

# Partons and jets at strong coupling (III)

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# Outline

Outline

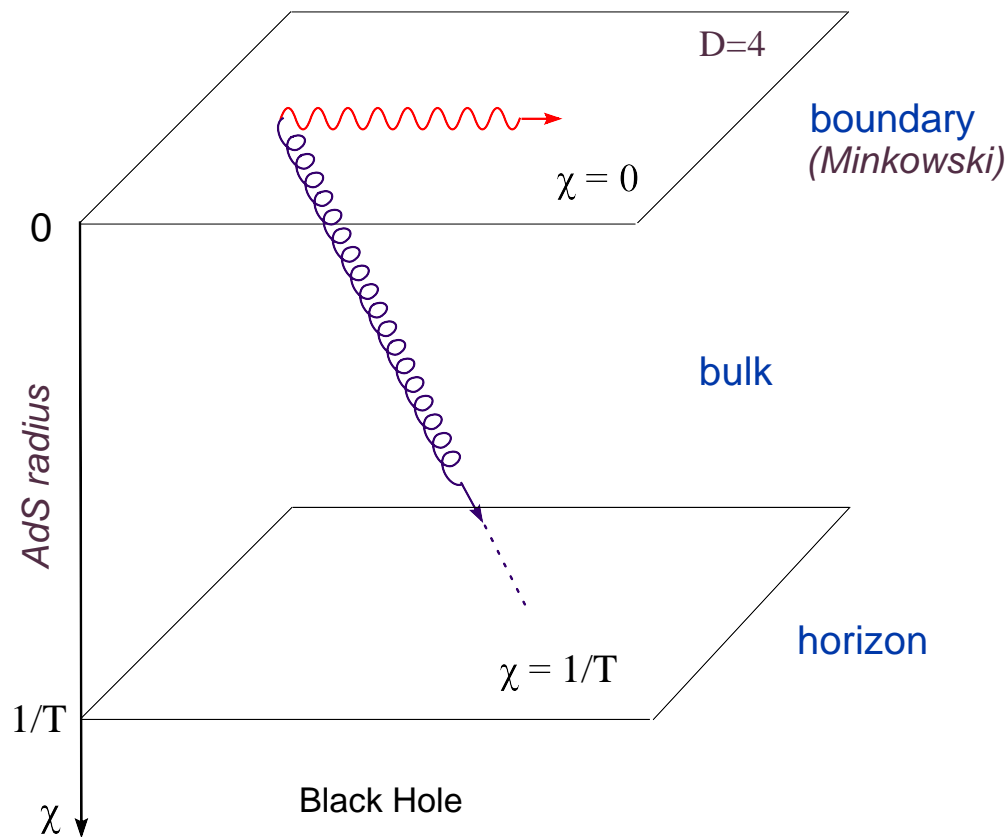
Current in the plasma

Conclusions

- Lecture I : Partons and jets in QCD at weak coupling
  - ◆ Introduction & Motivations
  - ◆ The situation at weak coupling (pQCD, phenomenology)
- Lecture II : A high–energy current in AdS/CFT
  - ◆ Invitation towards strong coupling
  - ◆ Methodology (black hole, wave equations)
  - ◆ The vacuum problem as a warm up
- Lecture III :  $\mathcal{R}$ –current in a strongly–coupled plasma
  - ◆ Results & Physical discussion
    - ▷ Medium induced branching
    - ▷ Parton saturation at strong coupling
    - ▷ Jet energy loss & momentum broadening
  - ◆ General consequences for high–energy scattering

# Maxwell wave in a Black Hole

- $\mathcal{N} = 4$  SYM at temperature  $T \iff$  a Black Hole in the 'radial' dimension of  $AdS_5$ : horizon at  $\chi_0 = 1/\pi T$  or  $\tilde{\chi}_0 = 2$
- Abelian current  $J_\mu$  in 4D  $\iff$  Maxwell wave falling into  $AdS_5$





# Electromagnetic current in a plasma

- Abelian  $\mathcal{R}$ -current :  $J_\mu(x) \propto e^{-i\omega t + ikz}$  with  $q^2 = \omega^2 - k^2$
- Retarded polarization tensor: thermal expectation value

$$\Pi_{\mu\nu}(q) \equiv \int d^4x e^{-iq \cdot x} i\theta(x_0) \langle [J_\mu(x), J_\nu(0)] \rangle_T$$

- 'Hard probe' :  $Q^2 \equiv |q^2| \gg T^2 \implies$  short distances
- High energy :  $\omega \simeq k \gg Q$  (most interesting)
- $\text{Im } \Pi_{\mu\nu}$  : absorption of the current by the plasma
  - ◆ time-like current ( $q^2 > 0$ ) : jets, meson
  - ◆ space-like current ( $q^2 < 0$ ) : DIS & parton structure
- 'The fall of the Maxwell wave in the Black Hole'
- The trajectory of this fall (+ the UV/IR correspondence)  
 $\implies$  the physical mechanism responsible for absorption

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# Wave equation in the AdS BH

$$\psi'' + \frac{1}{4\chi^2} \psi + \frac{\omega^2 - k^2 f}{f^2} \psi + \frac{f'}{f} \left( \psi' - \frac{1}{\chi} \psi \right) = 0$$

$$f = 1 - \left( \frac{\tilde{\chi}}{\tilde{\chi}_0} \right)^4 \quad \text{where} \quad \tilde{\chi}_0 = 2 \quad \text{or} \quad \chi_0 = 1/\pi T$$

- The most interesting dynamics at high  $Q^2$  takes place near the Minkowski boundary :  $\chi \lesssim 1/Q \ll \chi_0$

⇒ one can replace  $f \rightarrow 1$  everywhere except in the piece amplified by the energy :

$$\omega^2 - k^2 f = q^2 + k^2 \left( \frac{\chi}{\chi_0} \right)^4 = \mp Q^2 + \frac{k^2 \chi^4}{16}$$

upper sign: space-like ( $q^2 < 0$ ); lower sign: time-like ( $q^2 > 0$ )

- Competition between  $Q^2$  (virtuality) and  $k^2 \chi^4$  (interactions in the plasma) ⇒ critical point  $\chi_{\text{cr}} \sim \sqrt{Q/k} \ll \chi_0$

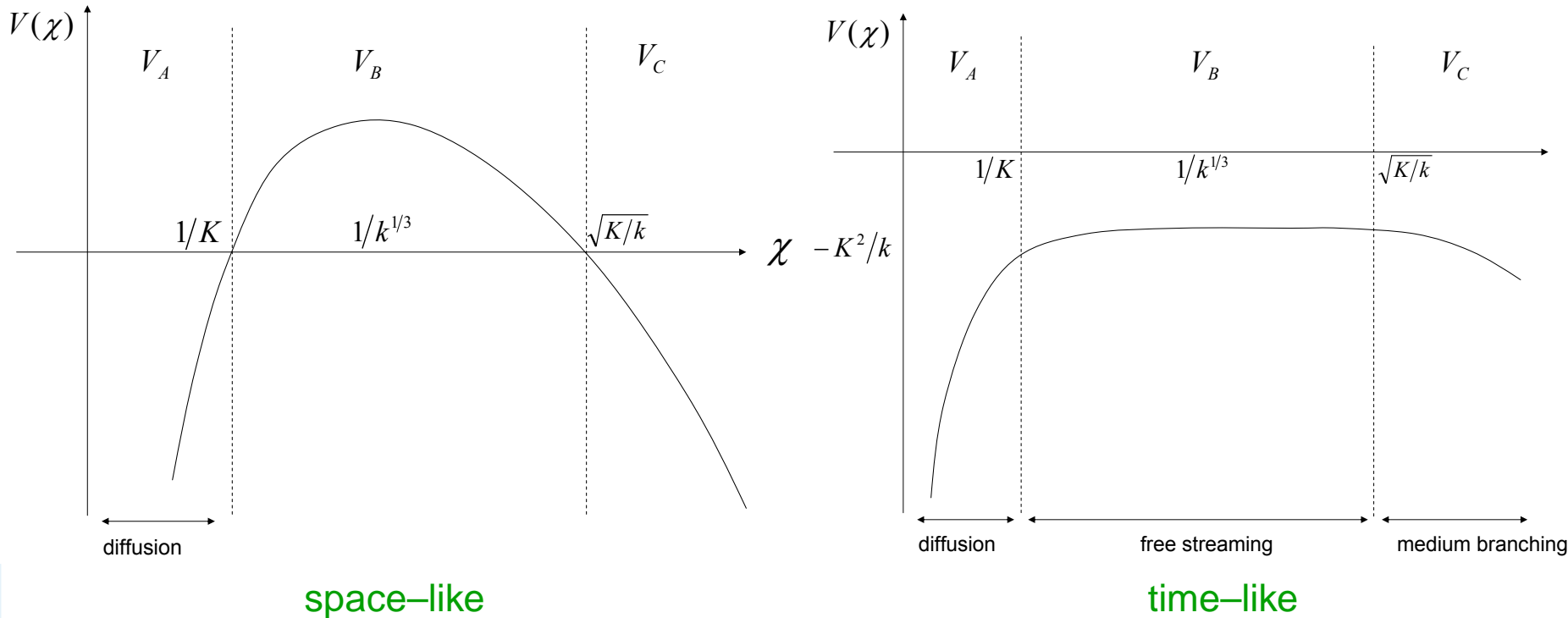
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■ Effective “Schrödinger equation” :  $-\psi'' + V_{\pm}(\chi)\psi = 0$



