## The Ternacular <br> of <br> the <br> 

Jacob L. Bourjaily
Cracow School of Theoretical Physics
LVI Course, 2016
A Panorama of Holography

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Wednesday, $25^{\text {th }}$ May Cracow School of Theoretical Physics, Zakopane
Part I: The Vernacular of the S-Matrix


















## Organization and Outline

(1) Spiritus Movens: a moral parable

- A Simple, Practical Problem in Quantum Chromodynamics
- The Shocking Simplicity of Scattering Amplitudes
(2) The Vernacular of the S-Matrix
- Physically Observable Data Describing Asymptotic States
- Massless Momenta and Spinor-Helicity Variables
- (Grassmannian) Geometry of Momentum Conservation
(3) The All-Orders S-Matrix for Three Massless Particles
- Three Particle Kinematics and Helicity Amplitudes
- Non-Dynamical Dependence: Coupling Constants \& Spin/Statistics

4. Consequences of Quantum Mechanical Consistency Conditions

- Factorization and Long-Range Physics: Weinberg's Theorem
- Uniqueness of Yang-Mills Theory and the Equivalence Principle
- The Simplest Quantum Field Theory: $\mathcal{N}=4$ super Yang-Mills


## Supercomputer Computations in Quantum Chromodynamics

Consider the amplitude for two gluons to collide and produce four: $g g \rightarrow g g g g$.

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Supercollider physics

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The Ohio State University, Columbus, Ohio 43210
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$D_{6}^{5}(7)=\frac{1}{s_{32} s_{4} t_{123}}\left[s_{16}-s_{48}+s_{36}\right]\left[s_{12}-s_{13}-s_{23}\right]$.
$D_{0}^{z}(8)=\frac{1}{s_{4} s_{23} s_{14}}\left[s_{13}+s_{13}-s_{24}\right]\left[s_{14}-s_{44}+s_{54}\right]$.
$D_{0}^{5}(9)=\frac{1}{s_{23} s_{54} t_{134}}\left[s_{14}+s_{54}-s_{13}\right]\left[s_{56}-s_{59}+s_{33}\right]$,
$D_{0}^{s}(10)-\frac{1}{s_{21} s_{s}}\left(p_{2}-p_{s}\right)\left(p_{1}-p_{s}\right)$.
$D_{6}^{5}(11)=\frac{1}{s_{1} s_{5} s_{6}}\left(p_{1}-p_{4}\right)\left(p_{s}-p_{v}\right)$,
$D_{0}^{s}(12)=\frac{1}{s_{4} s_{3 y}}\left(p_{t}-p_{1}\right)\left(p_{2}-p_{3}\right)$,
$D_{0}^{5}(13)=\frac{1}{s_{1}, s_{34}}\left(p_{3}-p_{1}\right)\left(p_{1}-p_{4}\right)$.
$D_{\delta}^{\Sigma}(14)=\frac{1}{s_{4} s_{4} s_{4}}\left(p_{2}-p_{y}\right)\left(p_{3}-p_{4}\right)$,
$D_{0}^{8}(15)=\frac{1}{s_{4} s_{3} s_{3}}\left\{\left[\left(p_{2}+p_{9}\right)\left(p_{3}-p_{0}\right)\right]\left[\left(p_{1}-p_{4}\right)\left(p_{2}-p_{3}\right)\right]\right.$
$+\left[\left(p_{2}-p_{9}\right)\left(p_{3}-p_{4}\right)\right]\left[\left(p_{1}-p_{4}\right)\left(p_{3}+p_{4}\right)\right]$
$\left.+\left[\left(p_{1}+p_{2}\right)\left(p_{2}-p_{3}\right)\right]\left[\left(p_{1}-p_{1}\right)\left(p_{3}-p_{1}\right)\right]\right\}$.
$D_{0}^{5}(16)-\frac{2}{s_{4} s_{4} 5_{3}}\left[\left\{\left(p_{2}-p_{j}\right)\left(p_{y},+p_{4}\right)\right\}\left(p_{1}-p_{b}\right)\left(p_{3}-p_{4}\right)\right]$
$+\left[\left(p_{1}+p_{p}\right)\left(p_{3}-p_{c}\right)\right]\left[\left(p_{1}-p_{0}\right)\left(p_{2}-p_{s}\right)\right]$
$\left.+\left[\left(p_{1}-p_{0}\right)\left(p_{2}+p_{3}\right)\right]\left[\left(p_{1}-p_{1}\right)\left(p_{2}-p_{3}\right)\right]\right)$.
The preceding list completes the result. Let us recapitualate now the numerical procedure of calculating the full cross section. First the diagrams $D$ are calculated by using eqs. (111)-(13). The result is substituted to eq. (8) to obtain the vectors $S_{0}$ permutations of momentas, eq. (6) is used to obtain the functions $A_{0}$ and $A_{2}$. Finaily, the total cross section is calculated by using eq. (5). The FORTRAN 5 program based on such a scheme generates ten Monte Carlo points in less than a second on the heterotic CDC CYBER $175 / 875$.
Given the complexity of the final result, it is very important to have some reliable testing procedures available for numerical calculations. Usually in QCD, the multisloon amplitudes are tested by checking the gauge invariance. Due to the specifics

