Theoretical Physics, J. W. Goethe University, Max-von-Laue-Straße 1, 60438 Frankfurt am Main, Germany



(Received 4 April 2022; accepted 25 July 2022; published 10 August 2022)

While the ground-state tensor $(J^{PC} = 2^{++})$ mesons $a_2(1320)$, $K_2^*(1430)$, $f_2(1270)$, and $f_2'(1525)$ are well known experimentally and form an almost ideal nonet of quark-antiquark states, their chiral partners, the ground-states axial-tensor ($J^{PC} = 2^{--}$) mesons are poorly settled: only the kaonic member $K_2(1820)$ of the nonet has been experimentally found, whereas the isovector state ρ_2 and two isoscalar states ω_2 and ϕ_2 are still missing. Here, we study masses, strong, and radiative decays of tensor and axial-tensor mesons within a chiral model that links them: the established tensor mesons are used to test the model and to determine its parameters, and subsequently various predictions for their chiral partners, the axial-tensor mesons, are obtained. The results are compared to current lattice OCD outcomes as well as to other theoretical approaches and show that the ground-state axial-tensor mesons are expected to be quite broad, the vector-pseudoscalar mode being the most prominent decay mode followed by the tensor-pseudoscalar one. Nonetheless, their experimental finding seems to be possible in ongoing and/or future experiments.

DOI: 10.1103/PhysRevD.106.036008

TABLE I. Chiral multiplets, their currents, and transformations up to J = 3. [* and/or $f_0(1500)$; **a mix of.] The first two columns correspond to the assignment suggested in the Quark Model review of the PDG [8], to which we refer for further details and references (see also the discussion in the text).

J^{PC} , $^{2S+1}L_J$	$\begin{cases} I = 1(\bar{u}d, \bar{d}u, \frac{\bar{d}d - \bar{u}u}{\sqrt{2}}) \\ I = 1(-\bar{u}s, \bar{s}u, ds, \bar{s}d) \\ I = 0(\frac{\bar{u}u + \bar{d}d}{\sqrt{2}}, \bar{s}s)^{\star\star} \end{cases}$	Microscopic currents	Chiral multiplet	Transformation under $SU(3)_L \times SU(3)_R \times \times U(1)$
0 ⁻⁺ , ¹ S ₀	$\begin{cases} \pi \\ K \\ \eta, \eta'(958) \end{cases}$	$P^{ij} = \frac{1}{2} \bar{q}^j \mathrm{i} \gamma^5 q^i$	$\Phi = S + iP$	
$0^{++}, {}^{3}P_{0}$	$\begin{cases} a_0(1450) \\ K_0^*(1430) \\ f_0(1370), f_0(1710)^* \end{cases}$	$S^{ij}=rac{1}{2}ar{q}^jq^i$	$\Phi = 3 + 1P$ $(\Phi^{ij} = \bar{q}_{R}^{j} q_{L}^{i})$	$\Phi \to \mathrm{e}^{-2\mathrm{i}\alpha} U_{\mathrm{L}} \Phi U_{\mathrm{R}}^{\dagger}$
1, ¹ S ₁	$\begin{cases} \rho(770) \\ K^*(892) \\ \omega(782), \phi(1020) \end{cases}$	$V_{\mu}^{ij}=rac{1}{2}ar{q}^{j}\gamma_{\mu}q^{i}$	$egin{aligned} L_{\mu} &= V_{\mu} + A_{\mu} \ (L_{\mu}^{ij} &= ar{q}_{\mathrm{L}}^{j} \gamma_{\mu} q_{\mathrm{L}}^{i}) \end{aligned}$	$L_{\mu} \to U_{\rm L} L_{\mu} U_{\rm L}^{\dagger}$
$1^{++}, {}^{3}P_{1}$	$\begin{cases} a_1(1260) \\ K_{1,A} \\ f_1(1285), f_1(1420) \end{cases}$	$A^{ij}_{\mu}=rac{1}{2}ar{q}^{j}\gamma^{5}\gamma_{\mu}q^{i}$	$R_{\mu}=V_{\mu}-A_{\mu} \ (R_{\mu}^{ij}=ar{q}_{\mathrm{R}}^{j}\gamma_{\mu}q_{\mathrm{R}}^{i})$	$R_{\mu} \to U_{\rm R} R_{\mu} U_{\rm R}^{\dagger}$
1 ⁺⁻ , ¹ P ₁	$\begin{cases} b_1(1235) \\ K_{1,B} \\ h_1(1170), h_1(1380) \end{cases}$	$P_{\mu}^{ij} = -\frac{1}{2}ar{q}^{j}\gamma^{5} \stackrel{\leftrightarrow}{D_{\mu}} q^{i}$	$\Phi_{\mu} = S_{\mu} + \mathrm{i} P_{\mu}$	Ф2iau Ф. и†
1, ³ D ₁	$\begin{cases} \rho(1700) \\ K^*(1680) \\ \omega(1650), \phi(?) \end{cases}$	$S^{ij}_{\mu} = \frac{1}{2} \bar{q}^j \mathrm{i} \overset{\leftrightarrow}{D}_{\mu} q^i$	$(\Phi_{\mu}^{ij}=ar{q}_{ m R}^{j}{ m i} \overset{ ightarrow}{D_{\mu}}q_{ m L}^{i})$	$\Phi_{\mu} \to e^{-2i\alpha} U_{\rm L} \Phi_{\mu} U_{\rm R}^{\dagger}$
2 ⁺⁺ , ³ P ₂	$\begin{cases} a_2(1320) \\ K_2^*(1430) \\ f_2(1270), f_2'(1525) \end{cases}$	$V^{ij}_{\mu\nu} = \frac{1}{2}ar{q}^j(\gamma_\mu i \stackrel{\leftrightarrow}{D}_\mu + \cdots) q^i$	$egin{aligned} L_{\mu u} &= V_{\mu u} + A_{\mu u} \ (L^{ij}_{\mu u} &= ar{q}^j_{ m L}(\gamma_\mu { m i} \stackrel{\leftrightarrow}{D_ u} + \cdots) q^i_{ m L}) \end{aligned}$	$L_{\mu\nu} \to U_{\rm L} L_{\mu\nu} U_{\rm L}^\dagger$
2, ³ D ₂	$\begin{cases} \rho_2(?) \\ K_2(1820) \\ \omega_2(?), \phi_2(?) \end{cases}$	$A^{ij}_{\mu\nu}=rac{1}{2}ar{q}^{j}(\gamma^{5}\gamma_{\mu}i\overleftrightarrow{D}_{ u}+\cdots)q^{i}$	$R_{\mu\nu} = V_{\mu\nu} - A_{\mu\nu} \ (R^{ij}_{\mu\nu} = \bar{q}^j_{ m R}(\gamma_\mu \stackrel{\leftrightarrow}{D_ u} + \cdots) q^i_{ m R})$	$R_{\mu\nu} \to U_{\rm R} R_{\mu\nu} U_{\rm R}^{\dagger}$
2 ⁻⁺ , ¹ D ₂	$\begin{cases} \pi_2(1670) \\ K_2(1770) \\ \eta_2(1645), \eta_2(1870) \end{cases}$	$P^{ij}_{\mu\nu} = -\frac{1}{2}\bar{q}^{j}(i\gamma^{5}\overset{\leftrightarrow}{D_{\mu}}\overset{\leftrightarrow}{D_{\nu}} + \cdots)q^{i}$	$\Phi_{\mu u} = S_{\mu u} + \mathrm{i} P_{\mu u}$	A Piari A rit
2^{++} , 3F_2	$\begin{cases} a_2(?) \\ K_2^*(?) \\ f_2(?), f_2'(?) \end{cases}$	$S_{\mu\nu}^{ij} = -\frac{1}{2}\bar{q}^{j}(\stackrel{\leftrightarrow}{D_{\mu}}\stackrel{\leftrightarrow}{D_{\nu}} + \cdots)q^{i}$	$(\Phi^{ij}_{\mu\nu} = \bar{q}^j_{\mathrm{R}}(\stackrel{\leftrightarrow}{D}_{\mu}\stackrel{\leftrightarrow}{D}_{\nu} + \cdots)q^i_{\mathrm{L}})$	$\Phi_{\mu\nu} \to e^{-2i\alpha} U_{\rm L} \Phi_{\mu\nu} U_{\rm R}^{\dagger}$
3, ³ D ₃	$\begin{cases} \rho_3(1690) \\ K_3^*(1780) \\ \omega_3(1670), \phi_3(1850) \end{cases}$	1	1	