$(K \equiv Q \text{ in all the figures})$

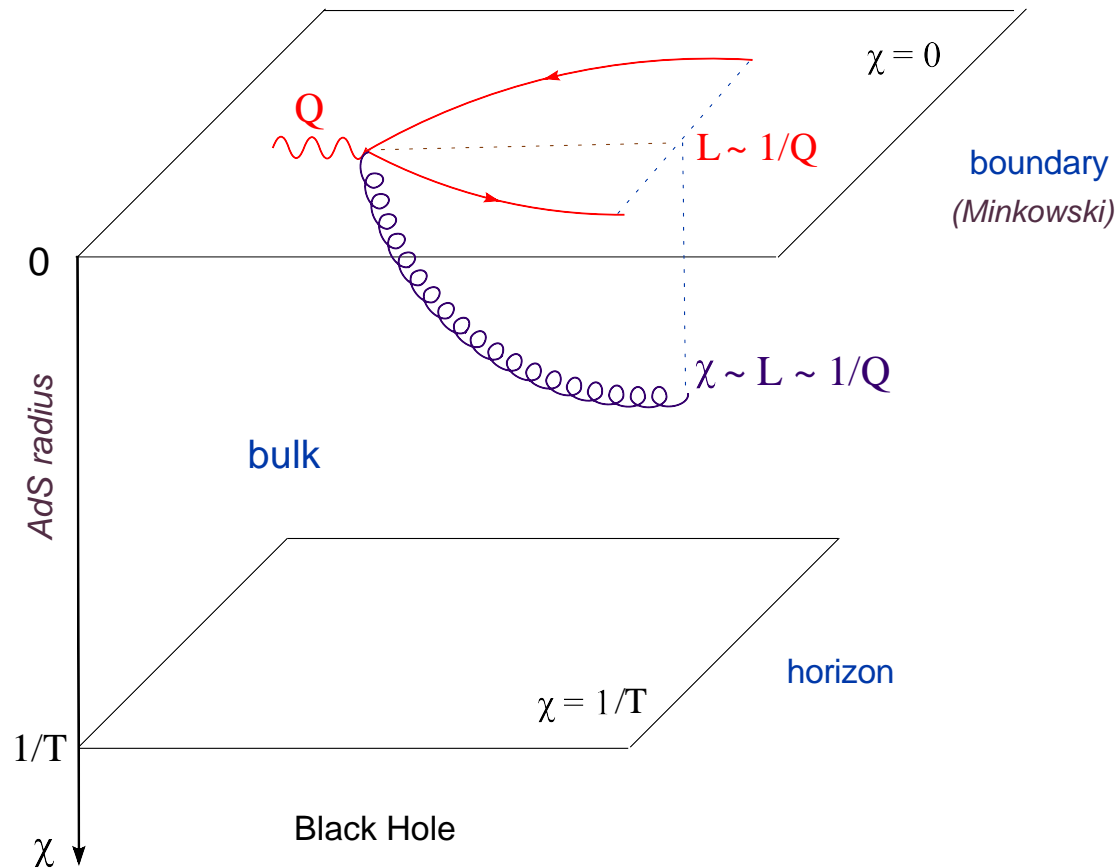
■ The potential becomes **attractive** at sufficiently **large**  $\chi$

◆ **small**  $\chi \ll \chi_{\text{cr}}$  : same dynamics as in the vacuum

◆ **larger**  $\chi$  : accelerated fall into the black hole

# Space-like current

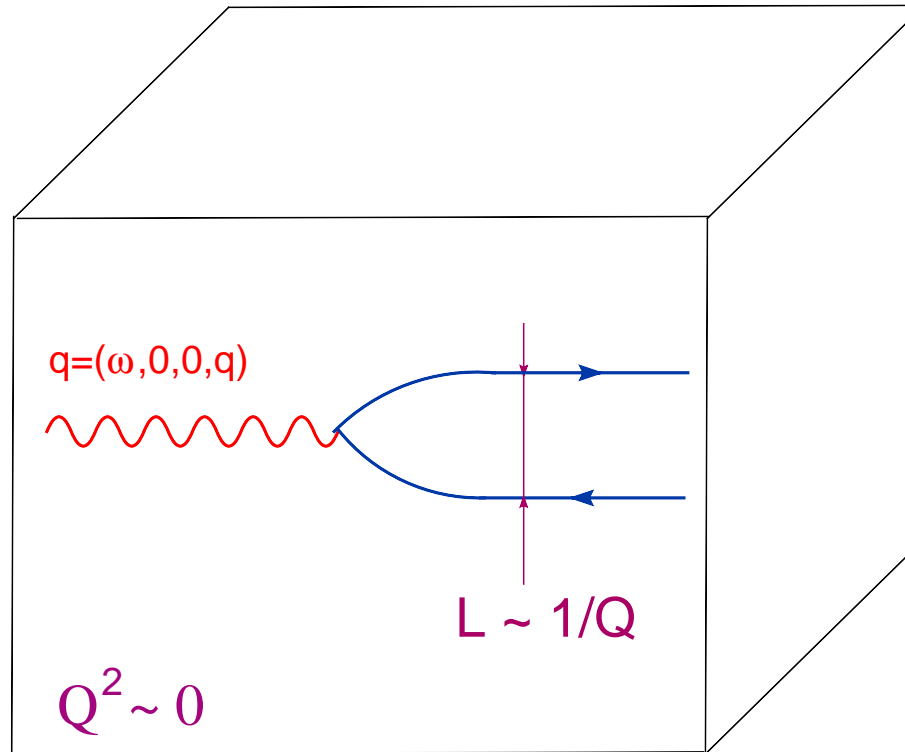
- After an early diffusion ( $L \sim \sqrt{t}$ ) at times  $t \lesssim \omega/Q^2$ , the wave gets stuck near the boundary:  $\chi \lesssim 1/Q \ll 1/T$



- No interaction with the BH ... except through tunneling

# 'Perfect color transparency'

- A small color dipole ('meson') with size  $L \sim 1/Q$  propagate through the strongly-coupled plasma **without interactions** !



- pQCD: the dipole cross-section vanishes too, but much more slowly :  $\sigma_{\text{dipole}} \sim 1/Q^2$  as  $Q^2 \rightarrow \infty$  (twist-2 operator)

