

Dark Matter and Colliders

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Counter-examples: axions; Gravitinos; FIMPs; dark atoms;
primordial black holes; keV neutrinos: not covered in this talk. **Note:**
Proves that LHC does *not* “recreate conditions of the early
universe”!

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- Indicates weak-scale $\chi\chi$ annihilation cross section:

$$\langle \sigma(\chi\chi \rightarrow \text{any})v \rangle \simeq (2 \text{ to } 4.5) \cdot 10^{-26} \text{ cm}^3 \text{ s}^{-1}$$

WIMPs and Early Universe

$\Omega_\chi h^2$ can be changed **a lot** in non-standard cosmologies (involving $T \gg T_{\text{BBN}}$):

- Increased: Higher expansion rate $H(T \sim T_F)$; additional non-thermal χ production at $T < T_F$; ...

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Determining $\sigma(\chi\chi \rightarrow \text{SM})$ allows probe of very early Universe, once χ has been established to be “the” DM particle! e.g. MD, Iminniyaz, Kakizaki, arXiv:0704.1590

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- Thermal WIMP: Only know *total* $\chi\chi \rightarrow$ SM cross section; contribution of specific final states (e^+e^- , $u\bar{u} + d\bar{d}$) *not* known
- $\Omega_\chi h^2$ determined from $\sigma(\chi\chi \rightarrow \text{SM})$ near threshold ($T_F \simeq m_\chi/20 \implies s \simeq 4m_\chi^2$). At colliders need ≥ 3 body final state to get signature (e.g. $e^+e^- \rightarrow \chi\chi\gamma$, $q\bar{q} \rightarrow \chi\chi g$) \implies typically need $\sigma(\chi\chi \rightarrow \text{SM})$ at $s \gtrsim 6$ to $10m_\chi^2$!

“Model-independent” approach

Goodman et al., arXiv:1005.1286 and 1008.1783; Bai, Fox, Harnik, arXiv:1005.3797; Wang, Li, Shao, Zhang, arXiv:1107.2048; Fox, Harnik, Kopp, Tsai, arXiv:1103.0240

Parameterize χ interaction with relevant SM fermion through dim-6 operator; e.g. for hadron colliders:

$$\mathcal{L}_{\text{eff}} = G_{\chi,q} \bar{\chi} \Gamma_{\chi} \chi \bar{q} \Gamma_q q$$

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χ Majorana $\implies \Gamma_{\chi} \in \{1, \gamma_5, \gamma_{\mu} \gamma_5\}$

$\Gamma_q \in \{1, \gamma_5, \gamma_{\mu}, \gamma_{\mu} \gamma_5\}$

If $\Gamma_{\chi}, \Gamma_q \in \{1, \gamma_5\}$: $G_{\chi,q} = m_q / (2M_*^3)$ (chirality violating!),
else $\Gamma_{\chi} = 1 / (2M_*^2)$ Rajamaran, Shepherd, Tait, Wijango, arXiv:1108.1196.

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Compare monojet signal from $q\bar{q} \rightarrow \chi\chi g$ with monojet limits (current bound) and background (ultimate reach)!

UV completion

Approach can only work if the “mediator” mass m_M is (much) larger than the highest relevant momentum scale:

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LHC can only hope to be sensitive if χ couples to quarks at tree level.