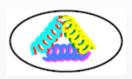
Table from:

F.G., R. Pisarski, A. Koenigstein Phys.Rev.D 97 (2018) 9, 091901 e-Print: 1709.07454

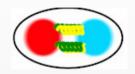
Non-conventional mesons: beyond qq



1) Glueballs



2) Hybrids



Compact diquark-antidiquark states



3) Four-quark states

Molecular states (a type of dynamical generation)



Companion poles (another type of dynamical generation)

Large-Nc: basics/1



- Instead of 3 colors, Nc colors. Then Nc is taken as a large number.
- Why to do that? Certain simplifications appear! (Yet QCD not solvable also in that limit).
- (Some) mesons become stable and slowly interacting.
- Confinement, symmetry breaking, etc...are believed to hold in large-Nc as well.

Large-Nc: basics/3



Running coupling and the 't Hooft limit

$$N_c \to \infty$$
, $g_{\rm QCD}^2 N_c \to {\rm finite}$.

$$\mu \frac{dg}{d\mu} = -bg^3$$

$$b = \frac{1}{2} \frac{1}{8\pi^2} \left(\frac{11}{3} N_c - \frac{2}{3} N_f \right)$$

$$g^{2}(\mu) = \frac{8\pi^{2}}{\left(\frac{11}{3}N_{c} - \frac{2}{3}N_{f}\right)} \frac{1}{\ln\frac{\mu}{\Lambda_{QCD}}}$$

For large- N_c we get:

$$g^{2}(\mu) = \frac{8\pi^{2}}{\left(\frac{11}{3}N_{c}\right)} \frac{1}{\ln\frac{\mu}{\Lambda_{QCD}}} \propto \frac{1}{N_{c}}$$

Large-Nc: consequences



- Constituent quark mass Nc^0
- Masses of conventional quark-antiquark states mesons and glueballs (and hybrids): Nc^0 (with one notable exception...)
- Decay width of these states decreases with Nc
- Masses of baryons proprtional to Nc; meson-baryon coupling proportional to Nc¹(1/2)

Large Nc at nonzero T



PHYSICAL REVIEW D 85, 056005 (2012)

Restoration of chiral symmetry in the large- N_c limit

Achim Heinz, ¹ Francesco Giacosa, ¹ and Dirk H. Rischke^{1,2}

¹Institute for Theoretical Physics, Goethe University, Max-von-Laue-Str. 1, D-60438 Frankfurt am Main, Germany

²Frankfurt Institute for Advanced Studies, Ruth-Moufang-Str. 1, D-60438 Frankfurt am Main, Germany

- NJL model
- Sigma model(s)
- Comparison and improvements



Is 3 a large number?

