

The Selforganized Universe

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What is QG?

Dyson: Whenever I try to imagine a gedanken experiment which probes the quantum nature of gravity I end up with a black hole.

String theory offers (in my opinion) disappointingly little insight in QG (if it exists)

How does $\langle g_{\mu\nu} \rangle$ emerge from nothing and how do we study the quantum fluctuations around E_{sp} ?

(This will be the topic of my lectures!)

Conformal inv. (symmetry of ST) (2)

tells us around which $\langle g_{\mu\nu}(x) \rangle$ we can have a consistent ST

$$S = \int d^2_3 \sqrt{h} h^{\alpha\beta} \langle g_{\mu\nu}(x) \rangle \partial_\alpha X^\mu \partial_\beta X^\nu$$

But SFT, the "meta" theory which determines the dynamics of $\langle g_{\mu\nu}(x) \rangle +$ fluctuations is not really developed.

This is the old program of 86-89:
The natural space of SFT: space of 2d field theories, $\langle g_{\mu\nu}(x) \rangle$ being the extremum and our universe the natural minimum

Many reasons this program did not succeed.

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The second string revolution did not really improve the situation from p.o.f of $\mathbb{Q}G$: An even larger zoo of choices.

In fact it seems that more and more string people give up and appeal to the anthropic principle for $\langle g_{\text{max}} \rangle$

LHC is an interesting test. (SUSY)

For me: SUSY = Take ST serious

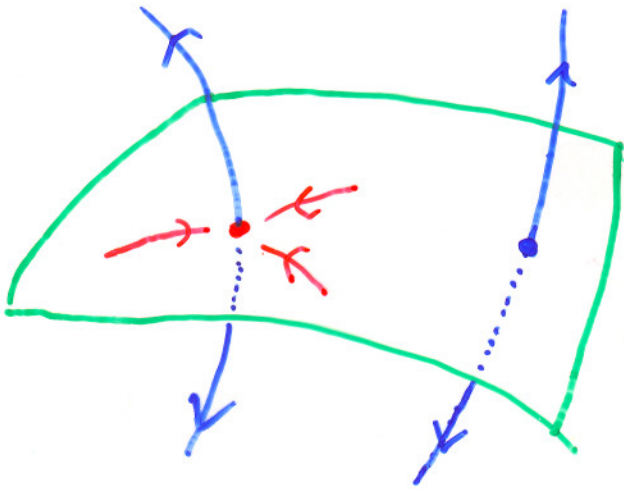
No SUSY: ST \equiv epicycles

Awaiting these results (and anticipating that SUSY will not be obs.) we have an obligation to look elsewhere. Maybe simpler selection?

Asymptotic safety

(Weinberg 79)

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Wilsonian picture:
critical surface of
finite co-dimension

In principle it provides us with a
non-perturbative def. of a QFT:

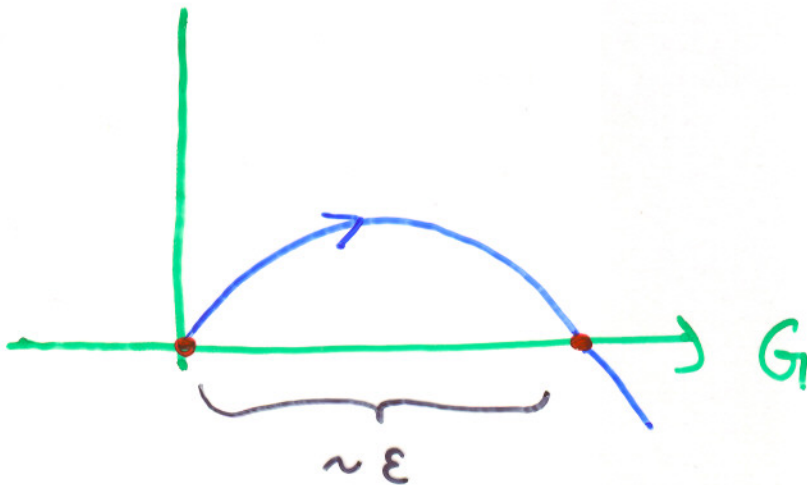
①: Find a fixed point

②: prove CS has finite co-dimension

Success: Fisher-Wilson fixed point in 3d

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Weinberg used $2+\epsilon$ d expansion to argue that there exists a non-trivial fixed point for QG in $d=4$ ($\epsilon=2$)



Elaborated on by Kawai et al (early 90s)

Reuter et al, Litern, last 10 years:

Exact renormalization group

Some evidence

Problem: truncation of effective action.