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# Increasing the energy

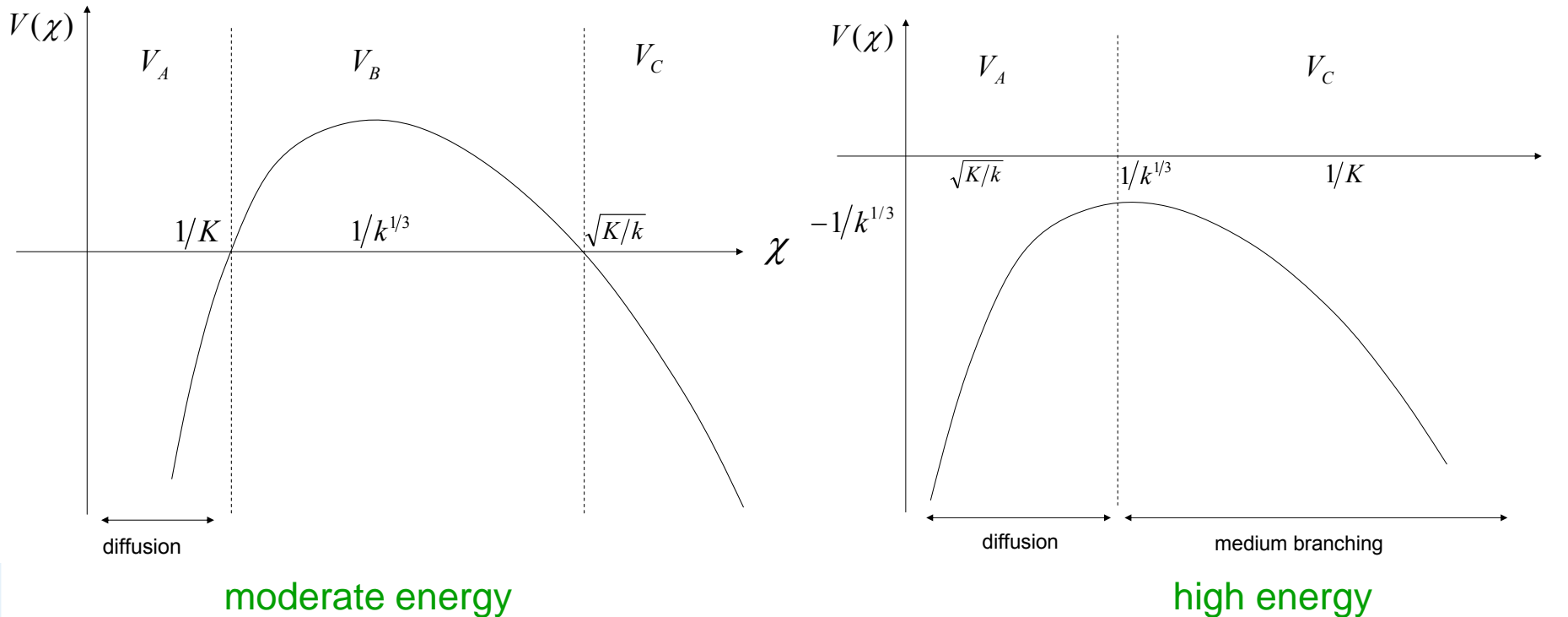
- By increasing the energy  $\omega$ , the interactions in the plasma become stronger and stronger

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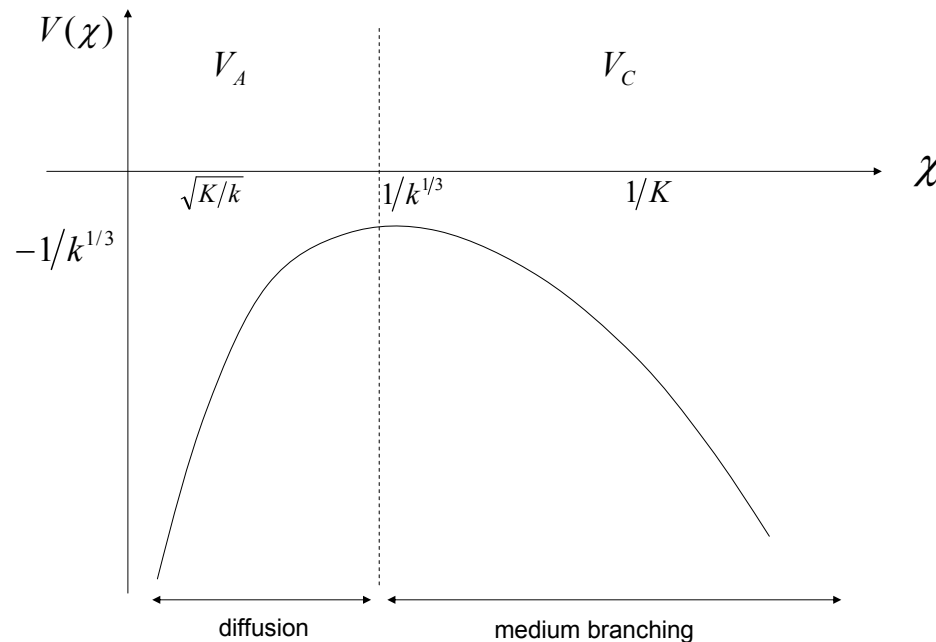
$(K \equiv Q \text{ in all the figures})$

- The barrier disappears when the penetration length  $\chi \sim 1/Q$  gets close to the 'point of no return'  $\chi_{\text{cr}} \sim \sqrt{Q/k}$

- This requires a **minimal energy**  $\omega_s$  (for given  $Q^2$  and  $T$ )

$$\frac{1}{Q} \sim \frac{1}{T} \sqrt{\frac{Q}{\omega}} \implies \omega_s \sim \frac{Q^3}{T^2}$$

- For  $\omega > \omega_s$ , the wave is **falling into the BH**  $\implies$  **energy loss**



- The virtuality plays no role  $\implies$  **the high-energy dynamics is the same for both space-like and time-like currents**



# Saturation momentum

$$\omega_s \sim \frac{Q^3}{T^2} \iff Q_s \simeq (\omega T^2)^{1/3} \simeq \frac{T}{x}$$

- Bjorken's  $x$  for a plasma (plasma rest frame) :  $x = \frac{Q^2}{2\omega T}$
- More on the physical meaning of this condition
- On the supergravity side:

$$\underbrace{Q}_{\text{potential barrier}} \gtrsim \underbrace{\frac{\omega T^2}{Q^2}}_{\text{gravitational potential}}$$

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# Saturation momentum

$$\omega_s \sim \frac{Q^3}{T^2} \iff Q_s \simeq (\omega T^2)^{1/3} \simeq \frac{T}{x}$$

## ■ On the gauge theory side

$$\underbrace{Q}_{\text{virtuality barrier}} \lesssim \underbrace{\frac{\omega}{Q^2}}_{\text{lifetime}} \times \underbrace{T^2}_{\text{plasma force}}$$

The partonic fluctuation must live long enough to feel the effects of the plasma

### Outline

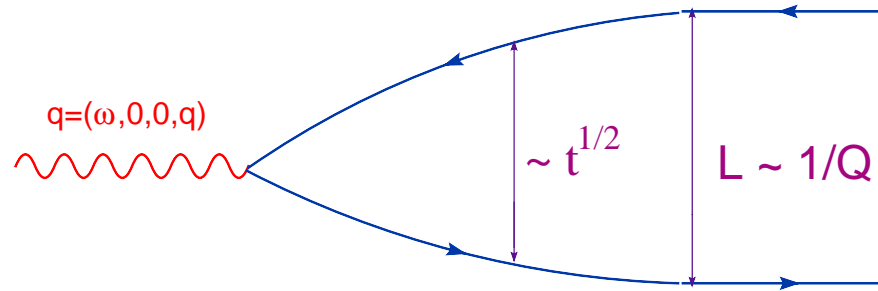
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$$\omega_s \sim \frac{Q^3}{T^2} \iff Q_s \simeq (\omega T^2)^{1/3} \simeq \frac{T}{x}$$

- $L_s \sim 1/Q_s$  : the minimal size for the dipole to feel the plasma



- Dipoles with size  $L \gtrsim L_s$  cannot survive in the plasma.

$$L_s \sim \frac{1}{Q_s} \quad \& \quad \gamma \sim \frac{\omega}{Q} \implies L_s \sim \frac{1}{\sqrt{\gamma}T} = \frac{(1 - v_z^2)^{1/4}}{T}$$

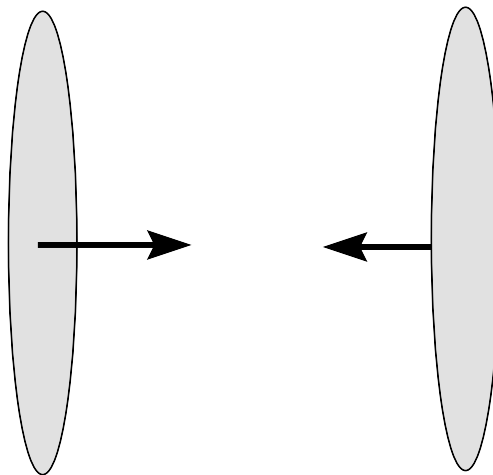
“meson screening length”, “limiting velocity”

[cf. Liu, Rajagopal, Wiedemann; Chernicoff et al; Caceres et al (2006)]