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${ }^{20}$
of our calculation, the most powerfil test does not rely on the gauge symmetry, but on the appropriate permutation symmetries. The function $A_{0}\left(p_{1}, p_{2}, p_{4}, p_{1}, p_{3}, p_{0}\right)$
must be symmetric under arbitrary permutations of the momenta $\left(p_{1}, p_{2}, p_{1}\right)$ and separately, ( $p_{4}, p_{s}, p_{4}$ ), whereas the function $A_{2}\left(p_{1}, p_{3}, p_{3}, p_{4}, p_{s}, p_{0}\right)$ must be symmetric under the permutations of ( $\left.p_{1}, p_{2}, p_{3}, p_{4}\right)$ and separately, $\left(p_{3}, p_{3}\right)$. This test is extremely powerful, because the required permutation symmetries are hidden in our supersymmetry relations, eqs. (1) and (3), and in the structure of amplitudes involving different species of particles. Another, very important test relies on the absence of the double poles of the form $\left(g_{0}\right)^{-2}$ in the cross section, as required by general arguments based on the helicity conservation. Further, in the leading $\left(s_{y}\right)^{-1}$ $[3,4]$, convoluted with the appropriate Altarelli-Parisi probabilities [5]. Our result has saccesfully passed both these numerical checks.
Details of the calculation, together with a full exposition of our techniques, will be given in a forthcoming article. Furthermore, we hope to obtain a simple analytic form for the answer, making our result not only an experimentalist's, but also a theorist's deligh
We thank Keith Ellis, Chris Quigs and especially, Estia Eichten for many usefal discussions and encouragement during the course of this work. We acknowledge the hospitality of Aspen Center for Physics, where this work was being completed in a pleasant, strung-out atmosphere.

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Physically Observable Data Describing Asymptotic States

## On What Data Does a Scattering Amplitude Depend?

A scattering amplitude, $\mathcal{A}_{n}$, can be a generally complicated(?) function of all the physically observable data describing each of the particles involved.

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Physically Observable Data Describing Asymptotic States

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Physically Observable Data Describing Asymptotic States

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\end{array}\right) \equiv\binom{\lambda^{1}}{\lambda^{2}} \quad \tilde{\lambda} \equiv\left(\begin{array}{llll}
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\end{array} \widetilde{\lambda}_{n}\right) \equiv\binom{\widetilde{\lambda}^{\mathrm{i}}}{\widetilde{\lambda}^{\dot{2}}}
$$

writing $\lambda_{a} \in \mathbb{C}^{2}$ for a column, $\lambda^{\alpha} \in \mathbb{C}^{n}$ for a row.

- Because Lorentz transformations mix the rows of each matrix, $\lambda^{\alpha}, \widetilde{\lambda}^{\dot{\alpha}}$, and the little group allows for rescaling, the invariant content of the data is:


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 the span of $k$ vectors in $\mathbb{C}^{n}$- Momentum conservation:



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Thus, all the kinematical data can be described by a pair of $(2 \times n)$ matrices:

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## Building Blocks: the S-Matrix for Three Massless Particles

Momentum conservation and Poincaré-invariance uniquely fix the kinematical dependence of the amplitude for three massless particles (to all loop orders!).