Spoiler: in some cases yes, in some cases no!

Finite T, NJL model



$$\mathcal{L}_{NJL}(N_c) = \bar{\psi}(i\gamma^{\mu}\partial_{\mu} - m_q)\psi + \frac{3G}{N_c}[(\bar{\psi}\psi)^2 + (\bar{\psi}i\gamma_5\psi)]^2$$

$$1 = \frac{m_q}{m^*} + \frac{3G}{N_c} \left(2N_c + \frac{1}{2} \right) \int_0^{\Lambda} \frac{dkk^2}{\pi^2} \frac{2 \tanh\left(\frac{\sqrt{k^2 + m^{*2}}}{2T} \right)}{\sqrt{k^2 + m^{*2}}},$$

$$T_c(N_c) \simeq \Lambda \sqrt{\frac{3}{\pi^2}} \sqrt{1 - \frac{\pi^2}{6\Lambda^2 G}} \propto N_c^0.$$

FInite T, sigma model



$$\mathcal{L}_{\sigma}(N_c) = \frac{1}{2} (\partial_{\mu} \Phi)^2 + \frac{1}{2} \mu^2 \Phi^2 - \frac{\lambda}{4} \frac{3}{N_c} \Phi^4$$

$$\Phi^t = (\sigma, \vec{\pi})$$

$$0 = \varphi(T)^2 - \frac{N_c}{3\lambda}\mu^2 + 3\int (G_{\sigma} + G_{\pi}).$$

$$\int G_i = \int_0^\infty \frac{dkk^2}{2\pi^2 \sqrt{k^2 + M_i^2}} \left[\exp\left(\frac{\sqrt{k^2 + M_i^2}}{T}\right) - 1 \right]^{-1}$$

$$T_c(N_c) = \sqrt{2} f_{\pi} \sqrt{\frac{N_c}{3}} \propto N_c^{1/2}.$$

How to cure it?



Modify the mass term:

$$\mu^2 \to \mu(T)^2 = \mu^2 \left(1 - \frac{T^2}{T_0^2}\right)$$

Use a quark-meson model (see later)

Introduce the Polyakov loop

$$l(x) = N_c^{-1} \operatorname{Tr} \left[\mathcal{P} \exp \left(i g_{\text{QCD}} \int_0^{1/T} A_0(\tau, x) d\tau \right) \right], \quad \text{For I = 0 conf, I = 1 deconf.}$$

$$\mathcal{L}_{\sigma\text{-Pol}}(N_c) = \mathcal{L}_{\sigma}(N_c) + \frac{\alpha N_c}{4\pi} |\partial_{\mu} l|^2 T^2 - \mathcal{V}(l) - \frac{h^2}{2} \Phi^2 |l|^2 T^2. \qquad T_c = \frac{\mu}{\sqrt{h^2 |l(T_c)|^2 + \frac{6\lambda}{N}}}$$

$$T_c = \frac{\mu}{\sqrt{h^2|l(T_c)|^2 + \frac{6\lambda}{N_c}}}$$



PHYSICAL REVIEW D **106**, 116016 (2022)

Editors' Suggestion

Fate of the critical endpoint at large N_c

Péter Kovács® and Győző Kovács®

Institute for Particle and Nuclear Physics, Wigner Research Centre for Physics, 1121 Budapest, Hungary and Institute of Physics, Eötvös University, 1117 Budapest, Hungary

Francesco Giacosa

Institute of Physics, Jan Kochanowski University, ulica Uniwersytecka 7, P-25-406 Kielce, Poland and Institute for Theoretical Physics, Goethe-University, Max-von-Laue-Straße 1, D-60438 Frankfurt am Main, Germany



(Received 22 September 2022; accepted 18 November 2022; published 23 December 2022)

- Using a model a sigma-model that is as complete as possible (with (psuedo)scalar, (axial-)vector) d.o.f.)
- Linear realization of chiral symmetry
- Vacuum: D. Parganlija et al., Phys.Rev.D 87 (2013) 1, 014011 e-Print: 1208.0585 [hep-ph]
- Extension to the medium: P. Kovacs, Phys.Rev.D 93 (2016) 11, 114014 • e-Print: 1601.05291 [hep-ph]: coupling to quarks and to the Polyakov loop.