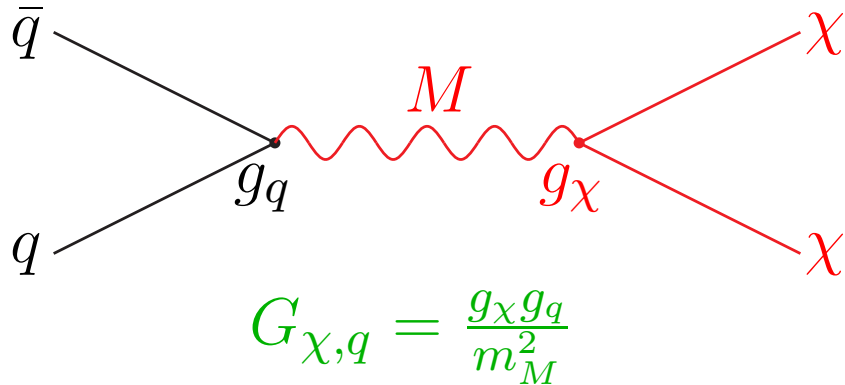
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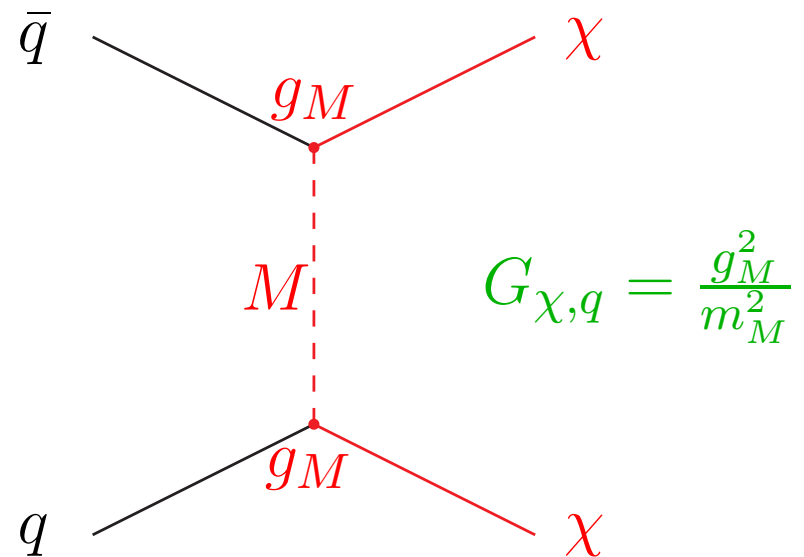
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s -channel



t -channel



Types of interactions

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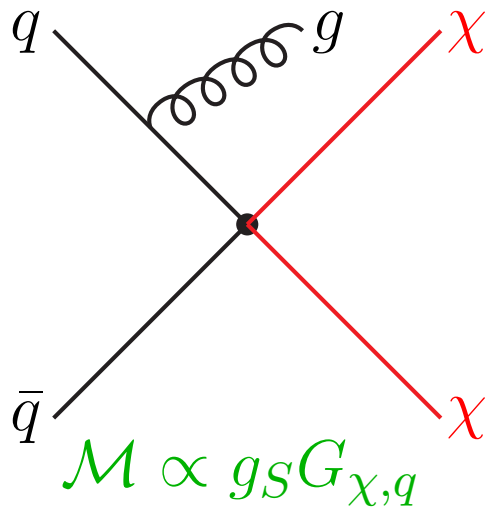
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Case $\Gamma_\chi = \Gamma_q = \gamma_5$ also has poor LHC reach. Gives velocity-dependent interaction for $\chi p \rightarrow \chi p \implies$ very poor reach in direct detection as well

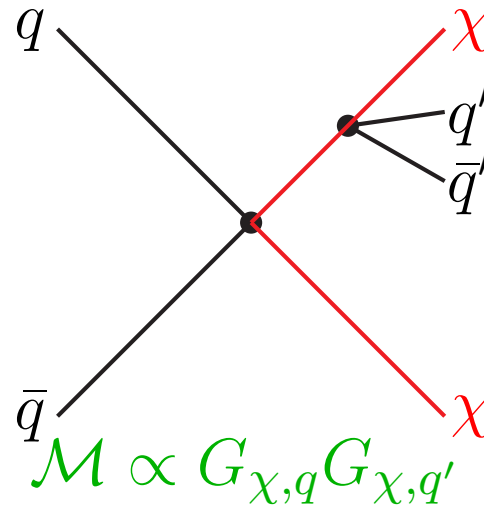
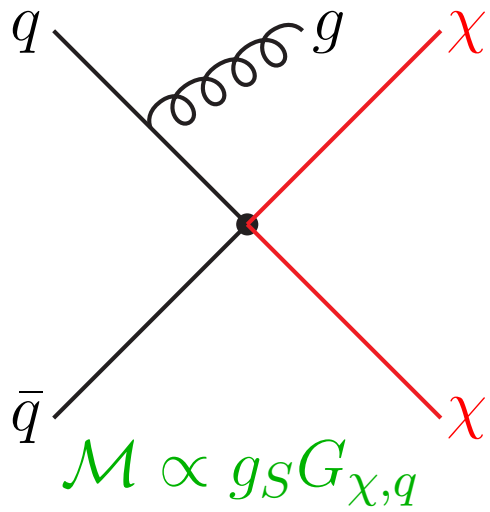
Monojet analysis is not model-independent!