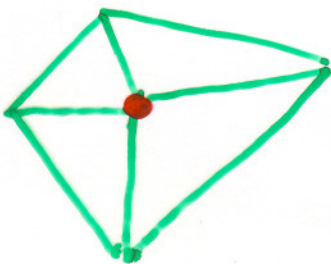
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Another approach, in the same spirit, comes from performing an explicit lattice-regularization.

Several Lattice approaches. Here only **Dynamical triangulation (DT)**.

How to put a diff. inv. theory on lattice?
realize that diff. inv. is the desire to deal directly with geometry.

Piecewise linear geometry (PLG) can be described without the use of coordinates:

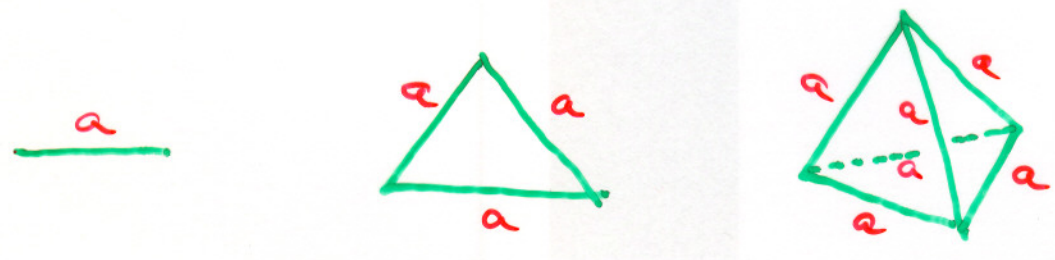


2d: just provide length of sides and a table of neighboring triangles.

curvature lives on vertices

This can be generalized to higher dimensions : curvature lives on $d-2$ dimensional hinges.

We will make it even simpler by restricting ourselves to simple Building Blocks (BB)

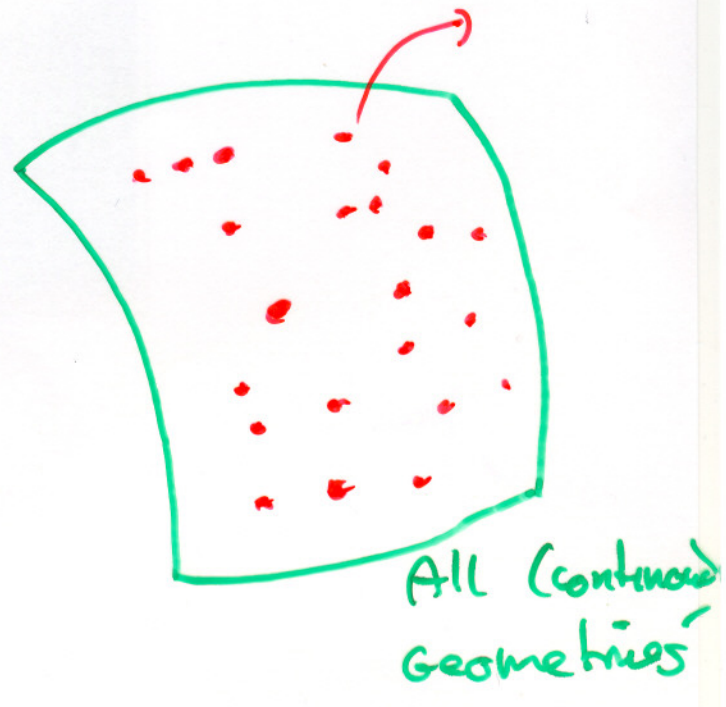


a = Lattice spacing

PLG-BB

Does it work?