eLSM Lagrangian, etc. Actually just a complicated vs of the Mexican hat ©



$$\begin{split} \mathcal{L}_{m} &= \mathrm{Tr}[(D_{\mu}M)^{\dagger}(D^{\mu}M)] - m_{0}\mathrm{Tr}(M^{\dagger}M) - \lambda_{1}[\mathrm{Tr}(M^{\dagger}M)]^{2} - \lambda_{2}[\mathrm{Tr}(M^{\dagger}M)^{2}] + c(\det M + \det M^{\dagger}) + \mathrm{Tr}[H(M + M^{\dagger})] \\ &- \frac{1}{4}\mathrm{Tr}[L_{\mu\nu}L^{\mu\nu} + R_{\mu\nu}R^{\mu\nu}] + \mathrm{Tr}\left[\left(\frac{m_{1}^{2}}{2} + \Delta\right)(L_{\mu}L^{\mu} + R_{\mu}R^{\mu})\right] + \frac{h_{1}}{2}\mathrm{Tr}(\phi^{\dagger}\phi)\mathrm{Tr}[L_{\mu}L^{\mu} + R_{\mu}R^{\mu}] \\ &+ h_{2}\mathrm{Tr}[(MR_{\mu})^{\dagger}(MR^{\mu}) + (L_{\mu}M)^{\dagger}(L^{\mu}M)] + 2h_{3}\mathrm{Tr}[R_{\mu}M^{\dagger}L^{\mu}M] - 2g_{2}\mathrm{Tr}\{L_{\mu\nu}[L^{\mu},L^{\nu}]\} + \mathrm{Tr}\{R_{\mu\nu}[R^{\mu},R^{\nu}]\}, \end{split}$$

$$\mathcal{L}_Y = \bar{\psi}(i\gamma_\mu \partial^\mu - g_F(S + i\gamma_5 P))\psi.$$

$$\begin{split} M &= S + iP = \sum_a (S_a + iP_a) T_a, \\ L^\mu &= V^\mu + A^\mu = \sum_a (V_a^\mu + A_a^\mu) T_a, \\ R^\mu &= V^\mu - A^\mu = \sum_a (V_a^\mu - A_a^\mu) T_a, \end{split}$$

$$\begin{split} D^{\mu} &= \partial^{\mu} M - i g_1(L_{\mu} M - M R_{\mu}) - i e A^{\mu} [T_3, M], \\ L^{\mu\nu} &= \partial^{\mu} L^{\nu} - i e A^{\mu} [T_3, L^{\nu}] - \{ \partial^{\nu} L^{\mu} - i e A^{\nu} [T_3, L^{\mu}] \}, \\ R^{\mu\nu} &= \partial^{\mu} R^{\nu} - i e A^{\mu} [T_3, R^{\nu}] - \{ \partial^{\nu} R^{\mu} - i e A^{\nu} [T_3, R^{\mu}] \}, \end{split}$$

 $\Omega(T, \mu_a) = U(\langle M \rangle) + \Omega_{\bar{a}a}^{(0)}(T, \mu_a) + U(\langle \Phi \rangle, \langle \bar{\Phi} \rangle)$

• Polyakov loop potential.
$$\Omega(T, \mu_q) = U_{Cl} + \Omega_{\bar{q}q}(T, \mu_q) + U_{Pol}(T, \mu_q)$$
 (2)

$$\Omega_{\bar{q}q}^{\mathbf{v}} = -2N_c \sum_{f=u,d,s} \int \frac{d^3p}{(2\pi)^3} E_f(p),$$

$$\Omega_{\bar{q}q}^{\mathbf{T}}(T,\mu_q) = -2T \sum_{f=u,d,s} \int \frac{d^3p}{(2\pi)^3} \mathrm{Tr}_c \left[\ln \left(1 + L^{\dagger} e^{-\beta (E_f(p) - \mu_q)} \right) + \ln \left(1 + L e^{-\beta (E_f(p) + \mu_q)} \right) \right]$$

Parameters and large-Nc scaling



TABLE I. Parameter sets. Left column is taken from [11] (set A) and right column is taken from [38] (set B).

Parameter	Set A	Set B
ϕ_N [GeV]	0.1411	0.1290
ϕ_S [GeV]	0.1416	0.1406
m_0^2 [GeV ²]	2.3925_{E-4}	-1.2370_{E-2}
m_1^2 [GeV ²]	6.3298_{E-8}	0.5600
λ_1	-1.6738	-1.0096
λ_2	23.5078	25.7328
c_1 [GeV]	1.3086	1.4700
δ_S [GeV ²]	0.1133	0.2305
g_1	5.6156	5.3295
g_2	3.0467	-1.0579
h_1	37.4617	5.8467
h_2	4.2281	-12.3456
h_3	2.9839	3.5755
g_F	4.5708	4.9571
M_0 [GeV]	0.3511	0.3935

TABLE II.	N_c dependence of the parameters.				
	m_0^2, m_1^2, δ_S	N_c^0			
	g_1, g_2, g_f	$1/\sqrt{N_c}$			
	λ_2, h_2, h_3	N_c^{-1}			
	λ_1, h_1	N_c^{-2}			
	c_1	$N_c^{-3/2}$			
	$h_{N/S}$	$\sqrt{N_c}$			

 $1/\sqrt{N_c}$

Chiral condensate vs T



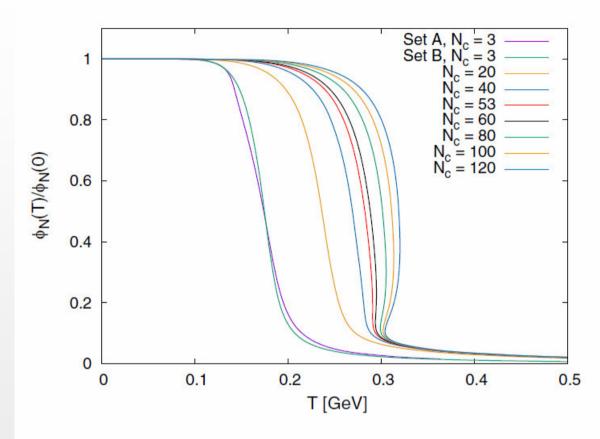


FIG. 4. The temperature dependence of the normalized chiral condensate ϕ_N .