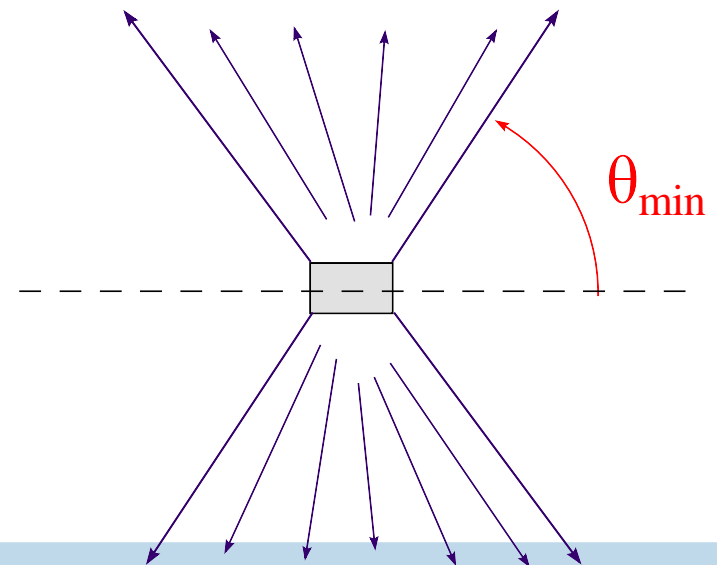
# Large- $x$ : No partons, no jets !

$$Q_s(x) \simeq \frac{T}{x} \iff x_s(Q) \simeq \frac{T}{Q}$$

- For  $Q > Q_s(x)$  or  $x > x_s(Q)$  :  $F_2(x, Q^2) \approx 0$   
 $\implies$  no partons with large momentum fractions  $x > x_s$
- ‘All partons have branched down to small values of  $x$ ’  
*(Polchinski and Strassler, 02; Hatta, E.I., Mueller, 07; see below !)*
- No forward/backward jets in hadron-hadron collisions !



$t < 0$



$\theta_{\min}$

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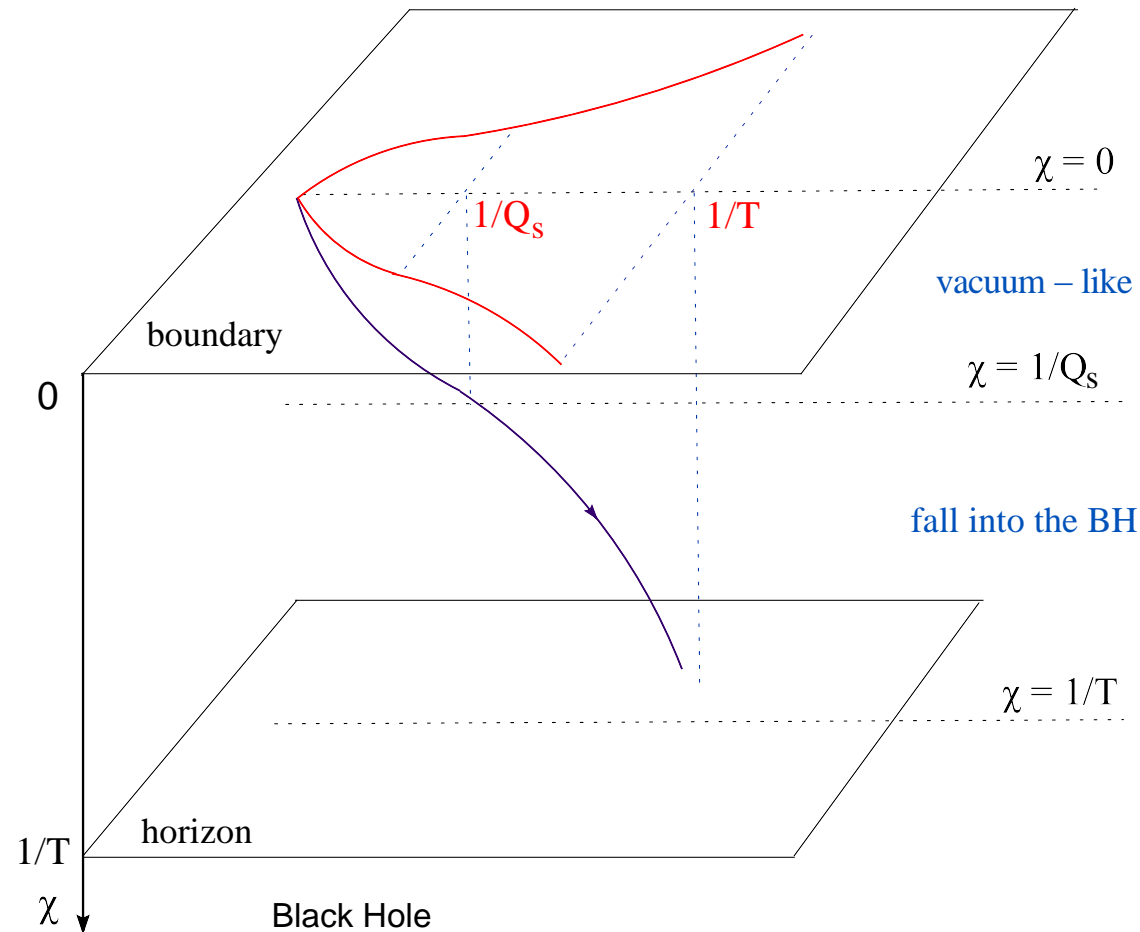
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# High energy: The fall

- For  $Q > Q_s(x)$  or  $x > x_s(Q)$  : the wave diffuses up to  $\chi \sim 1/Q_s$ , then it falls along a **massless geodesics**



- The same as the 'trailing string' (*Herzog et al; Gubser, 06*)

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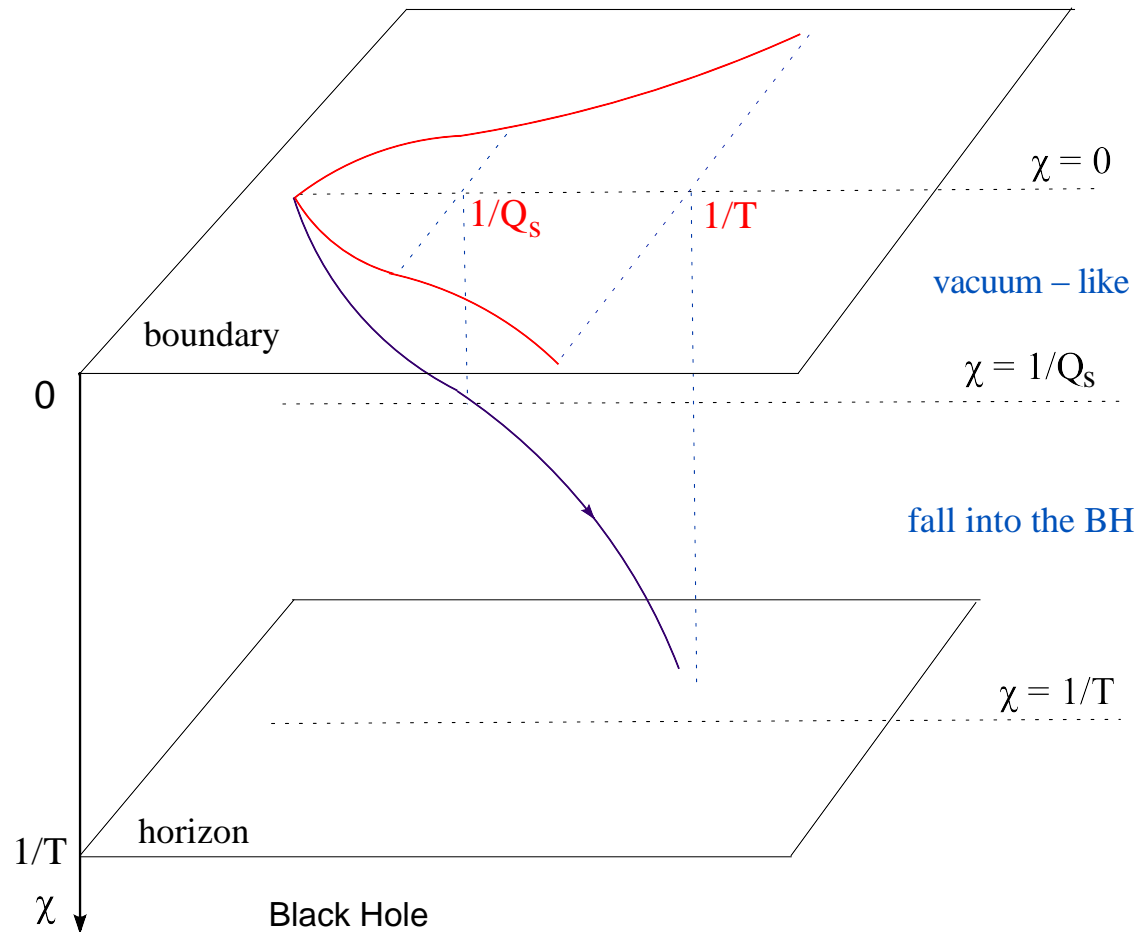
### ● Trailing string

