# Relativistic Quantum Mechanics of the Majorana Particle 

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## Introduction

1. The Dirac equation

$$
i \gamma^{\mu} \partial_{\mu} \psi(x)-m \psi(x)=0, \quad \psi(x)=\left(\begin{array}{c}
\psi^{1} \\
\psi^{2} \\
\psi^{3} \\
\psi^{4}
\end{array}\right), \quad \psi^{\alpha}(x) \in \mathbb{C}
$$

Scalar product: $\left\langle\psi_{1} \mid \psi_{2}\right\rangle=\int d^{3} x \psi_{1}(\mathbf{x}, t)^{\dagger} \psi_{2}(\mathbf{x}, t)$ with arbitrary $t$. The Hamiltonian form: $i \partial_{t} \psi=\hat{H} \psi$, Hermitian $\hat{H}=-i \gamma^{0} \gamma^{k} \partial_{k}+m \gamma^{0}$.
General solution of the equation

$$
\begin{equation*}
\psi(\mathbf{x}, t)=\frac{1}{(2 \pi)^{3 / 2}} \int d^{3} p e^{i \mathbf{p} \mathbf{x}}\left(e^{-i E_{\rho} t} v^{(+)}(\mathbf{p})+e^{i E_{\rho} t} v^{(-)}(\mathbf{p})\right) \tag{1}
\end{equation*}
$$

where

$$
\left(\gamma^{0} \gamma^{\prime} p^{\prime}+m \gamma^{0}\right) v^{( \pm)}(\mathbf{p})= \pm E_{p} v^{( \pm)}(\mathbf{p}), \quad E_{p}=+\sqrt{m^{2}+\mathbf{p}^{2}}
$$

(1) is important as the starting point for QFT of the Dirac field.

Two views on (1): expansion in the basis of common eigenvectors of commuting observables $\hat{\mathbf{p}}=-i \nabla$ and $\hat{H}$ (physics); or merely the Fourier transformation (mathematics).

## Introduction

2. There exist Majorana representations for $\gamma^{\mu}$ matrices in which they are purely imaginary. For example,

$$
\begin{aligned}
& \gamma^{0}=\left(\begin{array}{cc}
0 & \sigma_{2} \\
\sigma_{2} & 0
\end{array}\right), \gamma^{1}=i\left(\begin{array}{cc}
-\sigma_{0} & 0 \\
0 & \sigma_{0}
\end{array}\right), \gamma^{2}=i\left(\begin{array}{cc}
0 & \sigma_{1} \\
\sigma_{1} & 0
\end{array}\right), \\
& \gamma^{3}=-i\left(\begin{array}{cc}
0 & \sigma_{3} \\
\sigma_{3} & 0
\end{array}\right), \quad \text { and } \gamma_{5}=i \gamma^{0} \gamma^{1} \gamma^{2} \gamma^{3}=i\left(\begin{array}{cc}
0 & \sigma_{0} \\
-\sigma_{0} & 0
\end{array}\right) .
\end{aligned}
$$

$\sigma_{k}-$ the Pauli matrices, $\sigma_{0}-$ the $2 \times 2$ unit matrix. In the Majorana representations charge conjugation $C$ is represented simply by the complex conjugation.
By definition, the Majorana bispinors are invariant under $C$. Thus, they have real components in the Majorana repr. The Hilbert space of the Majorana bispinors is over $\mathbb{R}$, not $\mathbb{C}$. The scalar product has the form $\left\langle\psi_{1} \mid \psi_{2}\right\rangle=\int d^{3} x \psi_{1}(\mathbf{x}, t)^{T} \psi_{2}(\mathbf{x}, t)$. QM without complex numbers.
The Dirac bispinor is a composite object:

- $\psi=\psi_{1}+i \psi_{2}$ with real $\psi_{1,2}, \quad \boldsymbol{C} \psi_{1}=\psi_{1}, \quad \boldsymbol{C}\left(i \psi_{2}\right)=-i \psi_{2}$;
- the decomposition $\psi=\psi_{R}+\psi_{L}$ into the Weyl bispinors;
- two irreducible unitary $s=1 / 2$ reprs of the Poincaré group.