Contributing diagrams:



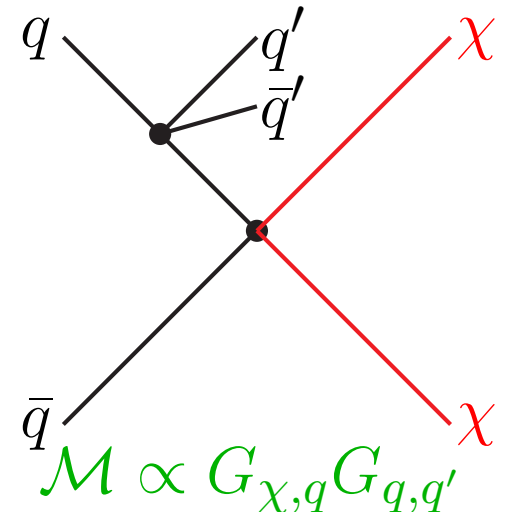
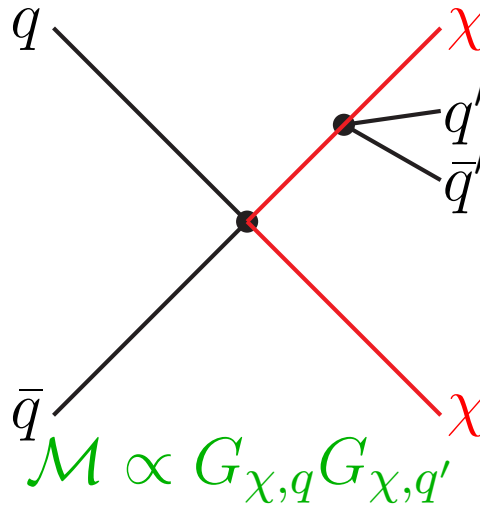
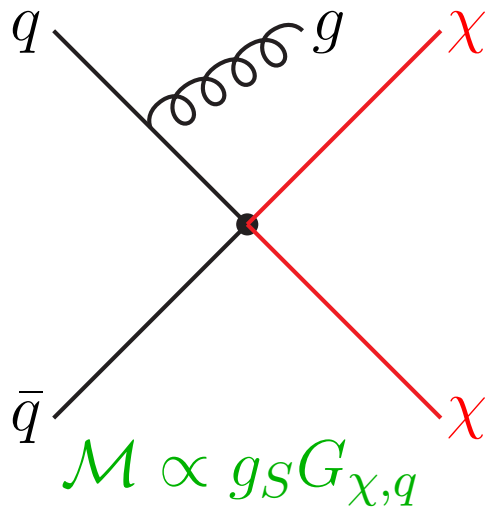
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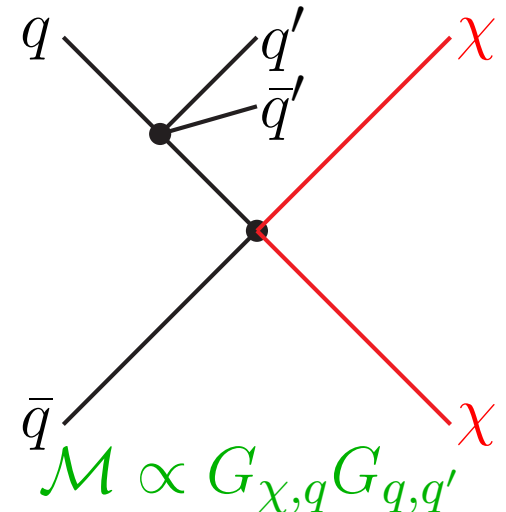
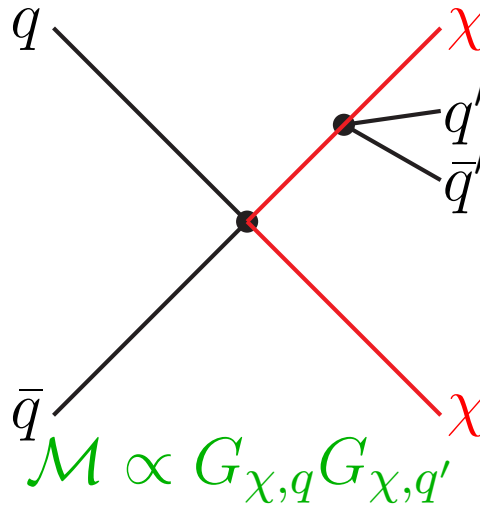
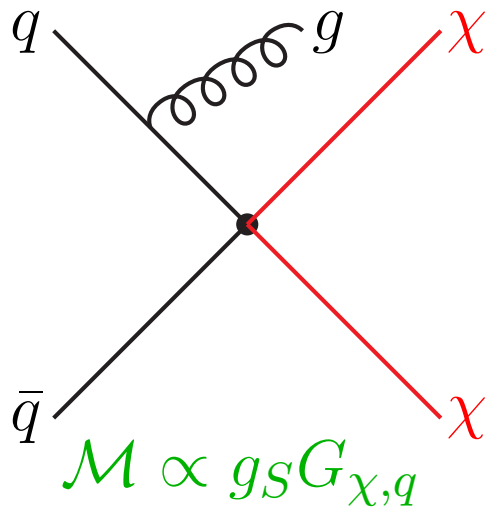
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$$G_{q,q'} = \begin{cases} \frac{g_q g_{q'}}{m_M^2}, & s - \text{channel} \\ 0, & t - \text{channel} \end{cases}$$