PLG-BB geometries a small subset of all geometries.



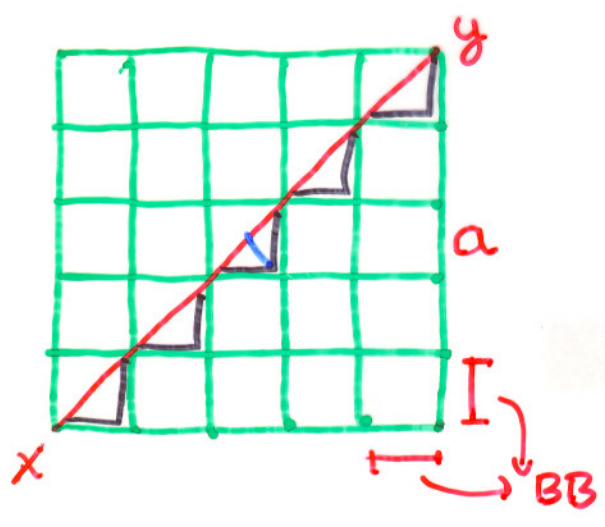
$$\sum_{\text{PLG}(a \rightarrow 0)} \longrightarrow \int \mathcal{D}[g_{\mu\nu}] \quad (?)$$

Clear that we can no longer approx.
a given (say) smooth geometry in an
obvious way with BB

Ex Free particle in D Eucl. dimensions
with geometric action:

$$S(P) = m \int_P d\ell = m L(P)$$

Hypercubic lattice: BB !



$$S(P_L) \geq \sqrt{2} S(P_{cont})$$

In what sense is the lattice P_L close to P_{cont} for $a \rightarrow 0$?

$$d(P_{cont}, P_L) = \max \{ d(x, P_{cont}) \mid x \in P_L \}$$

$$d(x, P_{cont}) = \min \{ d(x, y) \mid y \in P_{cont} \}$$

This defines a distance measure on the space of paths from x to y .

One can easily show that the BB-approach works for a free particle, but a more serious test is Euclidean 2d QG.

$$Z(\Lambda, G) = \int \mathcal{D}[g_{\mu\nu}] e^{-S[g_{\mu\nu}]}$$

$$S[g] = \Lambda \int d^2x \sqrt{g} - \frac{1}{G} \int d^2x \sqrt{g} R$$

As long as no topology change: $\int d^2x \sqrt{g} R = \text{const.}$

$$Z(\Lambda) = \int_0^\infty dV e^{-\Lambda V} \cdot \int \mathcal{D}[g] \cdot "1"$$

$V_g = V$
 $V_g = \int d^2x \sqrt{g}$

$$Z(\Lambda) = \int_0^\infty dV e^{-\Lambda V} \mathcal{N}([g] \sim V)$$

Entirely combinatorial

Our BB-approach is ideal for counting $\mathcal{N}(\Gamma_g) = V$

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$$Z(\Lambda) = \sum_T e^{-\Lambda V} = \sum_T e^{-\mu N_T}$$

$$V = \frac{\sqrt{3}}{4} a^2 N_T, \quad \Lambda a^2 = \mu$$

$$\begin{aligned} Z(\mu) &= \sum_N e^{-\mu N} \sum_{T \sim N} 1 \\ &= \sum_N e^{-\mu N} \mathcal{N}(T \sim N) \end{aligned}$$

This counting can be done combinatorially, (Tutte) or using matrix models (Kazakov, David, ...)

Results agree with continuum, "standard" calculations of $\int \mathcal{D}[\Gamma_m] e^{-S[\Gamma_m]}$, also for more complicated observables:

Conclusion: Lattice approach works in 2d

2d: No dynamical graviton

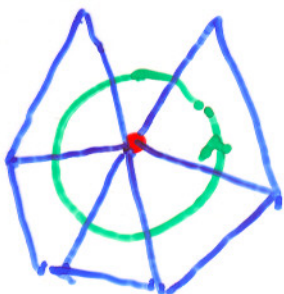
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Let us jump to 4d:

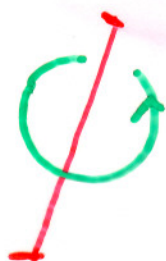
BB: Four-simplices of length a .
glue together to form space of
fixed topology (S^4 , say)
in all possible ways.