## Axial momentum

Questions: • What is the 'eigenvector' version of expansion (1) for the Majorana bispinors?

- $\hat{\mathbf{p}}=-i \nabla \rightarrow$ ?

1. There exists $1: 1$ mapping $M$ between linear spaces of the Majorana and right-handed (or left-handed) Weyl bispinors. Take arbitrary right-handed Weyl bispinor $\phi, \gamma_{5} \phi=\phi$, and form $\psi=\phi+\phi^{*} \equiv M(\phi) . \quad \psi$ is a Majorana bispinor. $M$ is invertible: $\phi=\left(I+\gamma_{5}\right) \psi / 2 \equiv \psi_{R}$. $M$ preserves linear combinations only if their coefficients are real.
The Weyl bispinors are complex, hence the standard momentum operator $\hat{\mathbf{p}}=-i \nabla$ is well-defined for them. It commutes with $\gamma_{5}$, therefore also $\hat{\mathbf{p}} \phi$ is right-handed Weyl bispinor. Let us find the Majorana bispinor that corresponds to $\hat{\mathbf{p}} \phi$ :

$$
M(\hat{\mathbf{p}} \phi)=\hat{\mathbf{p}} \phi+(\hat{\mathbf{p}} \phi)^{*}=\hat{\mathbf{p}}\left(\psi_{R}-\psi_{L}\right)=-i \gamma_{5} \nabla \psi \equiv \hat{\mathbf{p}}_{5} \psi .
$$

Thus the standard momentum operator in the space of right-handed Weyl bispinors gives rise to $\hat{\mathbf{p}}_{5}=-i \gamma_{5} \nabla$ - the axial momentum operator - in the space of Majorana bispinors.

## Axial momentum

$\hat{\mathbf{p}}_{5}$ commutes with $\gamma_{5}$, therefore it can be used also in the space of right-handed Weyl bispinors. It is invariant under the mapping $M$ : if $\psi=M(\phi)$ then $\hat{\mathbf{p}}_{5} \psi=M\left(\hat{\mathbf{p}}_{5} \phi\right)$ because the matrix $-i \gamma_{5}$ is real.
2. Are the two quantum mechanics, Majorana and Weyl, equivalent? No, because the mapping $M$ does not preserve scalar product. Take $\psi_{1}=M\left(\phi_{1}\right), \psi_{2}=M\left(\phi_{2}\right)$,

$$
\int d^{3} x \psi_{1}^{T} \psi_{2}=\int d^{3} x\left(\phi_{1}^{\dagger} \phi_{2}+\left(\phi_{1}^{\dagger} \phi_{2}\right)^{*}\right)
$$

Moreover, there are differences in evolution equations. In the Weyl case, evolution equation has the form (1) with $m=0$; in the Majorana case $m \neq 0$ is allowed. Using $M^{-1}$ one can transform Eq. (1) for $\psi$ to the space of right-handed Weyl bispinors:

$$
i \gamma^{\mu} \partial_{\mu} \phi-m \phi^{*}=0 .
$$

This equation is known as the Majorana equation for $\phi$ ( $\phi^{*}$ is charge conjugation of $\phi$ ). It can not be accepted as quantum evolution equation for the Weyl $\phi$ because it is not linear over $\mathbb{C}$. The Hilbert space of the Weyl bispinors is linear over $\mathbb{C}$ - it includes all bispinors such that $\gamma_{5} \phi=\phi$.