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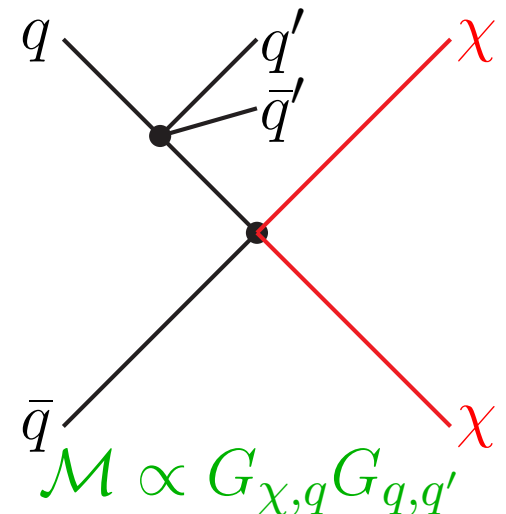
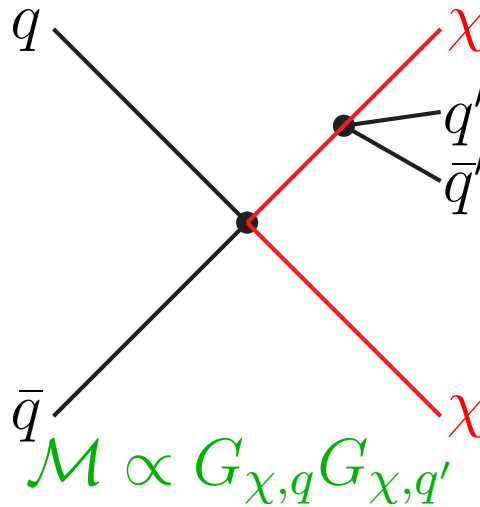
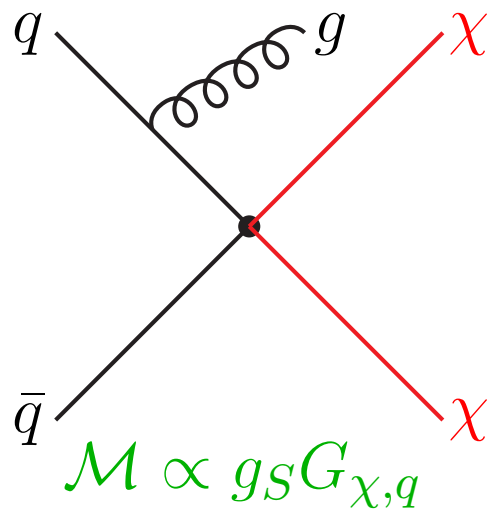