What action to choose?

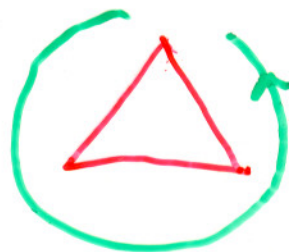
There is a natural geometric action
on PLGs, first discussed by Regge,
expressing the curvature, living on the
 $d-2$ -dim. hinges as defect angles.



2d



3d



4d

Using identical BB the Regge action becomes trivial:

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$$S(T) = -C_1 N_2(T) + C_2 N_4(T) = -k_0 N_0(T) + k_4 N_4(T)$$

$$Z(\Lambda, G) = \int \mathcal{D}[T] e^{-S[T]}$$

$$Z(k_0, k_4) = \sum_T e^{-S(T)}$$

$$= \sum_{N_0, N_4} e^{k_0 N_0 - k_4 N_4} \sum_{T \sim N_0, N_4} 1$$

$$= \sum_{N_0, N_4} x^{N_0} y^{N_4} \mathcal{N}(T \sim N_0, N_4)$$

$$x \sim e^{k_0}, y \sim e^{-k_4} :$$

$Z(x, y)$ is the generating function for the numbers $\mathcal{N}(N_0, N_4)$

Unfortunately no analytic solution to this counting problem.

Fortunately it is possible to count approximately using MC-simulations

All the results I report about will be based on such numerical "experiments"

After dinner tonight Andrzej Görlich will give a wonderful talk with nice "experimental" pictures and tell you how such experiments are actually conducted.

In 2d case it is well tested that MC-simulations can produce the critical exponents of 2d QG.

In the first half of 90 the 5-6 groups carried out such simulation:

Conclusion

It almost worked, but not quite

k_0 small ($G_0 \sim \frac{1}{k_0}$ large):

Crumpled universe

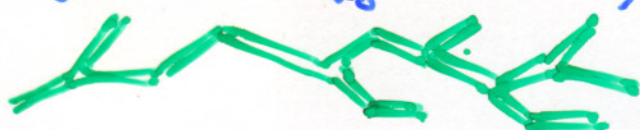


$D(N_4) \sim \text{const.}$

The natural entropic state of the (Euclidean) universe

k_0 large ($G_0 \sim \frac{1}{k_0}$ small)

Elongated universe



A phase transition separated these two phases.

Had it only been second order

Re-think:

Some general principles which could lead to a different selection of universes, with the same action.

CDT

Causal Dynamical Triangulations

Start in Minkowski space-time, insist on the requirement of causality for each space-time history (Trotter) and the existence of a global time coordinate.

(This causality requirement is different from the one we would impose on matter fields. !)

Only then rotate to Euclidean signature if needed.

Recall numerical setup: (Andrzej's talk)

foliation of $S^3 \times R$ ($S^3 \times S^1$)

For computer technical reasons we perform the simulations for fixed values of N_y :

$$Z(G, \Lambda) = \int \mathcal{D}[g_{\mu\nu}] e^{-\Lambda V_g + \frac{1}{G} \int \sqrt{g} R}$$