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# The lifetime of a high-energy current

Lifetime of the current: 
$$\Delta t \sim \frac{\omega}{Q_s^2} \sim \frac{1}{T} \left( \frac{\omega}{T} \right)^{1/3}$$



■ Same as the ‘gluon lifetime’ (*Gubser, Gulotta, Pufu, and Rocha,08*)

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# Finding the partons (plasma IMF)

- Total absorption of the current in the plasma:

Unitarity limit (maximal possible scattering) for DIS

- Saturation line:  $Q_s(x) \simeq T/x$  or  $x_s(Q) \simeq TQ$

$$F_2(x, Q^2) \simeq xN_c^2 Q^2 \begin{cases} \exp \{ - (x/x_s)^{1/2} \} & \text{for } x \gg x_s \equiv T/Q \\ 1 & \text{for } x \lesssim x_s \equiv T/Q, \end{cases}$$

- For given  $Q \gg T$ , all the partons have fallen at  $x \lesssim x_s \ll 1$
- For  $x \lesssim x_s$  : occupation numbers of  $\mathcal{O}(1) \implies$  **saturation**
- Energy sum rule:

$$\mathcal{E} = T^2 \int_0^1 dx F_2(x, Q^2) \sim T^2 \left[ xF_2(x, Q^2) \right]_{x=x_s} \sim N_c^2 T^4 \quad \checkmark$$

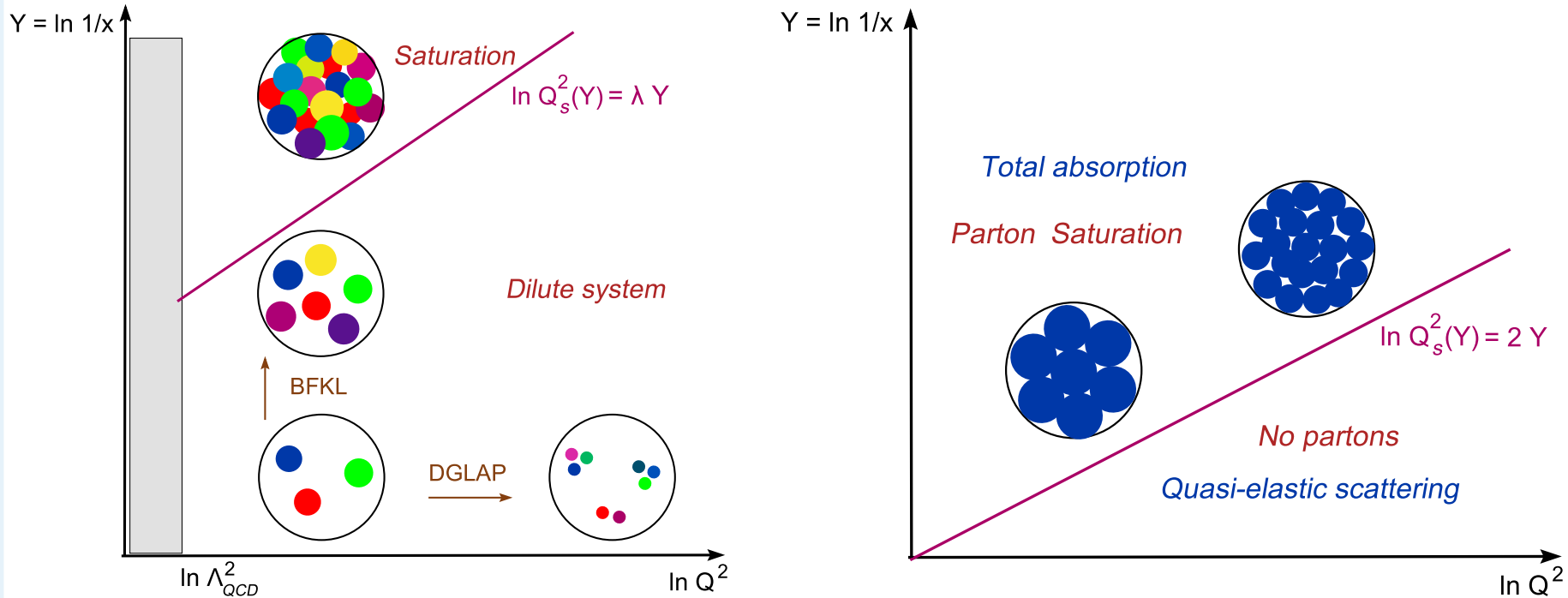
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# Saturation line: weak vs. strong coupling



■ Saturation exponent :  $Q_s^2(x) \propto 1/x^{\lambda_s} \equiv e^{\lambda_s Y}$

◆ weak coupling (lowest order):  $\lambda_s \approx 1.23 g^2 N_c$

◆ phenomenology & pQCD NLO:  $\lambda_s \approx 0.3$

◆ strong coupling (plasma):  $\lambda_s = 2$  (graviton)

# Time-like current at strong coupling

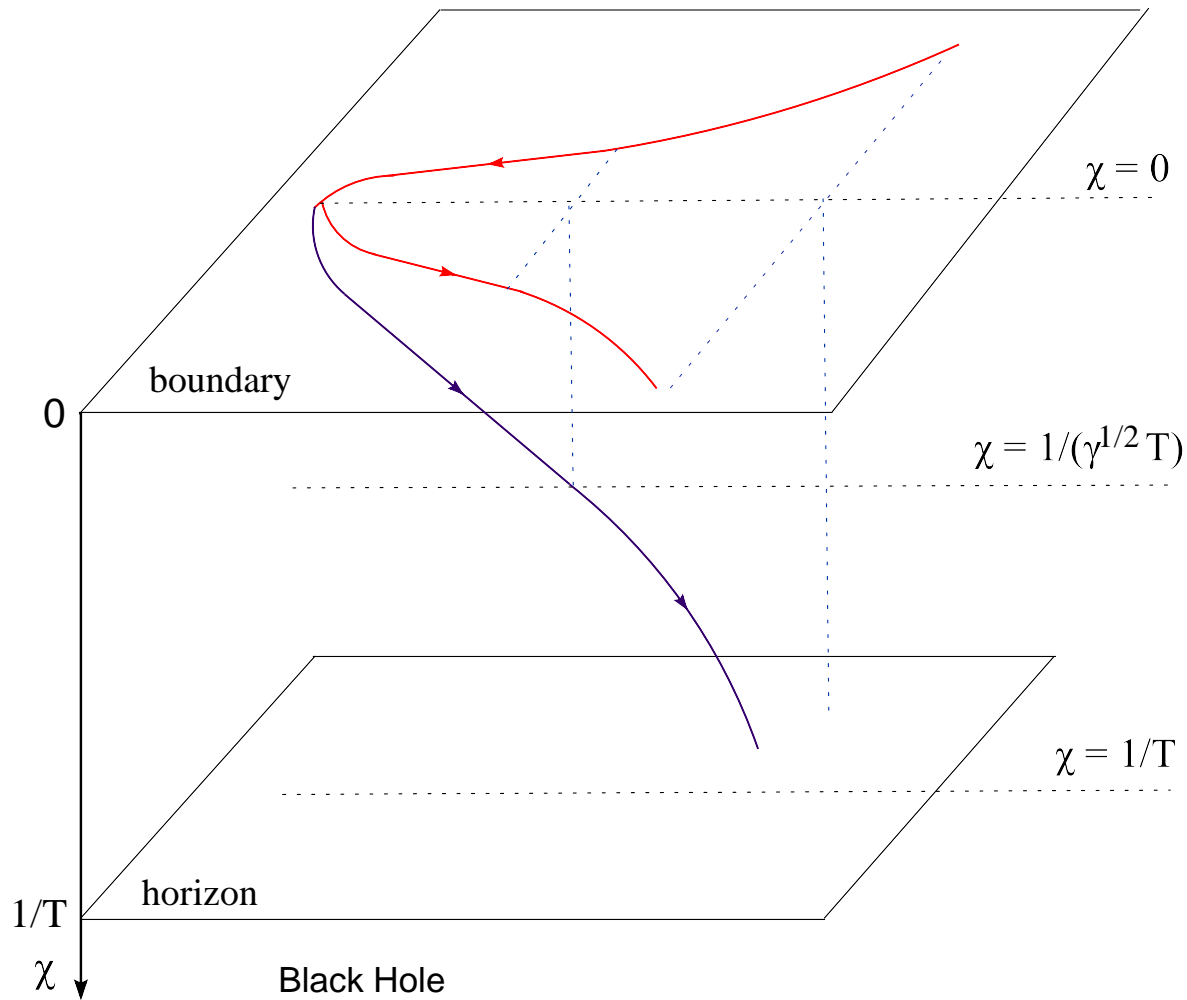
- A time-like current can decay (into partons of  $\mathcal{N} = 4$  SYM) already in the **vacuum**