## Axial momentum

3. The normalized eigenvectors of the axial momentum obey the conditions

$$
\hat{\mathbf{p}}_{5} \psi_{\mathbf{p}}(\mathbf{x})=\mathbf{p} \psi_{\mathbf{p}}(\mathbf{x}), \quad \int d^{3} x \psi_{\mathbf{p}}^{T}(\mathbf{x}) \psi_{\mathbf{q}}(\mathbf{x})=\delta(\mathbf{p}-\mathbf{q})
$$

They have the form

$$
\psi_{\mathbf{p}}(\mathbf{x})=(2 \pi)^{-3 / 2} \exp \left(i \gamma_{5} \mathbf{p} \mathbf{x}\right) v
$$

where $v$ an arbitrary real, constant, normalized $\left(v^{\top} v=1\right)$ bispinor, and

$$
\exp \left(i \gamma_{5} \mathbf{p x}\right)=\cos (\mathbf{p x}) I+i \gamma_{5} \sin (\mathbf{p x})
$$

We call them the axial plane waves.
4. Commutator of the axial momentum with position operator $\hat{\mathbf{x}}$ has the form

$$
\left[\hat{x}^{j}, \hat{p}_{5}^{k}\right]=i \delta_{j k} \gamma_{5} .
$$

The implied uncertainty relation is the well-known one

$$
\langle\psi|\left(\Delta \hat{x}^{j}\right)^{2}|\psi\rangle\langle\psi|\left(\Delta \hat{p}_{5}^{k}\right)^{2}|\psi\rangle \geq \frac{1}{4} \delta_{j k},
$$

where $\Delta \hat{x}^{j}=\hat{x}^{j}-\langle\psi| \hat{x}^{j}|\psi\rangle, \Delta \hat{p}_{5}^{k}=\hat{p}_{5}^{k}-\langle\psi| \hat{p}_{5}^{k}|\psi\rangle$.

## Axial momentum

5. The axial momentum commutes with the Hamiltonian $\hat{h}=-\gamma^{0} \gamma^{k} \partial_{k}-i m \gamma^{0}$ only in the massless case:

$$
\left[\hat{h}, \hat{\mathbf{p}}_{5}\right]=-2 i m \gamma^{0} \hat{\mathbf{p}}_{5} .
$$

The operator $\nabla$ commutes with $\hat{h}$. Therefore, in the Heisenberg picture, $\hat{\mathbf{p}}_{5}(t)=-i \hat{\gamma}_{5}(t) \nabla$. It turns out that the operator $\hat{\gamma_{5}}(t)$ has the following form

$$
\hat{\gamma}_{5}(t)=\gamma_{5}+i \frac{m}{\hat{\omega}} \gamma^{0} \gamma_{5}[\sin (2 \hat{\omega} t)+\hat{J}(1-\cos (2 \hat{\omega} t))] .
$$

Here $\hat{\jmath}=\hat{h} / \hat{\omega}, \hat{\omega}=\sqrt{m^{2}-\Delta} \rightarrow E_{p}=\sqrt{m^{2}+\vec{p}^{2}}$. Note that $\hat{\jmath}^{2}=-I$.
Oscillating matrix elements of the axial momentum:

$$
\begin{aligned}
& \int d^{3} x \psi_{p}^{T}(\vec{x}) \hat{\mathbf{p}}_{5}(t) \psi_{q}(\vec{x})=\mathbf{p}\left[1+\frac{m^{2}}{E_{p}^{2}}\left(\cos \left(2 E_{p} t\right)-1\right)\right]\left(v^{T} w\right) \delta(\vec{p}-\vec{q}) \\
& -\mathbf{p} \frac{m}{E_{p}}\left[i \sin \left(2 E_{p} t\right)\left(v^{T} \gamma^{0} w\right)+\left(1-\cos \left(2 E_{p} t\right)\right)\left(v^{T} \gamma_{5} \frac{\gamma^{j} p^{j}}{E_{p}} w\right)\right] \delta(\vec{p}+\vec{q})
\end{aligned}
$$

## General solution of the evolution equation

1. The expansion of $\psi(\mathbf{x}, t)$ into the axial plane waves
$\psi(\mathbf{x}, t)=\frac{1}{(2 \pi)^{3 / 2}} \sum_{\alpha=1}^{2} \int d^{3} p e^{i \gamma_{5} \mathbf{p x}}\left(v_{\alpha}^{(+)}(\mathbf{p}) c_{\alpha}(\mathbf{p}, t)+v_{\alpha}^{(-)}(\mathbf{p}) d_{\alpha}(\mathbf{p}, t)\right)$.
As the orthonormal basis of the real bispinors we take the eigenvectors of $\gamma^{0} \gamma^{k} p^{k}$ (this matrix commutes with $\mathbf{p}_{5}$ ):