Chiral condensate vs chemical potential



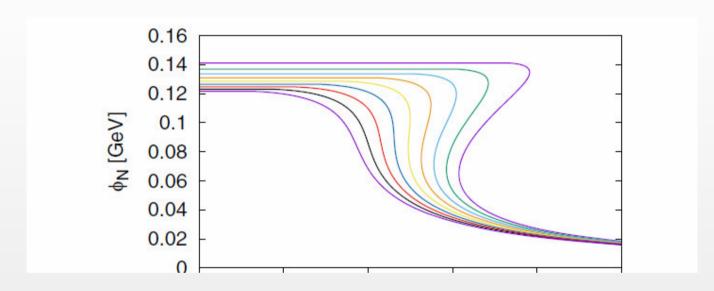
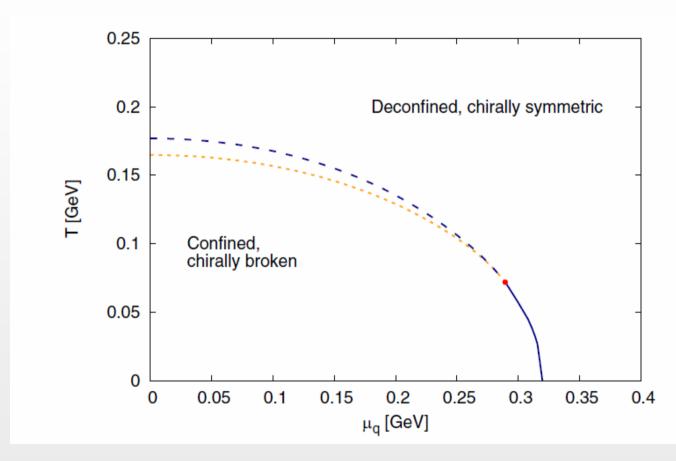


FIG. 3. The μ_q quark chemical potential dependence of the ϕ_N condensate at different N_c values. $N_c = 3.00$ corresponds to the rightmost curve, while $N_c = 3.45$ corresponds to the leftmost curve. The top figure is obtained with set A, while the bottom figure with set B of Table I.

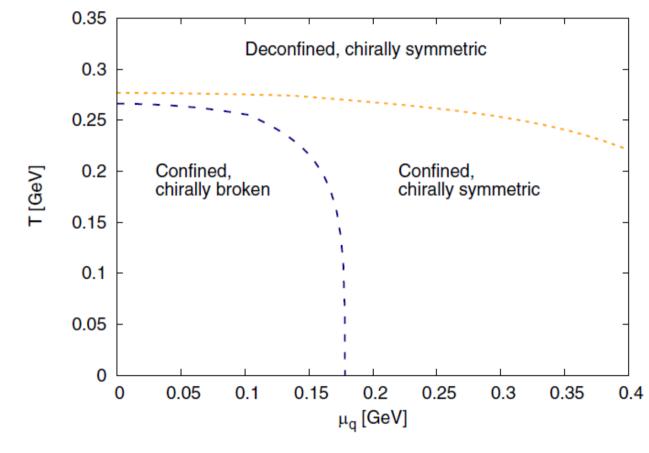
Phase diagram: Nc =3





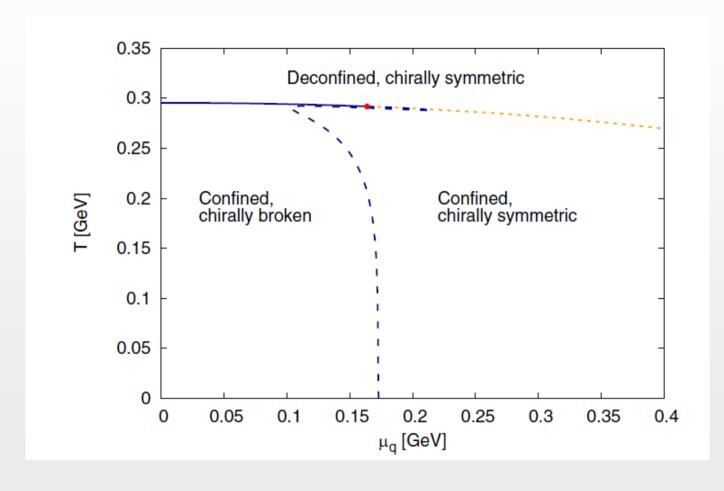
Phase diagram: Nc = 33 (only cross-over, no CP)