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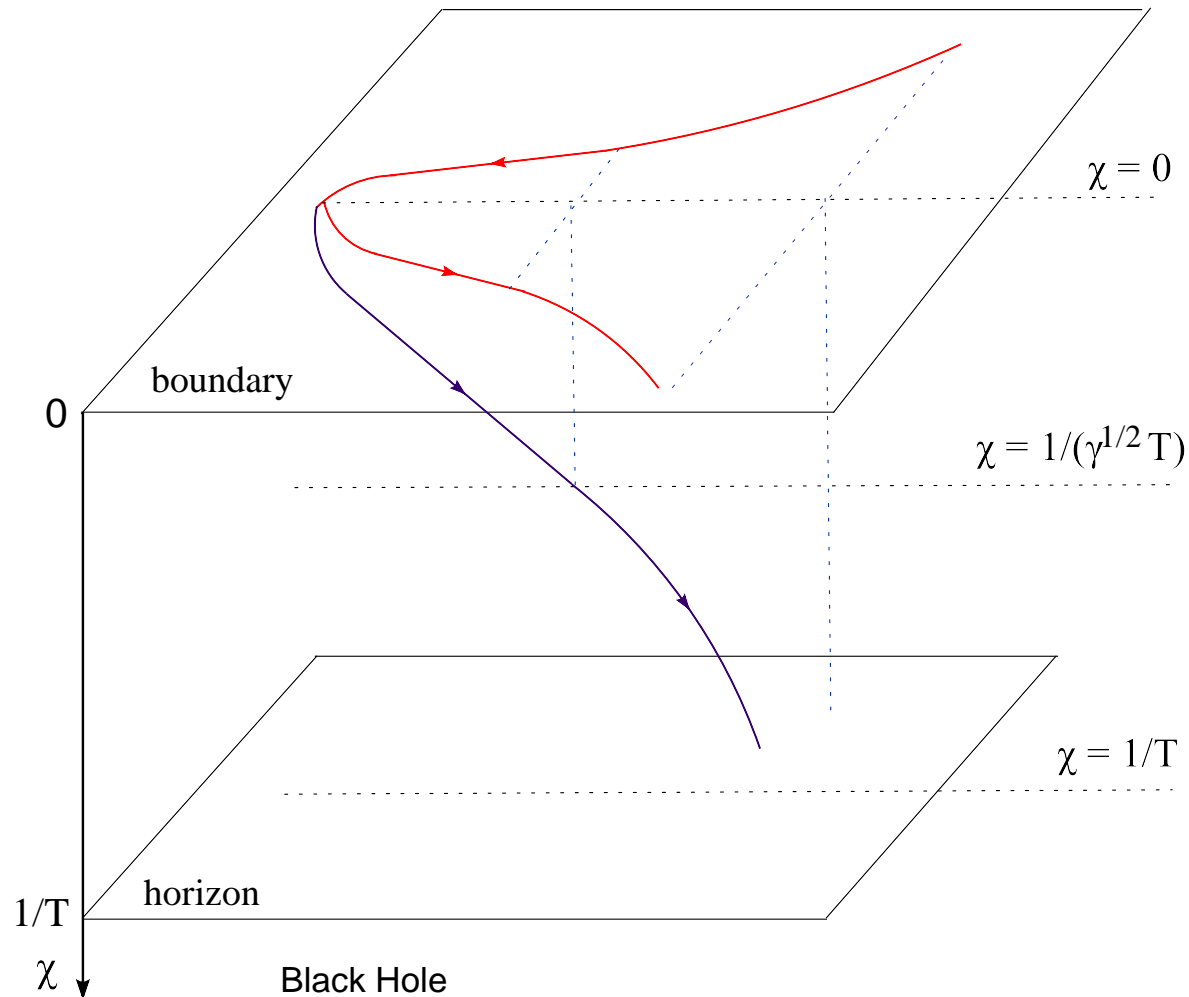
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# Early times: free streaming

## ■ Early times/small size $L$ :

free expansion up to the critical size  $L_{cr} \sim 1/\sqrt{\gamma} T$



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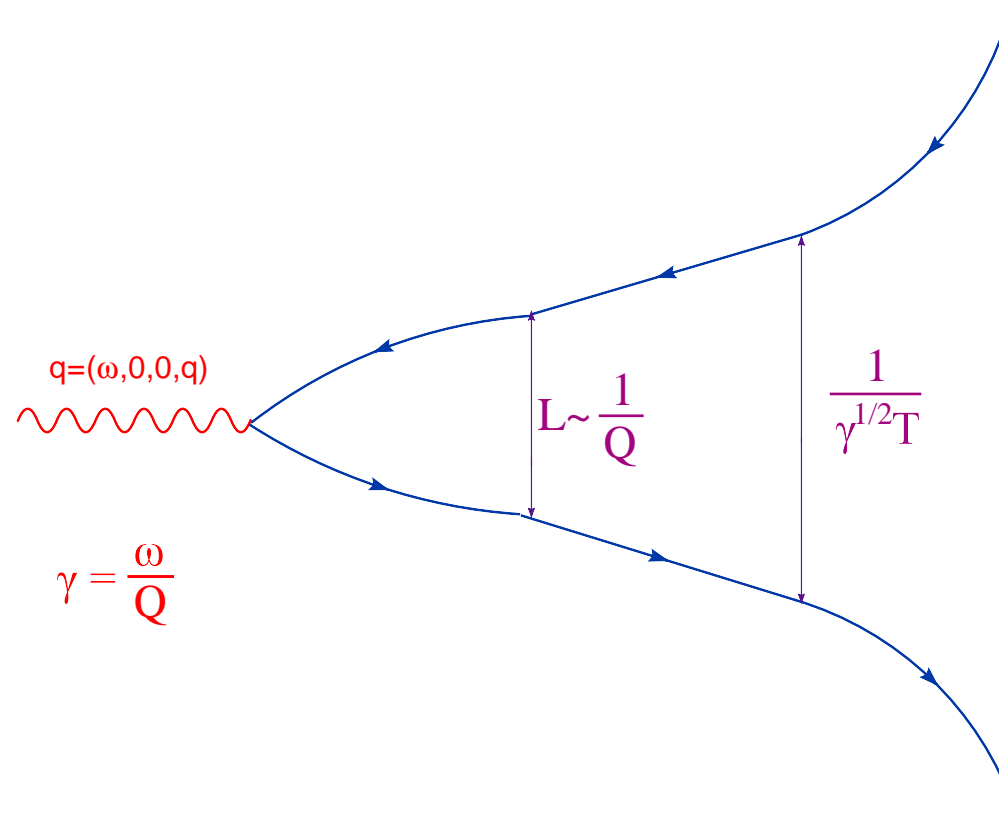
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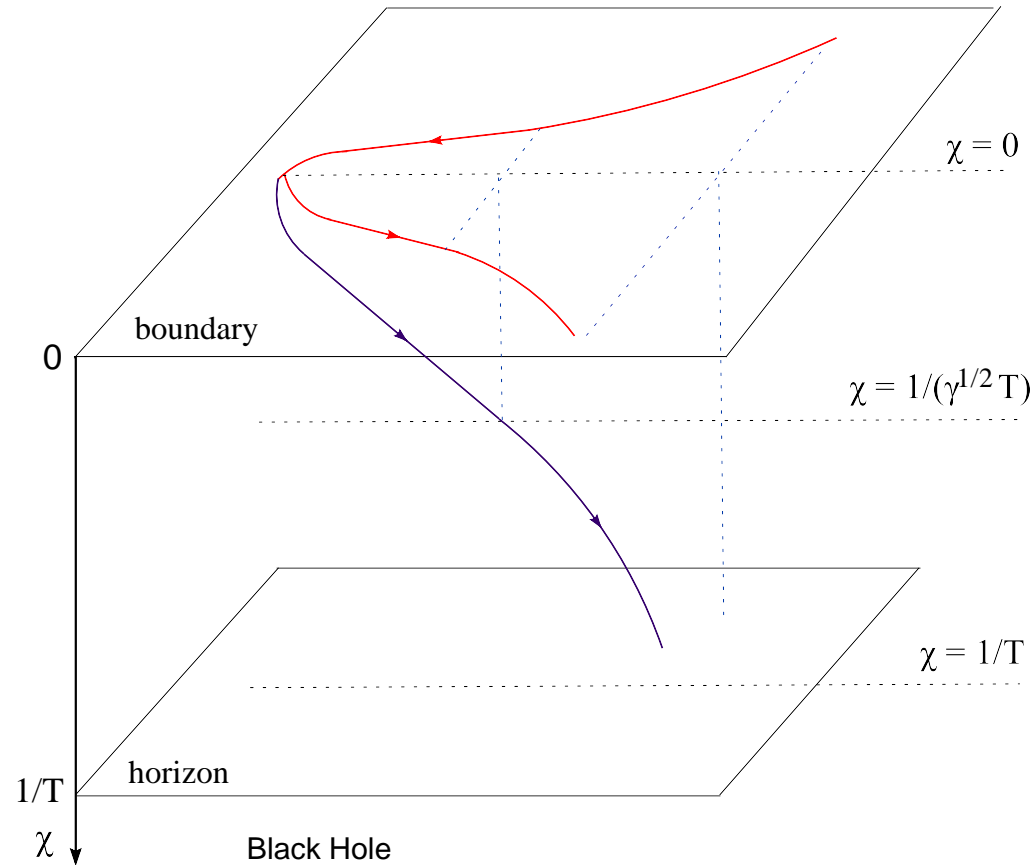