$$
\begin{aligned}
v_{1}^{(+)}(\mathbf{p}) & =\frac{1}{\sqrt{2|\mathbf{p}|\left(|\mathbf{p}|-p^{2}\right)}}\left(\begin{array}{c}
-p^{3} \\
p^{2}-|\mathbf{p}| \\
p^{1} \\
0
\end{array}\right), \quad v_{2}^{(+)}(\mathbf{p})=i \gamma_{5} v_{1}^{(+)}(\mathbf{p}), \\
v_{1}^{(-)}(\mathbf{p}) & =i \gamma^{0} v_{1}^{(+)}(\mathbf{p}), \quad v_{2}^{(-)}(\mathbf{p})=i \gamma_{5} v_{1}^{(-)}(\mathbf{p})=-\gamma_{5} \gamma^{0} v_{1}^{(+)}(\mathbf{p}) .
\end{aligned}
$$

This is rather interesting basis: scale invariant, and real. Can be used also in the Dirac quantum mechanics.
Another expansion (without $\hat{\mathbf{p}}_{5}$ ): L. Pedro, arXiv:1212.5465 (2012).

## General solution of the evolution equation

2. Time dependence of the real amplitudes $c_{\alpha}(\mathbf{p}, t), d_{\alpha}(\mathbf{p}, t)$ is found from the Dirac equation. The amplitudes are split into the even and odd parts,

$$
c_{\alpha}(\mathbf{p}, t)=c_{\alpha}^{\prime}(\mathbf{p}, t)+c_{\alpha}^{\prime \prime}(\mathbf{p}, t), \quad d_{\alpha}(\mathbf{p}, t)=d_{\alpha}^{\prime}(\mathbf{p}, t)+d_{\alpha}^{\prime \prime}(\mathbf{p}, t),
$$

where $c_{\alpha}^{\prime}(-\mathbf{p}, t)=c_{\alpha}^{\prime}(\mathbf{p}, t), \quad c_{\alpha}^{\prime \prime}(-\mathbf{p}, t)=-c_{\alpha}^{\prime \prime}(\mathbf{p}, t)$, similarly for $d^{\prime}, d^{\prime \prime}$.
Furthermore, we introduce
$\vec{c}(\mathbf{p}, t)=\left(\begin{array}{c}c_{1}^{\prime} \\ c_{1}^{\prime \prime} \\ c_{2}^{\prime} \\ c_{2}^{\prime \prime}\end{array}\right), \vec{d}(\mathbf{p}, t)=\left(\begin{array}{c}d_{1}^{\prime} \\ d_{1}^{\prime \prime} \\ d_{2}^{\prime \prime} \\ d_{2}^{\prime \prime}\end{array}\right), K_{ \pm}(\mathbf{p})=\left(\begin{array}{cccc}0 & -n^{1} & \pm n^{2} & \pm n^{3} \\ n^{1} & 0 & \mp n^{3} & \pm n^{2} \\ \mp n^{2} & \pm n^{3} & 0 & n^{1} \\ \mp n^{3} & \mp n^{2} & -n^{1} & 0\end{array}\right)$,
where $E_{p}=\sqrt{\mathbf{p}^{2}+m^{2}}$, and

$$
n^{1}=\frac{m p^{1}}{E_{p} \sqrt{\left(p^{1}\right)^{2}+\left(p^{3}\right)^{2}}}, \quad n^{2}=\frac{|\mathbf{p}|}{E_{p}}, \quad n^{3}=\frac{m p^{3}}{E_{p} \sqrt{\left(p^{1}\right)^{2}+\left(p^{3}\right)^{2}}} .
$$

## General solution of the evolution equation

The time dependence of the amplitudes is given by

$$
\vec{c}(\mathbf{p}, t)=\exp \left(t E_{p} K_{+}(\mathbf{p})\right) \vec{c}(\mathbf{p}, 0), \quad \vec{d}(\mathbf{p}, t)=\exp \left(t E_{p} K_{-}(\mathbf{p})\right) \vec{d}(\mathbf{p}, 0)
$$