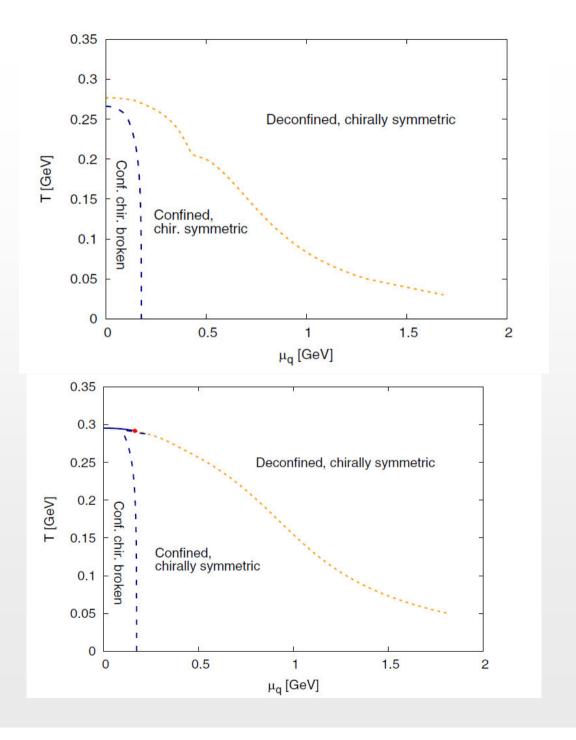




Phase diagram: Nc = 63









$$Nc = 33$$

$$Nc = 63$$

Schematic phase diagram at large Nc



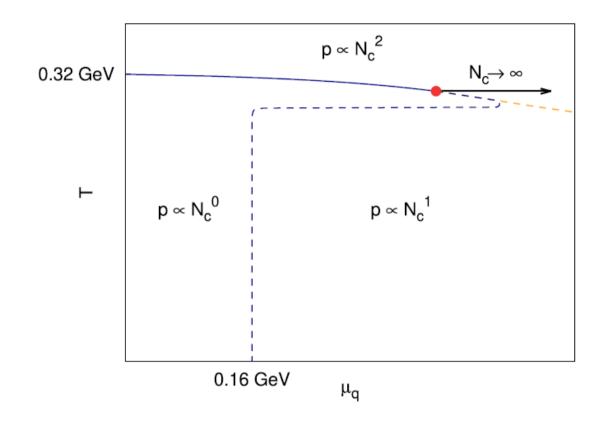
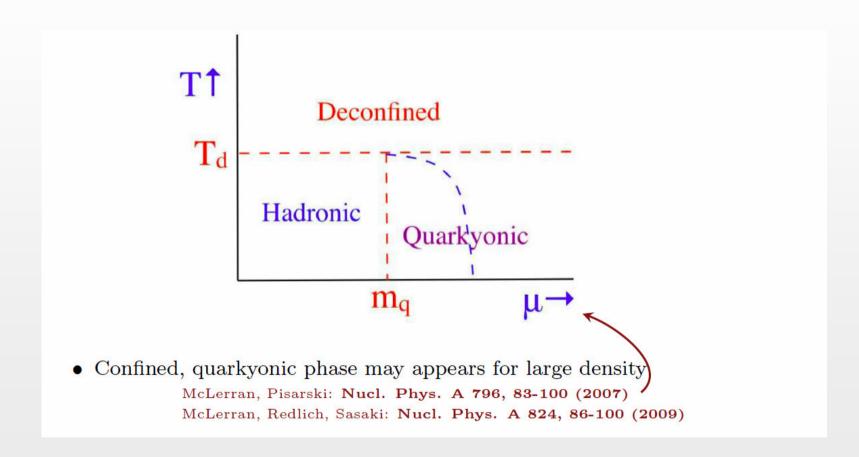


FIG. 13. The schematic phase diagram for large N_c and the N_c scaling of the pressure in the different phases.

Then, for the QCD diagram: 3 is not a large number!!!!

...agrees well with quarkyonic...













Nuclear Physics A 859 (2011) 49-62

www.elsevier.com/locate/nuclphysa

Does nuclear matter bind at large N_c ?

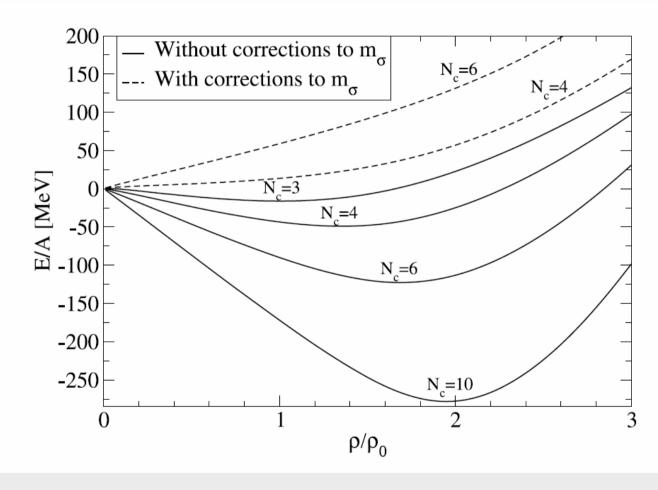
Luca Bonanno*, Francesco Giacosa

The Lagrangian of the Walecka model reads [9]:

$$\mathcal{L} = \bar{\psi} \left[\gamma^{\mu} (i \partial_{\mu} - g_{\omega} \omega_{\mu}) - (m_N - g_{\sigma} \sigma) \right] \psi + \frac{1}{2} \partial^{\mu} \sigma \partial_{\mu} \sigma - \frac{1}{2} m_{\sigma}^2 \sigma^2 - \frac{1}{4} F^{\mu \nu} F_{\mu \nu} + \frac{1}{2} m_{\omega}^2 \omega^{\mu} \omega_{\mu} - V_{\sigma}(\sigma),$$