- Free streaming up to the point of no return:

$$\chi_{\text{cr}} \sim \frac{1}{\sqrt{\gamma}} \implies L_{\text{cr}} \sim \frac{1}{\sqrt{\gamma} T} = \frac{(1 - v_z^2)^{1/4}}{T}$$

# Later times : falling into the black hole

- 'trailing string' [*Herzog, Karch, Kovtun, Kozcaz, Yaffe; Gubser, 2006*]



- Energy transfer from the partons to the plasma

Lifetime of the current/penetration length:  $\Delta t \sim \Delta z \sim \frac{\sqrt{\gamma}}{T}$

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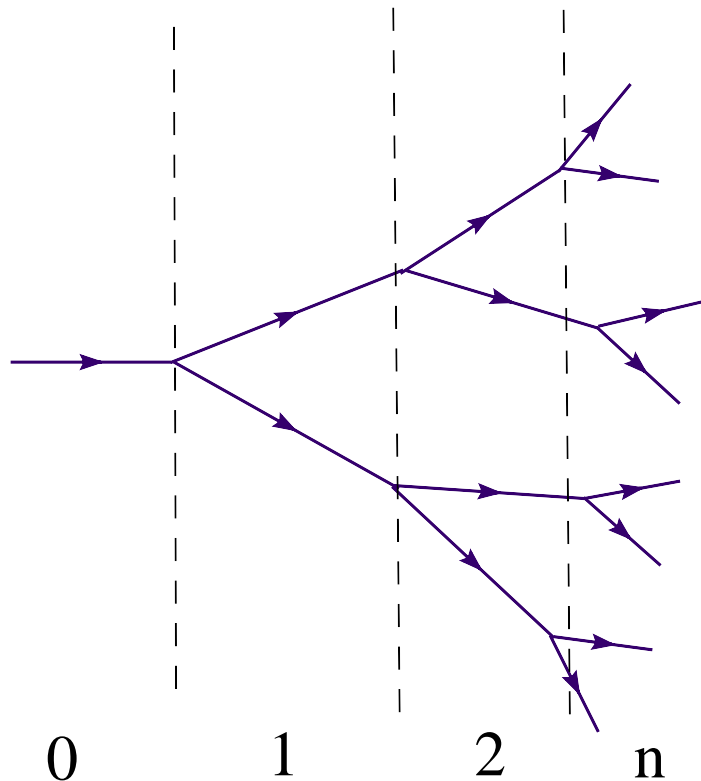
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# Medium induced branching

## ■ Universal energy loss mechanism, active at partonic level

- ◆ no reason why branching should stop at 2 parton level
- ◆ no reason to favour special corners of phase-space



$$\omega_n \sim \frac{\omega_{n-1}}{2} \sim \frac{\omega}{2^n}$$

$$Q_n \sim \frac{Q_{n-1}}{2} \quad (\text{vacuum})$$

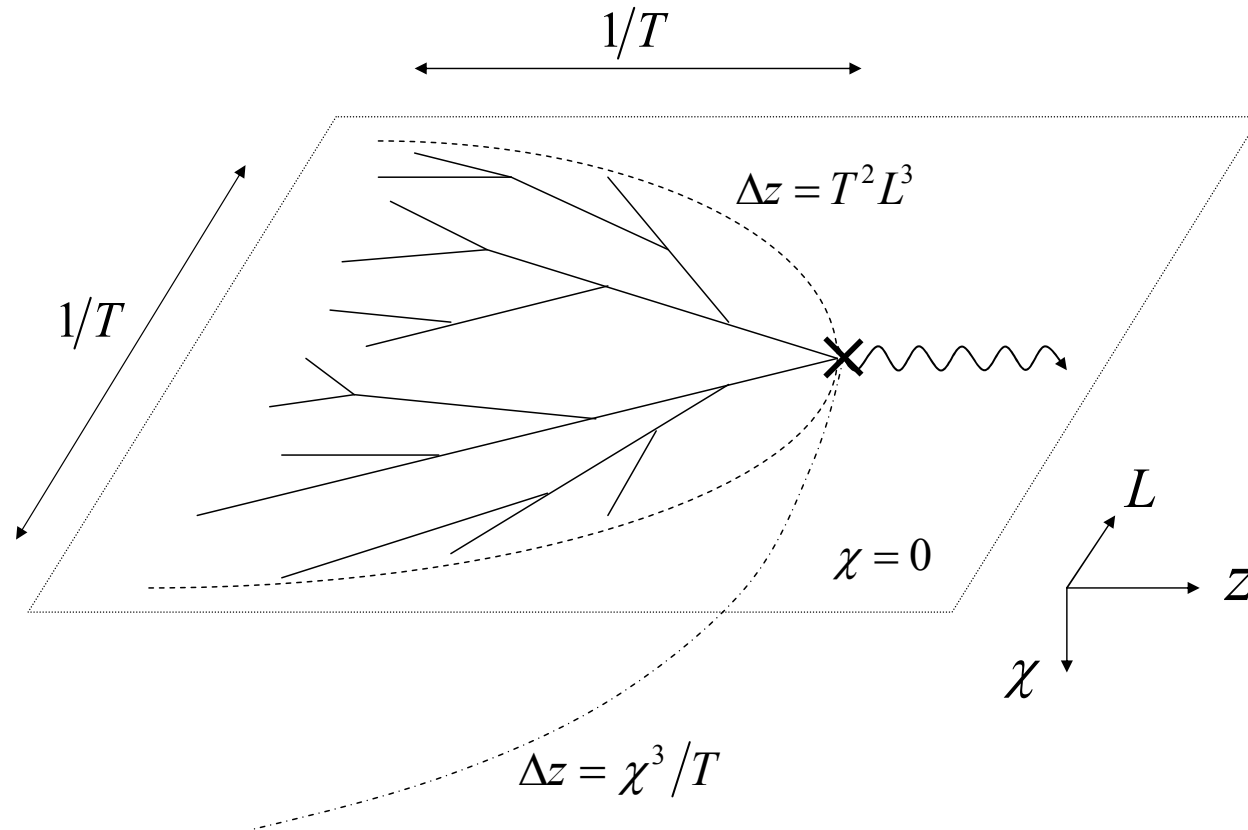
$$\Delta t_n \sim \frac{\omega_n}{Q_n^2}$$

$$\frac{\Delta Q_n}{\Delta t_n} \sim -T^2 \quad (\text{plasma})$$

$\Rightarrow Q_n \sim (\omega_n T^2)^{1/3}$  : saturation momentum at step  $n$

# Trailing string revisited

- The **enveloping curve** of the resulting partonic system coincides with the **'trailing string'** (*Herzog et al; Gubser, 2006*)



- ... as it should by virtue of the UV/IR correspondence

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# Energy loss: massless parton