The matrices $K_{ \pm}$are antisymmetric, the matrices $\exp \left(t E_{p} K_{ \pm}(\mathbf{p})\right)$ belong to the $S O(4)$ group.
3. This solution can be rewritten in terms of quaternions. The quaternionic units $\hat{i}, \hat{j}, \hat{k}$ are introduced as follows:

$$
\hat{i}=i \gamma_{5}, \quad \hat{j}=i \gamma^{0}, \quad \hat{k}=-\gamma_{5} \gamma^{0}
$$

They obey the usual conditions

$$
\hat{i}^{2}=\hat{j}^{2}=\hat{k}^{2}=-I, \quad \hat{i} j=\hat{k}, \quad \hat{k} \hat{i}=\hat{j}, \quad \hat{j} \hat{k}=\hat{i} .
$$

The bispinor basis $v_{\alpha}^{( \pm)}(\mathbf{p})$ is generated from $v_{1}^{(+)}(\mathbf{p})$ by acting with $\hat{i}, \hat{j}, \hat{k}$. Moreover, $K_{ \pm}(\mathbf{p})=-n^{1} \hat{k} \mp n^{2} \hat{i} \pm n^{3} \hat{j}$. Therefore, the time evolution of the amplitudes $\vec{c}, \vec{d}$ at each fixed value of the axial momentum $\mathbf{p}$ is given by a time dependent quaternion.

## General solution of the evolution equation

4. The general solution can be rewritten in the form of superposition of traveling plane waves:

$$
\begin{aligned}
\psi(\mathbf{x}, t)= & \frac{1}{2(2 \pi)^{3 / 2}} \int d^{3} p\left[\cos \left(\mathbf{p} \mathbf{x}-E_{p} t\right) A_{+}(\mathbf{p})+\cos \left(\mathbf{p} \mathbf{x}+E_{p} t\right) A_{-}(\mathbf{p})\right. \\
& \left.+\sin \left(\mathbf{p} \mathbf{x}-E_{p} t\right) B_{+}(\mathbf{p})+\sin \left(\mathbf{p} \mathbf{x}+E_{p} t\right) B_{-}(\mathbf{p})\right]
\end{aligned}
$$

where
$A_{ \pm}(\mathbf{p})=v_{1}^{(+)}(\mathbf{p}) A_{ \pm}^{1}(\mathbf{p})+v_{2}^{(+)}(\mathbf{p}) A_{ \pm}^{2}(\mathbf{p})+v_{1}^{(-)}(\mathbf{p}) A_{ \pm}^{3}(\mathbf{p})+v_{2}^{(-)}(\mathbf{p}) A_{ \pm}^{4}(\mathbf{p})$,
$B_{ \pm}(\mathbf{p})=v_{1}^{(+)}(\mathbf{p}) B_{ \pm}^{1}(\mathbf{p})+v_{2}^{(+)}(\mathbf{p}) B_{ \pm}^{2}(\mathbf{p})+v_{1}^{(-)}(\mathbf{p}) B_{ \pm}^{3}(\mathbf{p})+v_{2}^{(-)}(\mathbf{p}) B_{ \pm}^{4}(\mathbf{p})$, and

$$
A_{ \pm}^{1}=\left(1 \pm \frac{p}{E_{p}}\right) c_{1} \mp \frac{m}{E_{p}} d_{2}, \quad A_{ \pm}^{2}=\left(1 \pm \frac{p}{E_{p}}\right) c_{2} \mp \frac{m}{E_{p}} d_{1}
$$

