

**Limits on the invisible mass
in the inner Galaxy
from the microlensing
in the thin disk model**

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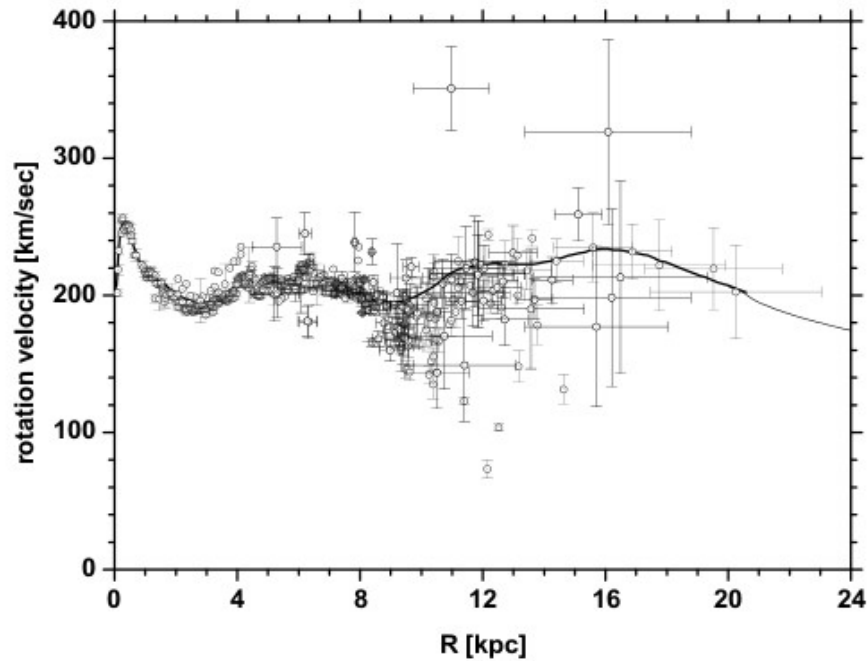
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Outline

- The global disk model of a spiral galaxy
- The microlensing as a method of measuring the mass of compact objects along a line of sight
- The microlensing towards Galactic bulge: the dark matter limit in the global disk model framework

The standard model of spiral galaxies



Spiral Galaxy NGC 4565
(FORS / VLT)

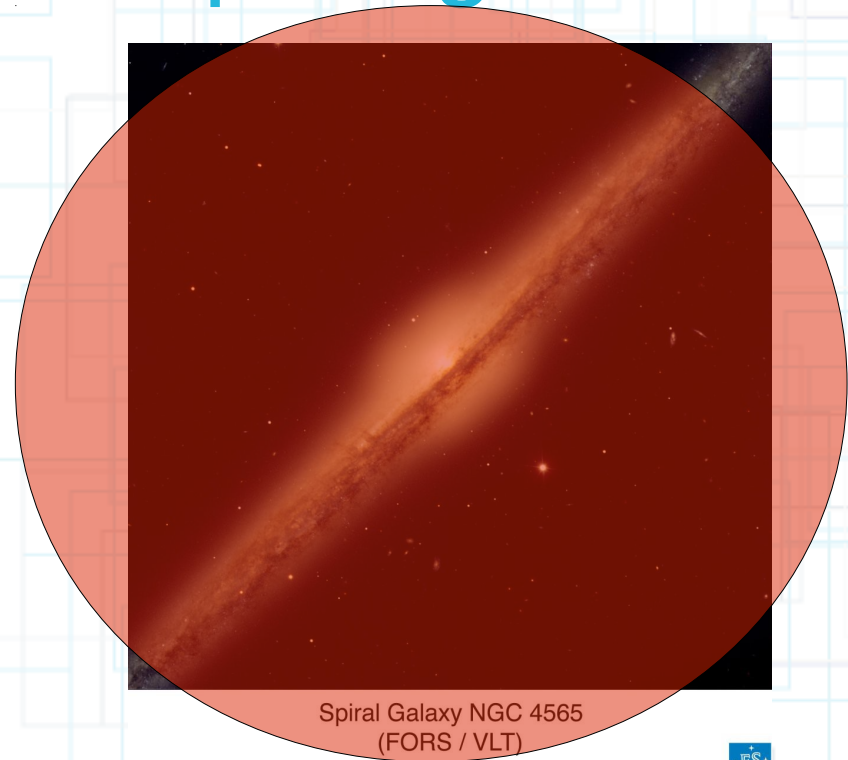
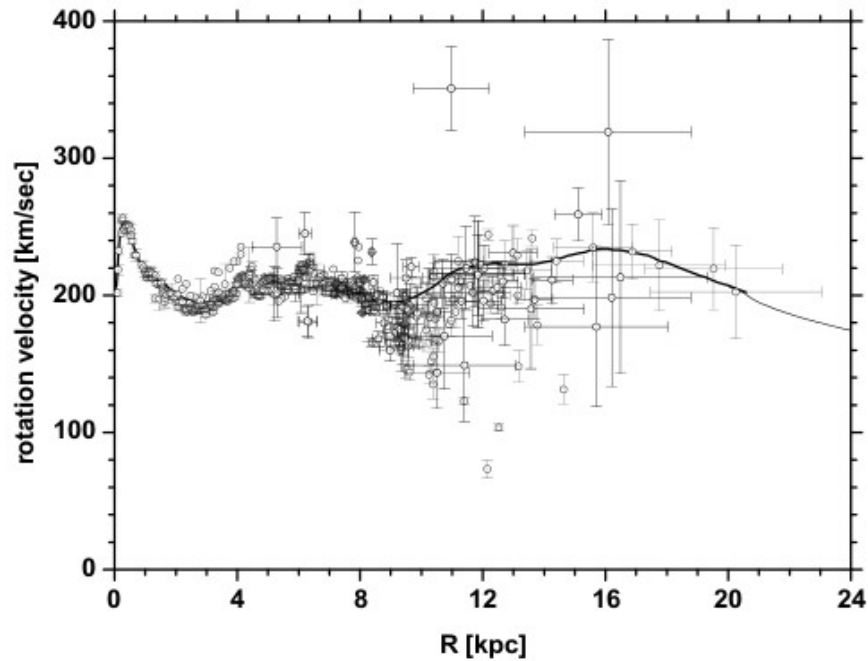
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The standard modeling of a spiral galaxy takes into account:

- the rotation curve
- the brightness profile
- M/L - the mass to light ratio
- analytical formulas with some free parameters, describing the mass density in a various components: bulge, disk, halo

The standard model of spiral galaxies

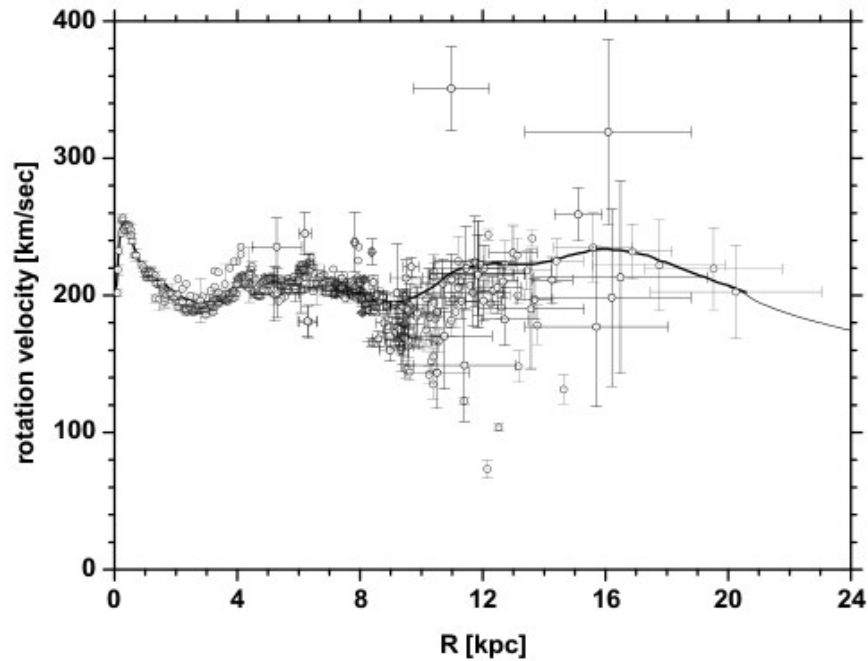


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The result of a fitting the free parameters is that the dominating contribution to the gravitational potential gives a dark matter halo