$$m_{\sigma} \longrightarrow m_{\sigma};$$
 $m_{\omega} \longrightarrow m_{\omega}, \qquad m_{N} \longrightarrow m_{N} \frac{N_{c}}{3};$
 $g_{\sigma} \longrightarrow g_{\sigma} \sqrt{\frac{N_{c}}{3}}, \qquad g_{\omega} \longrightarrow g_{\omega} \sqrt{\frac{N_{c}}{3}}.$





$$m_{\sigma}^{2}(N_{c}) = m_{\sigma}^{2} + b_{\sigma}^{2} \left(\frac{1}{3} - \frac{1}{N_{c}}\right).$$

Minimal variation of the scaling... quark model places this state higher. **Enough to unbind nuclear matter**

What if the lightest scalar is a tetraquark?



$$\mathcal{L} = \bar{\psi} \left[\gamma^{\mu} (i \partial_{\mu} - g_{\omega} \omega_{\mu}) - (m_N - g_{\chi} \chi) \right] \psi + \frac{1}{2} \partial^{\mu} \chi \partial_{\mu} \chi - \frac{1}{2} m_{\chi}^2 \chi^2$$
$$- \frac{1}{4} F^{\mu \nu} F_{\mu \nu} + \frac{1}{2} m_{\omega}^2 \omega^{\mu} \omega_{\mu},$$

$$\chi = [\bar{R}, \bar{B}][R, B] + [\bar{G}, \bar{B}][G, B] + [\bar{R}, \bar{G}][R, G]$$
 for Nc = 3

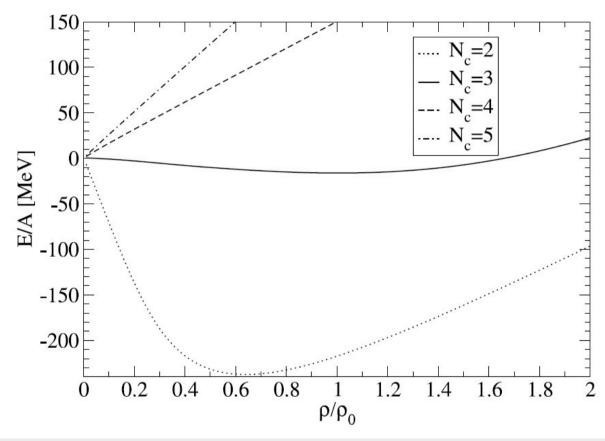
$$d_{a_1} = \varepsilon_{a_1 a_2 a_3 \cdots a_{N_c}} q^{a_2} q^{a_3} \cdots q^{a_{N_c}}$$
 with $a_2, \dots, a_{N_c} = 1, \dots, N_c$. for Nc> 3

$$\chi = \sum_{a_1=1}^{N_c} d_{a_1}^{\dagger} d_{a_1}$$

Extended 'tetraquark' version! (indeed, a well-defined one)

$$m_{\chi} \to m_{\chi} \frac{2N_c - 2}{4}, \qquad g_{\chi} \to g_{\chi}.$$





Summary: for nuclear matter, 3 is not a large number!!!!

Other scenarios



- Two scalar fields: tetraquark+quarkonium, no nuclear matter.
- f0(500) as pion-pion molecular states, dissolves at large Nc, no nuclear matter.
- One-pion-exchange: what does eventually happen at very large Nc? (not taken into account here because beyond MFE)





Available online at www.sciencedirect.com

ScienceDirect

Nuclear Physics A 968 (2017) 366–378



www.elsevier.com/locate/nuclphysa



Neutron stars in the large- N_c limit

Francesco Giacosa a,b, Giuseppe Pagliara c,*

$$p_q = b_1 N_c \mu_q^4 - N_c^2 B$$

$$b_1 = \frac{N_f}{12\pi^2}.$$

 $b_1 = \frac{N_f}{12\pi^2}$. Quark matter athigh density: free gas plus bag

$$p_b = a_1 \mu_b^{\alpha} - K$$

$$a_1(N_c) \propto \left(\frac{g_V^2}{m_V^2}\right)^{\frac{\alpha-4}{2}} \propto N_c^{\frac{\alpha-4}{2}}$$

$$K = \tilde{K} N_c^{(3\alpha - 4)/2}$$

Baryonic matter at high density (starting from 2p0): parameter α unknown

$$v_b = \sqrt{\frac{dp_b}{d\varepsilon_b}} = \frac{1}{\sqrt{\alpha - 1}}.$$
 $\alpha \ge 2$

$$\alpha \ge 2$$



The stiffest equation of state corresponds, in agreement with causality, to $\alpha = 2$.

$$p_b = \tilde{a}_1 N_c \mu_q^2 - \tilde{K} N_c$$

The speed of sound is 1 in this case

Stiffest equation and transition



The stiffest equation of state corresponds, in agreement with causality, to $\alpha = 2$.