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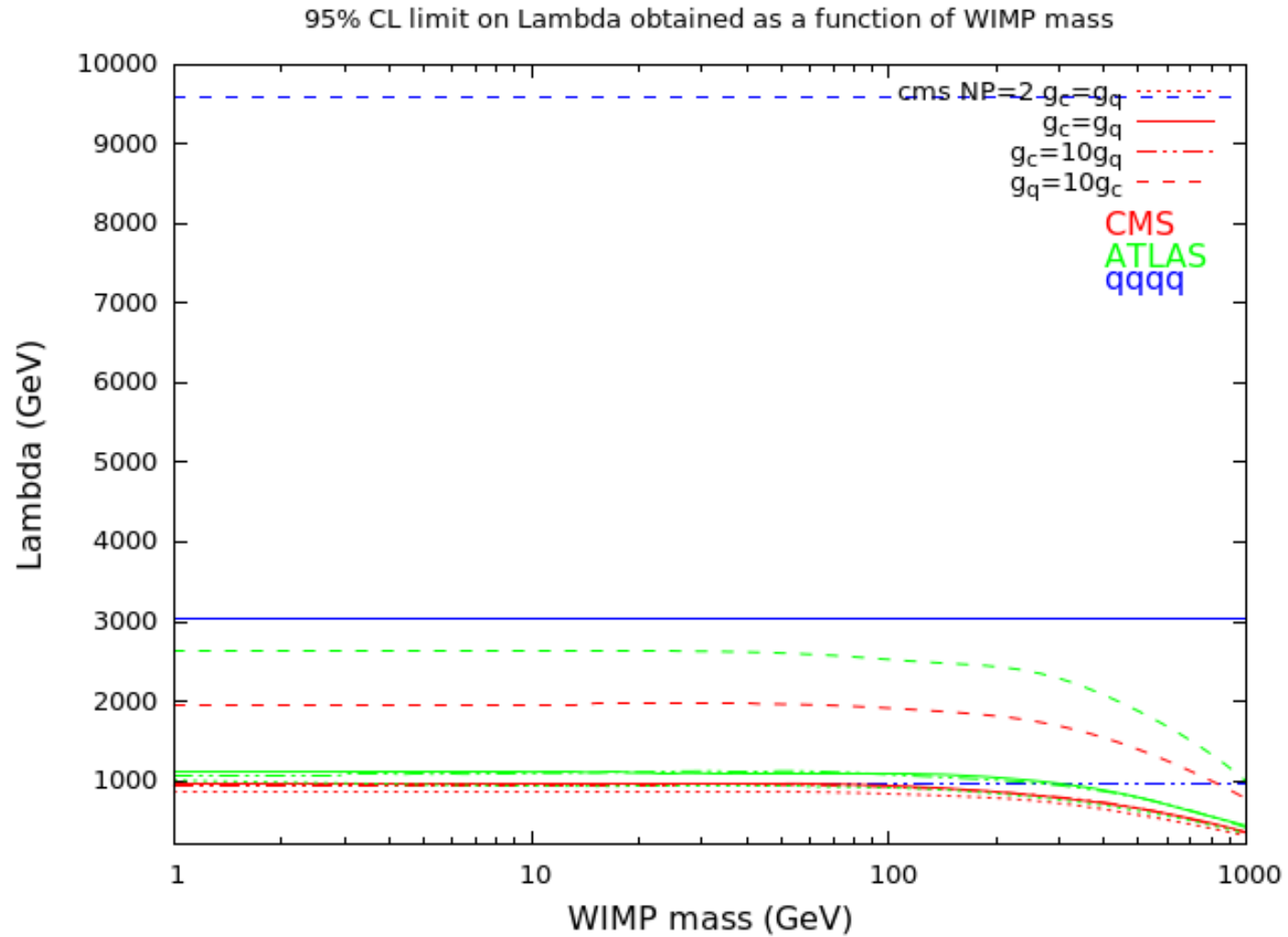
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For s -channel: bound on $4q$ contact interaction stronger than bound from monojet searches, unless $g_\chi \gg g_q$!