The global disk model



Spiral Galaxy NGC 4565
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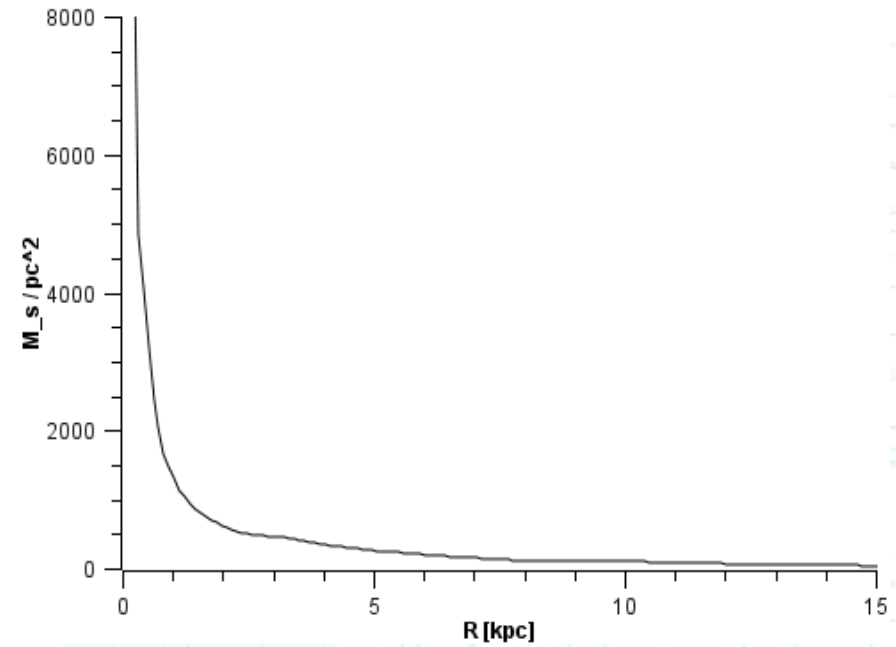
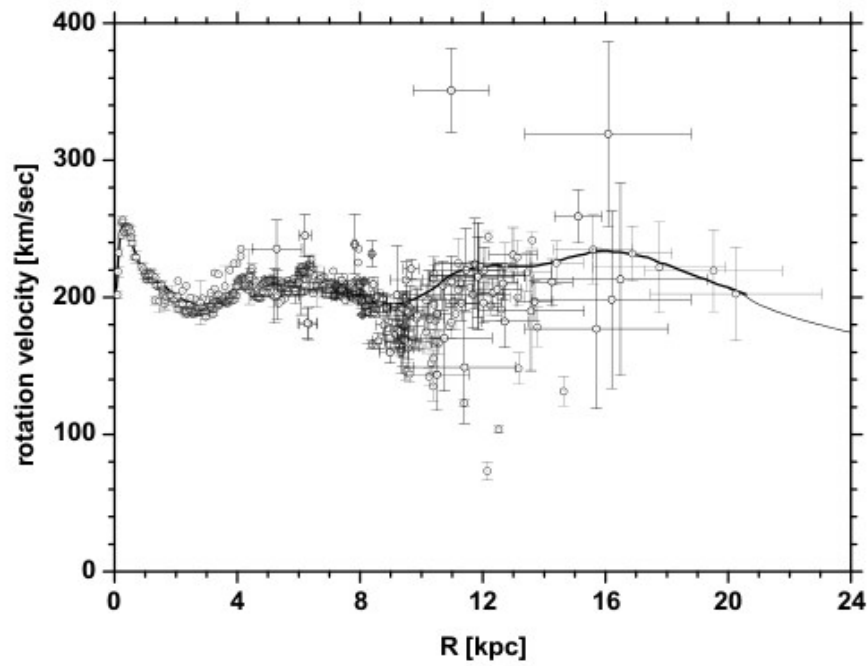
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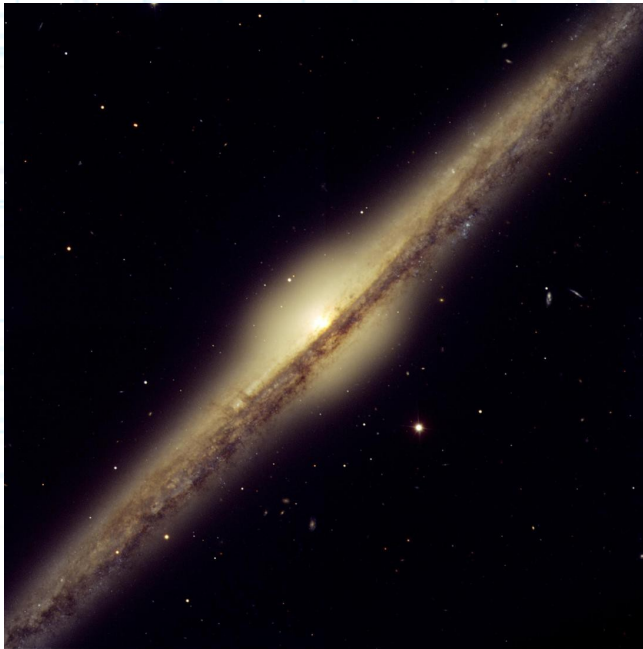
$$\sigma(R) = \frac{1}{\pi^2 G} \mathcal{P} \left[\int_0^R v^2(\chi) \left(\frac{K(\chi/R)}{R \chi} - \frac{R E(\chi/R)}{\chi R^2 - \chi^2} \right) d\chi + \int_R^\infty v^2(\chi) \frac{E(R/\chi)}{\chi^2 - R^2} d\chi \right]$$