$$p_b = \tilde{a}_1 N_c \mu_q^2 - \tilde{K} N_c$$

The speed of sound is 1 in this case

$$p_b = \tilde{a}_1 N_c^{\frac{3\alpha - 4}{2}} \mu_q^{\alpha} - \tilde{K} N_c^{\frac{3\alpha - 4}{2}} = b_1 N_c \mu_q^4 - N_c^2 B = p_q$$

$$\mu_q^{\text{crit}} = \left(\frac{BN_c}{b_1}\right)^{1/4} [1 + \dots] \text{ for } 2 \le \alpha \le \frac{16}{7}$$

Results



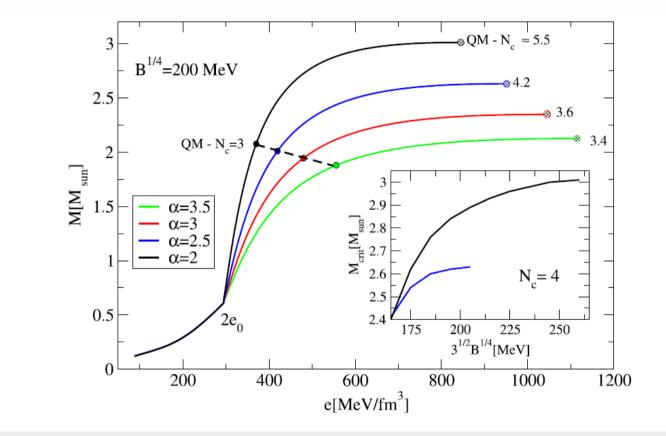


Fig. 1. Masses of neutron stars as functions of their central energy densities for different values of α . The filled points correspond to the critical mass for the transition to quark matter in the case $N_c = 3$. The filled dashed points correspond to the critical values of N_c for which the phase transition occurs in correspondence of the maximum mass configuration.

Neutron stars, main outcome



From the $N_c=3$ analysis, one further reduces the range (20), namely α must be *smaller* than about $\alpha_{\rm max}\simeq 2.5$ (the blue line) in order to explain the existence of $2M_{\odot}$ stars:

$$2 \le \alpha \lesssim 2.5$$
 for $N_c = 3$.

Moreover one should not observe stars with masses larger than about $2.1 M_{\odot}$

Also for neutron stars: Nc = 3 is not large!

Conclusions and outlook



Large-Nc

- useful tool for QCD (as well as for a variety of models/theories)
- Phenomenology in the vacuum can be better understood (i.e. OZI), certains terms appear as dominant, other are suppressed... 3 is a large number.
- Applications at nonzero temperature and density, in most cases 3 is not a large number.



Thanks!

Large-Nc: basics/1



invariant under SU(3) = white

Graniant under 50 (3)

IBARYON > -> IBARYON

Involuent under SUECNE

Large-Nc: basics/2



	DIAGRAM:																						3				
	9:1,2	3 .									a:																
						:																				*	
,	G-400N:	٠ . ک					F	A.	111		1		· : 1		5			•	1	?	1	2	N	2.			
																							.(-	
. 400000	g	3	-	A.F.	4			22	ee.	2	, i				\	0 100		_	Hu			, constraint		-			100
9 com	bination (-1 sle airo resta)		-)			Nc	? (.	-1)) ((in	n b	in	çi (. 2	7	i A est	F	(a)	b)	-	GI.	2,5 FÖRDER	3 4,	, No	0

Glueball production and decays: gluon-rich processes



Glueballs should be find in gluon-rich processes (such as J/ψ decays, proton-antiproton fusion, ...)

Glueball should have suppressed decay into flavor breaking channels (eg eta-etaprime)

Moreover, glueballs should have a suppressed (but nonzero!) decay into photons.



In the $N_c \to \infty$ limit eg.:

- Stable, noninteracting mesons and glue-balls (infinite number with fixed qn.) in the hadronic phase with $m \propto N_c^0$ masses.
- Baryon masses diverges as $m_B \propto N_c^1$.
- Hadronic phase built from noninteracting mesons and glueballs, energy density scales as $\propto N_c^0$
- Phase boundary to quark-gluon plasma at a temperature $\propto N_c^0$
- Energy density of quark-gluon phase N_c^2 . \Rightarrow First or second order phase transition expected.
- Quark loops are suppressed: the thermodynamics expected to became similar to Yang-Mills.
- Confined, quarkyonic phase may appears for large density

McLerran, Pisarski: Nucl. Phys. A 796, 83-100 (2007) McLerran, Redlich, Sasaki: Nucl. Phys. A 824, 86-100 (2009)





Available online at www.sciencedirect.com

ScienceDirect



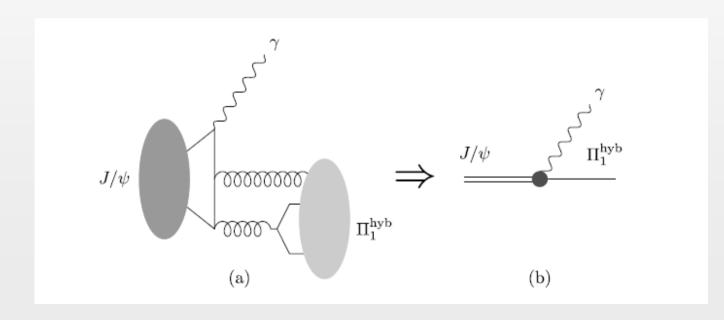






Radiative production and decays of the exotic $\eta'_1(1855)$ and its siblings

Vanamali Shastry a,*, Francesco Giacosa a,b



Glueball spectrum from lattice QCD