## General solution of the evolution equation

$$
\begin{aligned}
& A_{ \pm}^{3}=\left(1 \mp \frac{p}{E_{p}}\right) d_{1} \pm \frac{m}{E_{p}} c_{2}, \quad A_{ \pm}^{4}=\left(1 \mp \frac{p}{E_{p}}\right) d_{2} \pm \frac{m}{E_{p}} c_{1} \\
& B_{ \pm}^{1}=-\left(1 \pm \frac{p}{E_{p}}\right) c_{2} \mp \frac{m}{E_{p}} d_{1}, \quad B_{ \pm}^{2}=\left(1 \pm \frac{p}{E_{p}}\right) c_{1} \pm \frac{m}{E_{p}} d_{2} \\
& B_{ \pm}^{3}=-\left(1 \mp \frac{p}{E_{p}}\right) d_{2} \pm \frac{m}{E_{p}} c_{1}, \quad B_{ \pm}^{4}=\left(1 \mp \frac{p}{E_{p}}\right) d_{1} \mp \frac{m}{E_{p}} c_{2}
\end{aligned}
$$

Here $p \equiv|\mathbf{p}|, E_{p}=\sqrt{\mathbf{p}^{2}+m^{2}}$, the amplitudes $c_{1}, c_{2}, d_{1}, d_{2}$ are the ones introduced earlier. $\mathbf{p}$ is the eigenvalue of the axial momentum.

In the massive case one can not have $A_{-}=0=B_{-}$and $\left(A_{+}\right)^{2}+\left(B_{+}\right)^{2}>0$, or vice versa. Always a component propagating in the opposite direction is present. It can be weak, $\sim m / E_{p}$.

## Relativistic invariance

1. The Poincaré transformations of the real bispinor $\psi(x)$ have the standard form,

$$
\psi_{L, a}^{\prime}(x)=S(L) \psi\left(L^{-1}(x-a)\right),
$$

where $S(L)=\exp \left(\omega_{\mu \nu}\left[\gamma^{\mu}, \gamma^{\nu}\right] / 8\right)$, and $\omega_{\mu \nu}=-\omega_{\nu \mu}$ parameterize the proper orthochronous Lorentz group, $L=\exp \left(\omega^{\mu}{ }_{\nu}\right)$.

In order to identify the pertinent unitary irreducible representations of the Poincaré group we write

$$
\psi(\mathbf{x}, t)=\frac{1}{(2 \pi)^{3 / 2}} \int d^{3} p e^{i^{\gamma} \mathbf{p} \mathbf{x}} v(\mathbf{p}, t)
$$

where $v(\mathbf{p}, t)$ is a real bispinor. The Dirac equation $\Rightarrow$

$$
\dot{v}(\mathbf{p}, t)=-i \gamma^{0} \gamma^{k} \gamma_{5} p^{k} v(\mathbf{p}, t)-i m \gamma^{0} v(-\mathbf{p}, t) .
$$

In the last term we have $v(-\mathbf{p}, t)$ because $\gamma^{0}$ anticommutes with $\gamma_{5}$. It follows that $\ddot{v}(\mathbf{p}, t)=-E_{p}^{2} v(\mathbf{p}, t)$.

## Relativistic invariance

Solving the latter equation we obtain the explicit dependence on $t$

$$
\psi(\mathbf{x}, t)=\frac{1}{(2 \pi)^{3 / 2}} \int \frac{d^{3} p}{E_{p}} e^{i \gamma_{5} \mathbf{p} \mathbf{x}}\left(e^{-i \gamma_{5} E_{p} t} v_{+}(\mathbf{p})+e^{i \gamma_{5} E_{p} t} v_{-}(-\mathbf{p})\right),
$$

where the amplitudes $v_{ \pm}(\mathbf{p})$ are restricted by the conditions

$$
E_{p} v_{ \pm}(\mathbf{p})=\gamma^{0} \gamma^{k} p^{k} v_{ \pm}(\mathbf{p}) \pm m \gamma_{5} \gamma^{0} v_{\mp}(\mathbf{p})
$$

The Lorentz transformations of $\psi(x)$ are equivalent to the following transformation of the bispinor amplitudes,

$$
\begin{equation*}
v_{ \pm}^{\prime}(p)=S(L) v_{ \pm}\left(L^{-1} p\right) \tag{2}
\end{equation*}
$$

The spatial translations $\mathbf{x}^{\prime}=\mathbf{x}+\mathbf{a}$ are represented by

$$
v_{ \pm}^{\prime}(\mathbf{p}, t)=e^{\mp i \gamma_{5} \mathbf{p a}} v_{ \pm}(\mathbf{p}, t)
$$