Bounds on Λ

S. Belwal, MD, J.S. Kim, in preparation



Promising collider searches

Look for mediators!

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t -channel: mediator carries color \Rightarrow can be pair-produced!

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For perturbative couplings: search for off-shell mediator $\rightarrow \chi\chi$ is hopeless! E.g. SUSY: signal for $\tilde{\chi}\tilde{\chi}j$ is *much* smaller than $Z (\rightarrow \nu\bar{\nu})j$ background, even for 100 GeV higgsino-like $\tilde{\chi}$ (Baer, Mustafayev, Tata, arXiv:1401.1162)

DM and Light (Gauge) Bosons

(At least) 3 kinds of WIMP models require light ($m \leq \text{few GeV}$) (gauge) bosons U :

- MeV DM: Suggested as explanation of 511 keV line (\Rightarrow slow e^+) excess from central region of our galaxy (Boehm et al., astro-ph/0309686). Should have $m_\chi \leq 10 \text{ MeV}$ (γ constraints)
 $\Rightarrow m_\chi \leq m_U \leq 200 \text{ MeV}$ to mediate $\chi\chi \rightarrow e^+e^-$; fixes $g_{U\chi\chi}g_{Ue^+e^-}/m_U^2$! (Unless $2m_\chi \simeq m_U$.)

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- PAMELA/FermiLAT inspired TeV DM: Needs light boson for Sommerfeld enhancement (e.g. Arkani-Hamed et al., arXiv:0810.0713(4)) ($\chi\chi \rightarrow UU \rightarrow 4l$ is also somewhat less constrained by γ spectrum than $\chi\chi \rightarrow 2l$.)

- DAMA/CoGeNT inspired few GeV DM: Needs light mediator to achieve sufficiently large $\sigma_{\chi p}$. (2 different mediators for isospin violation to evade bounds: Cline, Frey, arXiv:1108.1391)

Light Gauge Bosons (cont'd)

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$U_{\chi\chi}$ coupling may well be large.

Signatures of light gauge bosons

If $m_U > 2m_\chi$: $U \rightarrow \chi\chi$ dominant! Is invisible \Rightarrow need extra tag, e.g. $e^+e^- \rightarrow \gamma U \rightarrow \gamma + \text{nothing}$.

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Borodatchenkova, Choudhury, MD, hep-ph/0510147