Bratek Ł., Jałocha J., Kutschera M., 2008, MNRAS, 391, 1373

The global disk model

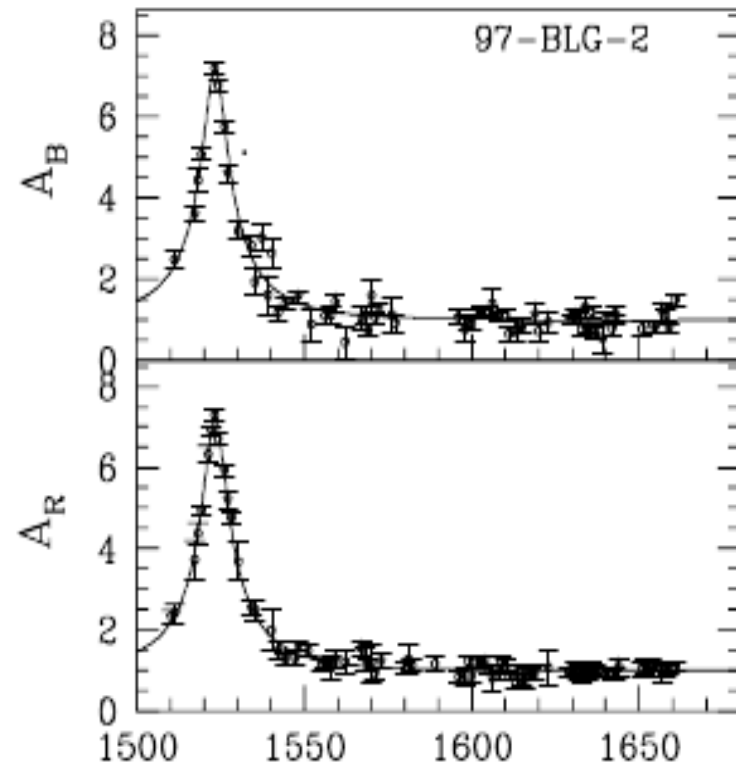


The microlensing



Spiral Galaxy NGC 4565
(FORS / VLT)

ESO PR Photo 24a/05 (August 10, 2005)