- Energy loss per branching:

$$\frac{\omega_n - \omega_{n-1}}{t_n - t_{n-1}} \sim - \frac{\omega_n}{\omega_n / Q_n^2} \sim - Q_n^2$$

$$\Rightarrow - \frac{d\omega(t)}{dt} \simeq Q_s^2(t) = (\omega T^2)^{2/3}$$

- Parton lifetime (the time to loose all its energy) :

$$\Delta t \sim \frac{1}{T} \left( \frac{\omega_0}{T} \right)^{1/3}$$

- Same as the 'gluon lifetime' (*Gubser, Gulotta, Pufu, and Rocha,08*)

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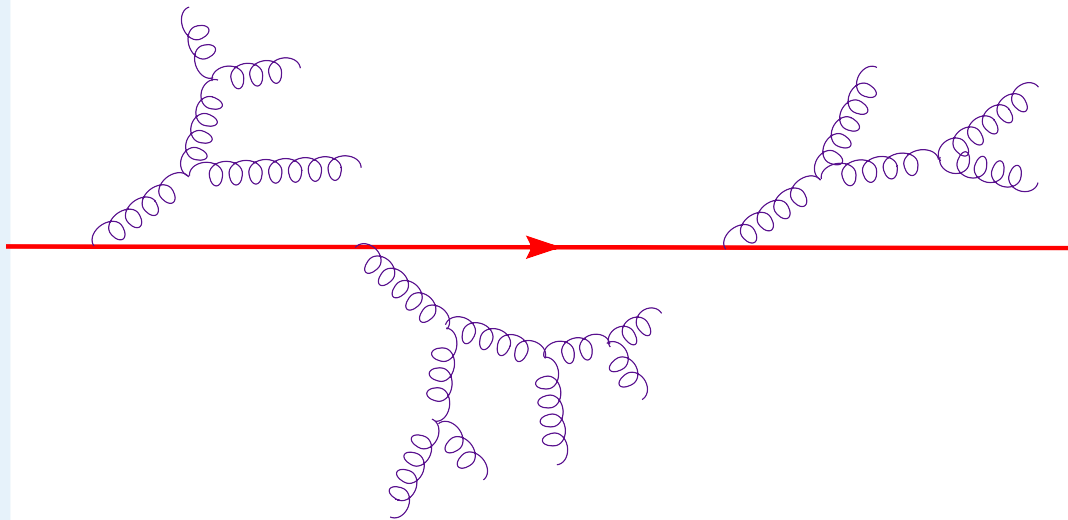
# Heavy Quark: Energy loss

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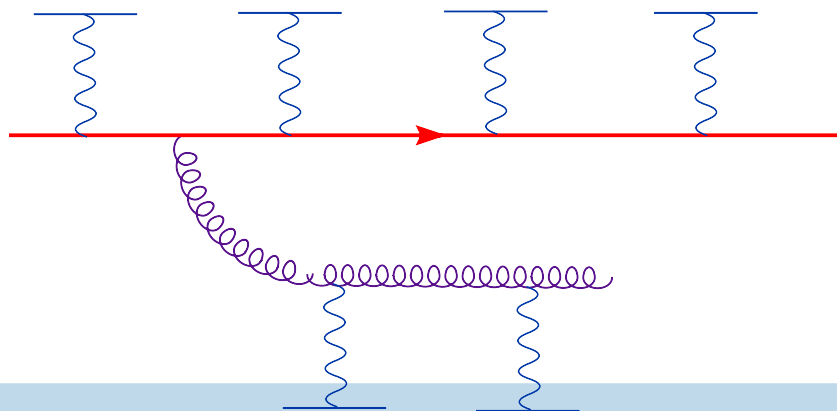
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$$Q_s^2 \sim (t_{\text{br}} T^2)^2 \sim \gamma T^2$$

$$-\frac{dE}{dt} \simeq \sqrt{\lambda} \frac{\omega}{(\omega/Q_s^2)} \simeq \sqrt{\lambda} Q_s^2 \quad (\text{Herzog et al; Gubser, 2006})$$

■ **pQCD** : same formula, but with  $\sqrt{\lambda} \rightarrow g^2 N_c$  and different  $Q_s$



$$Q_s^2 \sim t_{\text{br}} \hat{q} \quad (\text{quenching param.})$$

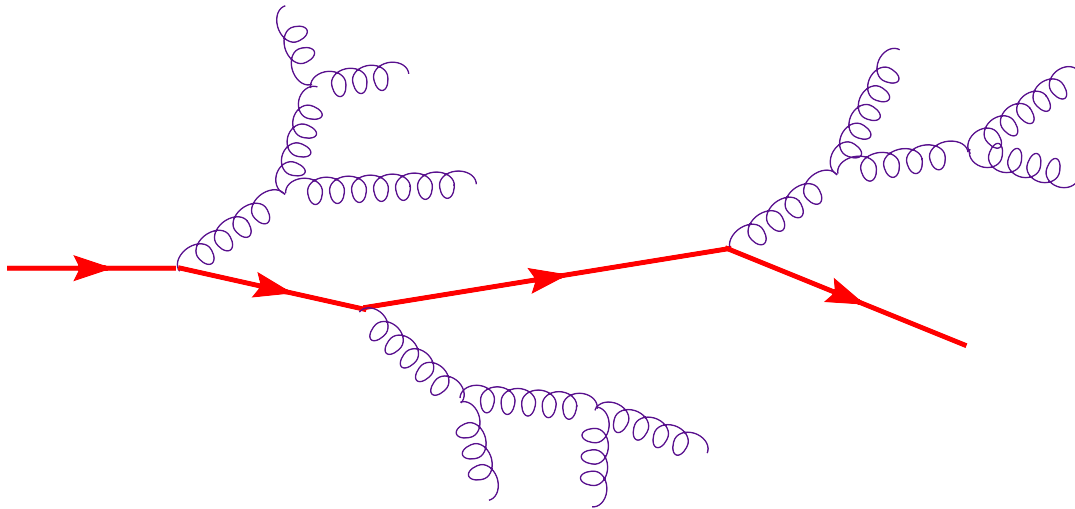
# Transverse momentum broadening

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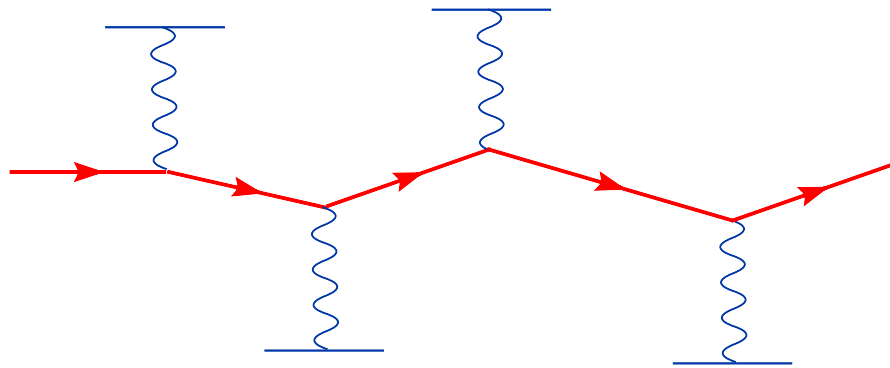
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$$\frac{dp_{\perp}^2}{dt} \sim \sqrt{\lambda} T^2 Q_s$$

*Casalderrey-Solana, Teaney, 2006; Gubser, 2006; Dominguez et al, 2008*

■ pQCD : different physics ! thermal rescattering



$$\frac{dp_{\perp}^2}{dt} \sim \hat{q}$$

- **Hard probes & high-energy physics** appears to be quite different at strong coupling as compared to QCD
  - ◆ no forward/backward particle production in HIC
  - ◆ no jets in  $e^+e^-$  annihilation
  - ◆ jet momentum broadening dominated by in-medium parton branching, as opposed to thermal rescattering
  
- Not necessarily surprising: by **asymptotic freedom**, hard & high-energy physics in QCD is weakly coupled
  
- Are AdS/CFT methods useless for HIC ? **Not necessarily so !**
  - ◆ Transition from partons to fluid (*Gubser et al, Chesler and Yaffe*)
  - ◆ Some observables receive contributions from several scales, from soft to hard: use **AdS/CFT** in the soft sector  
*Jet quenching: Liu, Rajagopal, Wiedemann, 06; Mueller, 08*
  - ◆ Separation scale  $Q_0 = ??$  Perhaps  $2\pi T_c \sim 1 \text{ GeV}$
  - ◆ ... and long-range properties: hydro, thermalization etc