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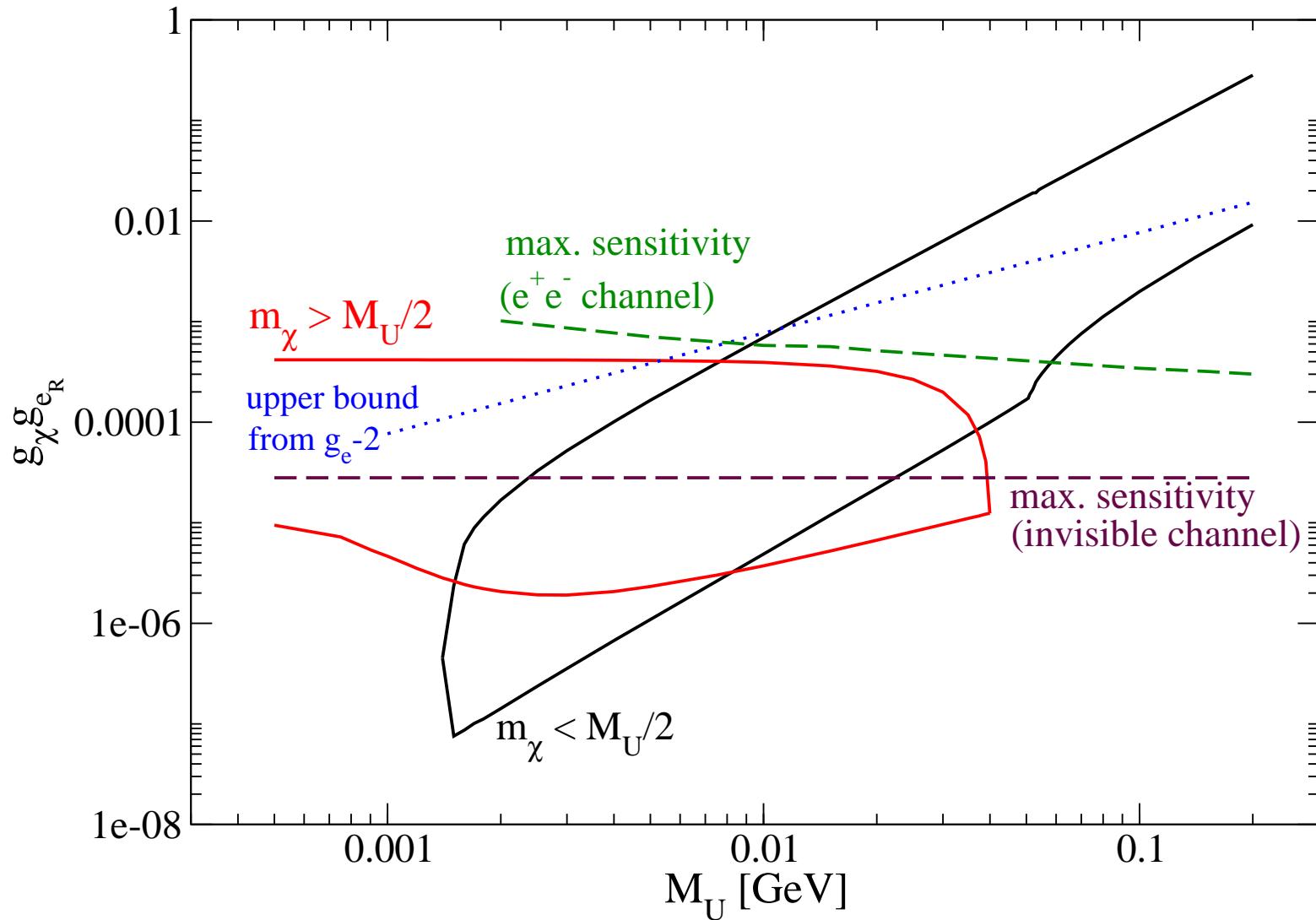
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- Instrumental backgrounds (not from e^+e^- annihilation) seem large

Sensitivity at B -factories (100 fb^{-1})



Red, black: Regions allowed by Ω_χ , $\sigma(\chi\chi \rightarrow e^+e^-)$.