From: MACHO project:
arXiv:astro-ph/0002510v2

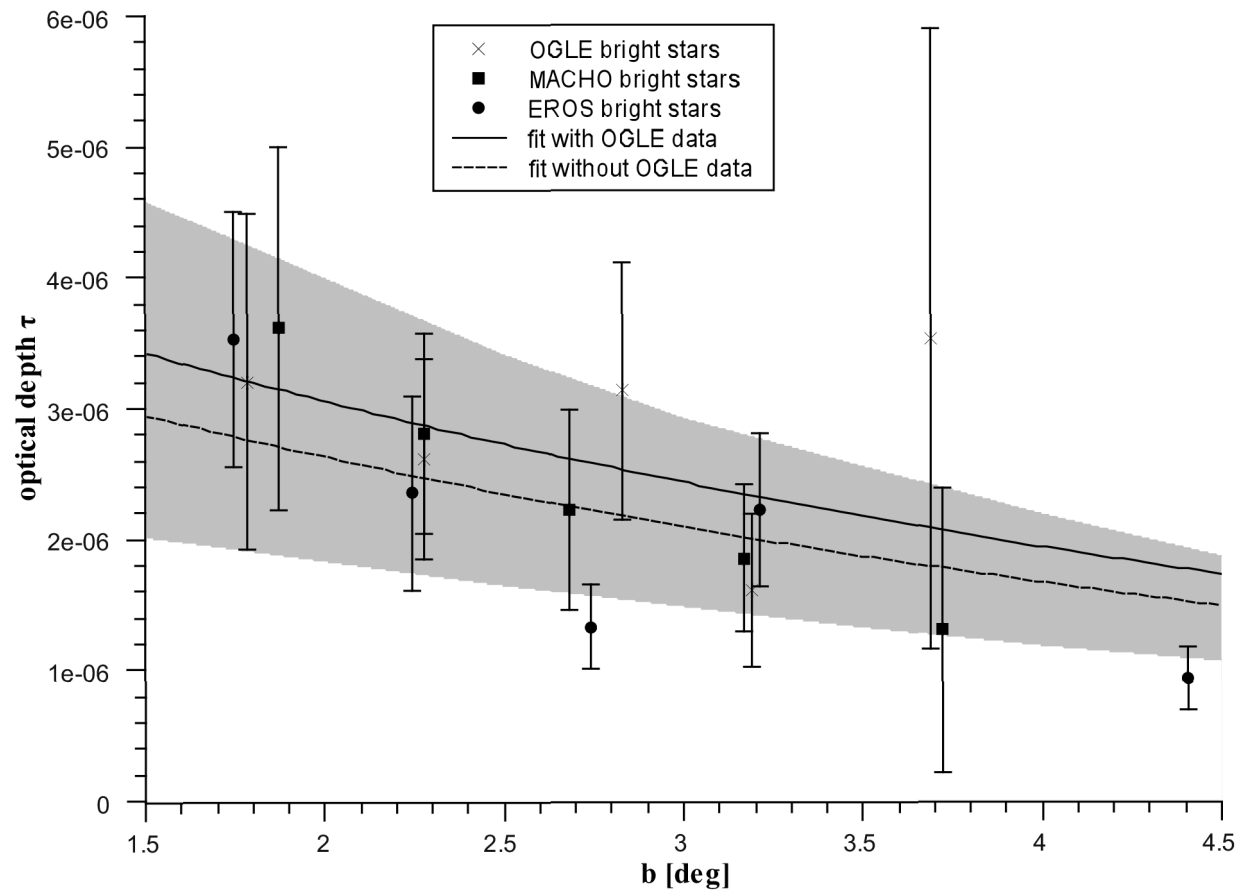
$$\tau = \int_0^{D_S} \frac{4\pi G}{c^2 D_L} \left(1 - \frac{D_L}{D_S}\right) \rho_{co}(\vec{r}(D_L)) D_L^2 dD_L$$

The microlensing

Within the global disk model we take into account:

- Various vertical density distribution profiles:
 - single exponential
 - double exponential
 - $1/\cosh(z/a)$
- The structure of central bulge:
 - spherically symmetric bulge
 - three-axial G2 bulge/bar model by Dwek et al.
- Subtraction of a gas contribution
- The Monte Carlo simulation of a sources distribution
- The blending effect – we restrict to the bright sources only

The results



Dark matter limit: $(0.11 \pm 0.41) \times 10^{10} M_{\odot}$

Additional dynamical mass: $(0.86 \pm 0.36) \times 10^{10} M_{\odot}$

The references

- [1] Sikora S., Bratek Ł., Jałocha J., Kutschera M., (2011), arXiv:1103.5056
„Gravitational microlensing as a test of a finite-width disk model of the Galaxy”
- [2] Bratek Ł., Jałocha J., Kutschera M., (2008), MNRAS, 391, 1373
„On the axisymmetric thin disk model of flattened galaxies”
- [3] Jałocha J., Bratek Ł., Kutschera M., Skindzier P., (2010), MNRAS, 407, 1689
„Transverse gradients of azimuthal velocity in a global disk model of the Milky Way”
- [4] Moni Bidin C., Carraro G., Mendez R. A., Smith R., (2012), arXiv:1204.3924
„Kinematical and chemical vertical structure of the Galactic thick disk II. A lack of dark matter in the solar neighborhood” (accepted in ApJ)