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\lambda \equiv\left(\begin{array}{lll}
\lambda_{1}^{1} & \lambda_{2}^{1} & \lambda_{3}^{1} \\
\lambda_{1}^{2} & \lambda_{2}^{2} & \lambda_{3}^{2}
\end{array}\right)
$$

$$
\widetilde{\lambda} \equiv\left(\begin{array}{ccc}
\widetilde{\lambda}_{1}^{i} & \widetilde{\lambda}_{2}^{i} & \widetilde{\lambda}_{3}^{i} \\
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\begin{aligned}
& \lambda^{\perp} \equiv(\langle 23\rangle\langle 31\rangle\langle 12\rangle) \supset \widetilde{\lambda} \\
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\lambda_{1}^{1} & \lambda_{2}^{1} & \lambda_{3}^{1} \\
\lambda_{1}^{2} & \lambda_{2}^{2} & \lambda_{3}^{2}
\end{array}\right) \\
& \text { or } \\
& \widetilde{\lambda} \equiv\left(\begin{array}{lll}
\widetilde{\lambda}_{1}^{i} & \widetilde{\lambda}_{2}^{i} & \widetilde{\lambda}_{3}^{i} \\
\widetilde{\lambda}_{1}^{i} & \widetilde{\lambda}_{2}^{i} & \widetilde{\lambda}_{3}^{\dot{i}}
\end{array}\right) \\
& \widetilde{\lambda}^{\perp} \equiv\left(\left[\begin{array}{lll}
23] & {[31]} & [12]) \supset \lambda
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$$
\begin{aligned}
& h_{1} \\
& \stackrel{\langle 12\rangle^{h_{3}-h_{1}-h_{2}}\langle 23\rangle^{h_{1}-h_{2}-h_{3}}\langle 31\rangle^{h_{2}-h_{3}-h_{1}}}{\longrightarrow} \mathcal{O}\left(\epsilon^{-\left(h_{1}+h_{2}+h_{3}\right)}\right) \quad \lambda \equiv\left(\begin{array}{lll}
\lambda_{1}^{1} & \lambda_{2}^{1} & \lambda_{3}^{1} \\
\lambda_{1}^{2} & \lambda_{2}^{2} & \lambda_{3}^{2}
\end{array}\right) \\
& \begin{aligned}
{[12]^{h_{1}+h_{2}-h_{3}}[23]^{h_{2}+h_{3}-h_{1}}[31]^{h_{3}+h_{1}-h_{2}} } & \widetilde{\lambda} \equiv\left(\begin{array}{lll}
\widetilde{\lambda}_{1}^{i} & \widetilde{\lambda}_{2}^{i} & \widetilde{\lambda}_{3}^{i} \\
\widetilde{\lambda}_{1}^{2} & \widetilde{\lambda}_{2}^{i} & \widetilde{\lambda}_{3}^{2}
\end{array}\right) \\
{[a b] \rightarrow \mathcal{O}(\epsilon) } & \mathcal{O}\left(\epsilon^{\left(h_{1}+h_{2}+h_{3}\right)}\right)
\end{aligned} \widetilde{\lambda}^{\perp} \equiv([23][31][12]) \supset \lambda, ~\left[\begin{array}{ll}
{[3]}
\end{array}\right.
\end{aligned}
$$

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\lambda_{1}^{2} & \lambda_{2}^{2} & \lambda_{3}^{2}
\end{array}\right) \\
& \begin{aligned}
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$$
\left.\begin{array}{rlrl}
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\lambda_{1}^{1} & \lambda_{2}^{1} & \lambda_{3}^{1} \\
\lambda_{1}^{2} & \lambda_{2}^{2} & \lambda_{3}^{2}
\end{array}\right), ~ l
$$

$$
h_{1}+h_{2}+h_{3} \leq 0
$$

or

$$
\begin{array}{cl}
{[12]^{h_{1}+h_{2}-h_{3}}[23]^{h_{2}+h_{3}-h_{1}}[31]^{h_{3}+h_{1}-h_{2}}} & \widetilde{\lambda} \equiv\left(\begin{array}{lll}
\widetilde{\lambda}_{1}^{\mathrm{i}} & \widetilde{\lambda}_{2}^{\mathrm{i}} & \widetilde{\lambda}_{3}^{\mathrm{i}} \\
\widetilde{\lambda}_{1}^{2} & \widetilde{\lambda}_{2}^{\dot{2}} & \widetilde{\lambda}_{3}^{2}
\end{array}\right) \\
h_{1}+h_{2}+h_{3} \geq 0 & \widetilde{\lambda}^{\perp} \equiv\left(\left[\begin{array}{lll}
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\end{array}\right.\right.
\end{array}
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{[12]^{h_{1}+h_{2}-h_{3}}[23]^{h_{2}+h_{3}-h_{1}}[31]^{h_{3}+h_{1}-h_{2}}} & \widetilde{\lambda} \equiv\left(\begin{array}{lll}
\widetilde{\lambda}_{1}^{\mathrm{i}} & \widetilde{\lambda}_{2}^{\mathrm{i}} & \widetilde{\lambda}_{3}^{\mathrm{i}} \\
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