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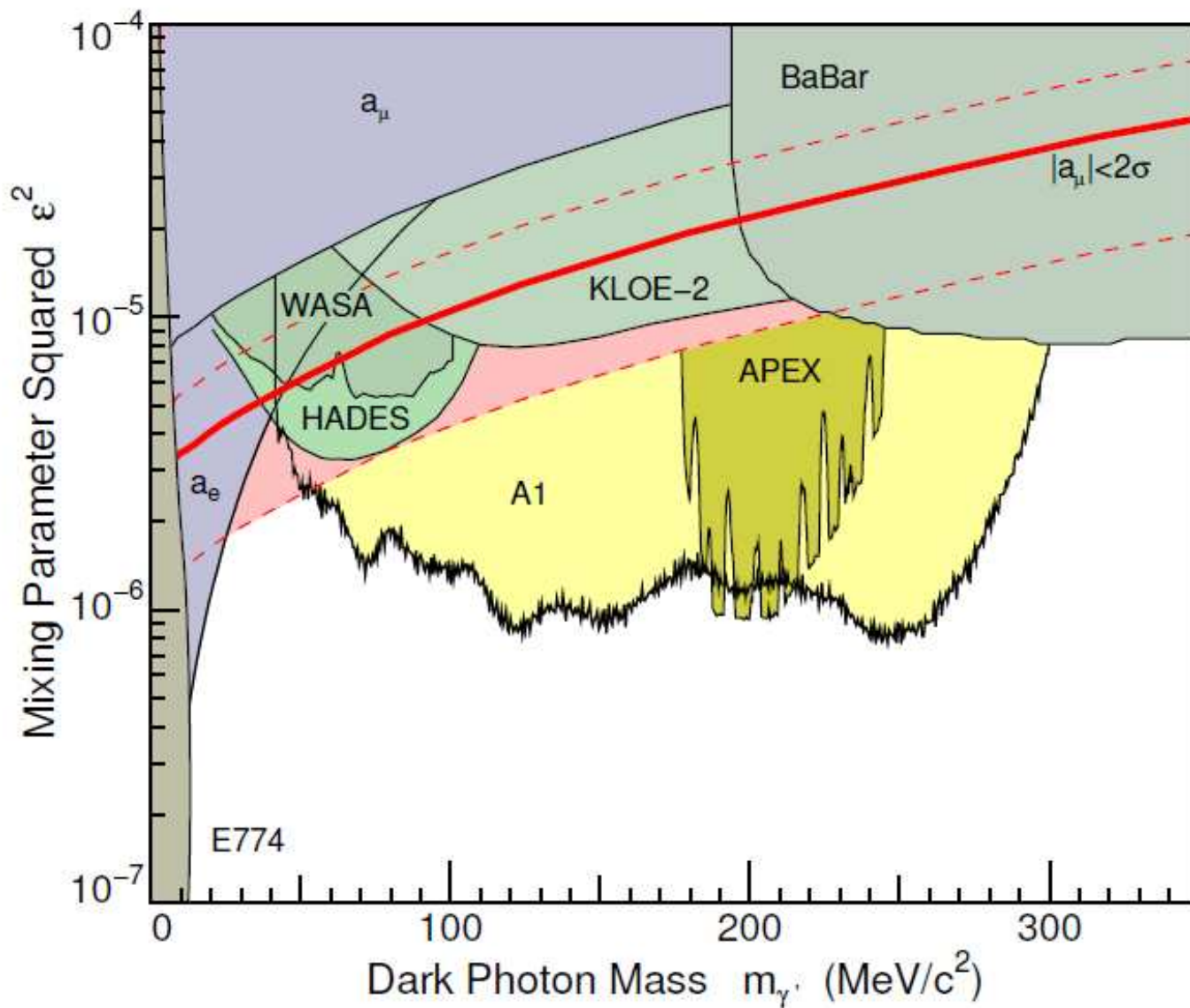
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Also, KLOE-2 performed search, mostly for $\phi \rightarrow U\eta$: no signal. [arXiv:1107.2531](#)

A1 results



SUSY DM and the LHC

Saw above: WIMP searches at colliders not promising, *if* WIMP is only accessible new particle. Fortunately, in many cases the WIMP is the lightest of *many* new particles! True in SUSY. (Also in Little Higgs, UED.)

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- HLS theorem, relation to superstrings: don't single out weak scale.

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- SUSY implies equal masses for partners \implies **SUSY must be broken**

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- In simplest, R -parity invariant scenario: lightest superparticle LSP is stable: satisfies one condition for DM candidate!

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- but DM–allowed regions of parameter space do exist even in constrained models!

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- **Note:** DM-allowed region of $(m_0, m_{1/2})$ plane of cMSSM depends on $A_0, \tan \beta$!

Generic SUSY searches at LHC

Strongly interacting sparticles have biggest cross sections; may decay via long decay “cascades”. Example:

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$\mathcal{O}(100)$ searches have been performed, but no signal has been found.

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- In pMSSM10: $m_{\tilde{\chi}_1^0} \simeq 50 \text{ GeV}$ still ok! de Vries et al., arXiv:1504.03260

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- **Most interesting to me: Predict $\Omega_\chi h^2$, compare with observation: Constrain very early universe!**

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- Scenarios with new light gauge bosons with suppressed couplings to SM fermions are now being probed at low- E colliders, fixed-target expts.
- Absence of missing E_T signal at LHC is disappointing, but plenty of parameter space in reasonably well motivated WIMP models left to explore