## Relativistic invariance

2. In the massive case, $v_{-}(\mathbf{p})$ can be expressed by $v_{+}(\mathbf{p})$. The scalar product acquires explicitly Lorentz invariant form

$$
\left\langle\psi_{1} \mid \psi_{2}\right\rangle=\frac{1}{m^{2}} \int \frac{d^{3} p}{E_{p}} \overline{v_{1+}(\mathbf{p})}\left(\gamma^{0} E_{p}-\gamma^{k} p^{k}\right) v_{2+}(\mathbf{p}),
$$

where $\overline{v_{1+}(\mathbf{p})}=v_{1+}(\mathbf{p})^{T} \gamma^{0}$.
Representation (2) is unitary with respect to this scalar product. It is irreducible, and equivalent to a real version of the standard spin $1 / 2$ unitary representation.
For comparison, in the case of massive Dirac particle there are two spin $1 / 2$ representations.
3. In the massless case, the bispinors $v_{+}(\mathbf{p}), v_{-}(\mathbf{p})$ are independent. They are restricted by the conditions

$$
E_{p} v_{ \pm}(\mathbf{p})=\gamma^{0} \gamma^{k} p^{k} v_{ \pm}(\mathbf{p}), \quad E_{p}=|\mathbf{p}| .
$$

The resulting linear subspaces of real bispinors are two dimensional. Each spans the representation (2).

## Relativistic invariance

These representations are irreducible, unitary, and characterized by helicities $\pm 1 / 2$. The Lorentz invariant form of the scalar product:

$$
\begin{aligned}
&\left\langle\psi_{1} \mid \psi_{2}\right\rangle=-2 \int \frac{d^{3} p}{|\mathbf{p}|}\left[\overline{w_{1+}(\mathbf{p})}\left(\gamma^{0}|\mathbf{p}|-\gamma^{k} p^{k}\right) w_{2+}(\mathbf{p})\right. \\
&\left.+\overline{w_{1-}(\mathbf{p})}\left(\gamma^{0}|\mathbf{p}|-\gamma^{k} p^{k}\right) w_{2-}(\mathbf{p})\right]
\end{aligned}
$$

where $w_{ \pm}(\mathbf{p})$ are related to $v_{ \pm}(\mathbf{p})$ by

$$
v_{ \pm}(\mathbf{p})=\left(\gamma^{0}|\mathbf{p}|-\gamma^{k} p^{k}\right) w_{ \pm}(\mathbf{p}) .
$$

Interestingly, the last formula determines $w_{ \pm}$up to a gauge transformation of the form

$$
w_{ \pm}^{\prime}(\mathbf{p})=w_{ \pm}(\mathbf{p})+\left(\gamma^{0}|\mathbf{p}|-\gamma^{k} p^{k}\right) \chi_{ \pm}(\mathbf{p})
$$

with arbitrary $\chi_{ \pm}(\mathbf{p})$, because $\left(\gamma^{0}|\mathbf{p}|-\gamma^{k} p^{k}\right)^{2}=0$ (nilpotency). The scalar product is invariant with respect to these transformations.

## Summary

- The axial momentum $\mathbf{p}_{5}=-i \gamma_{5} \nabla$ is a viable observable for the Majorana particle (to replace $\mathbf{p}=-i \nabla$ ). Sensitive to $m \neq 0$.
- The eigenvectors of $\mathbf{p}_{5}$ are not stationary states. The minimal stationary block in the Hilbert space is spanned by the two modes: $\mathbf{p},-\mathbf{p}$.
- The main features of the expansion into the eigenvectors of $\mathbf{p}_{5}$ include:
- the presence of quaternions
- SO(4) 'phase' factors
- the presence of pairs of plane waves traveling in the opposite directions ( $1: m / E$ )
- Relativistic invariance: as expected, and simple.
- Example of QM without complex numbers.
H. A., Phys. Lett. A 383 (2019) 1242-1246 (arxiv:1805:03016);
H.A. and Z. Świerczyński, a forthcoming paper.


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