$$V_g = \int \sqrt{g}$$

$$Z(G, \Lambda) = \int d\nu e^{-\Lambda \nu} Z(G, \nu)$$

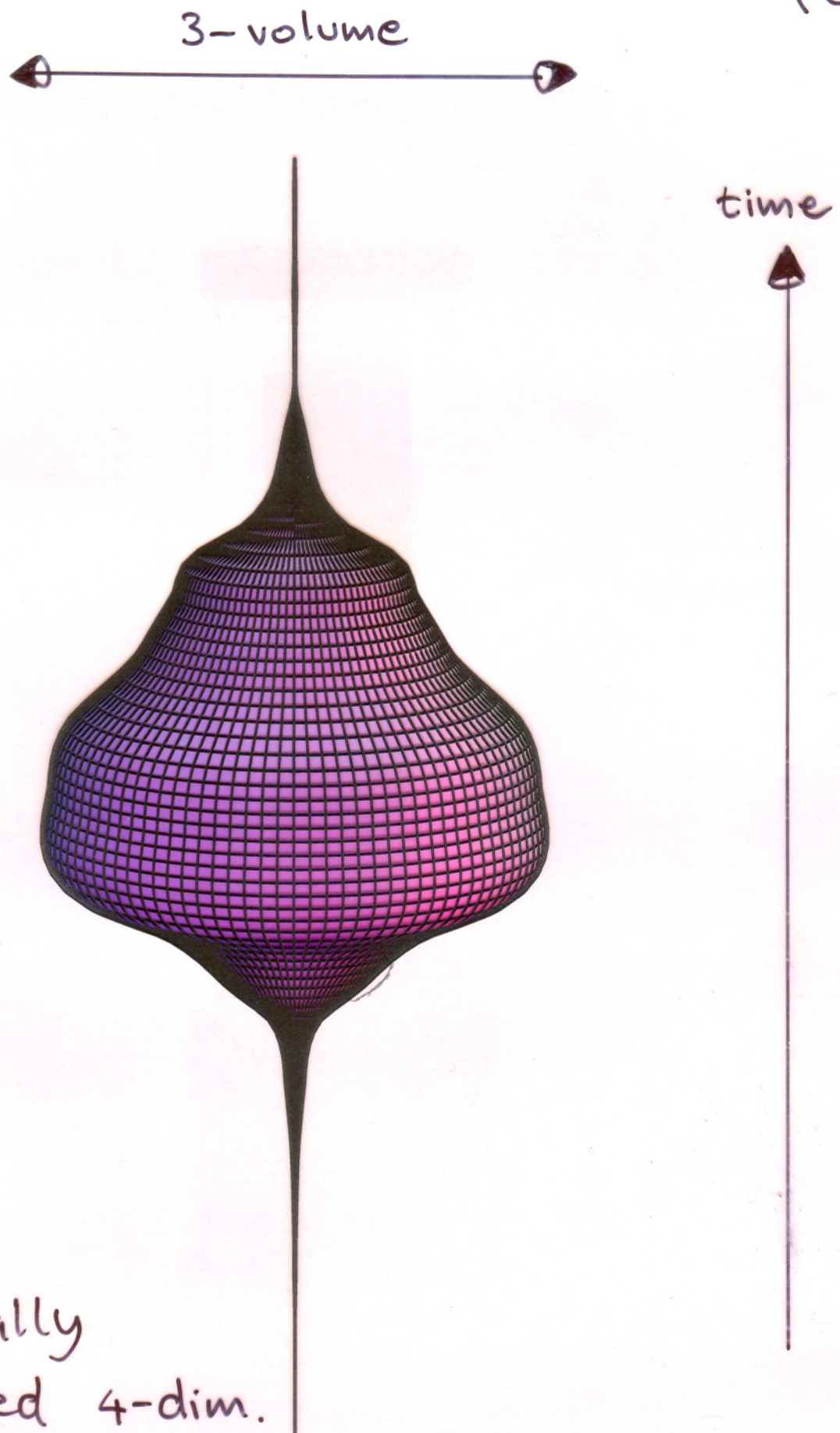
$$Z(G, \nu) = \int \mathcal{D}[g_{\mu\nu}] e^{\frac{1}{G} \int \sqrt{g} R} \delta(\int \sqrt{g} - \nu)$$

Thus:

$$\sum_{i=1}^{N_y} N_3(i) = N_y$$

What do we see?

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Dynamically
generated 4-dim.
quantum universe, obtained from a
causal path integral.

For different N_4 we have:

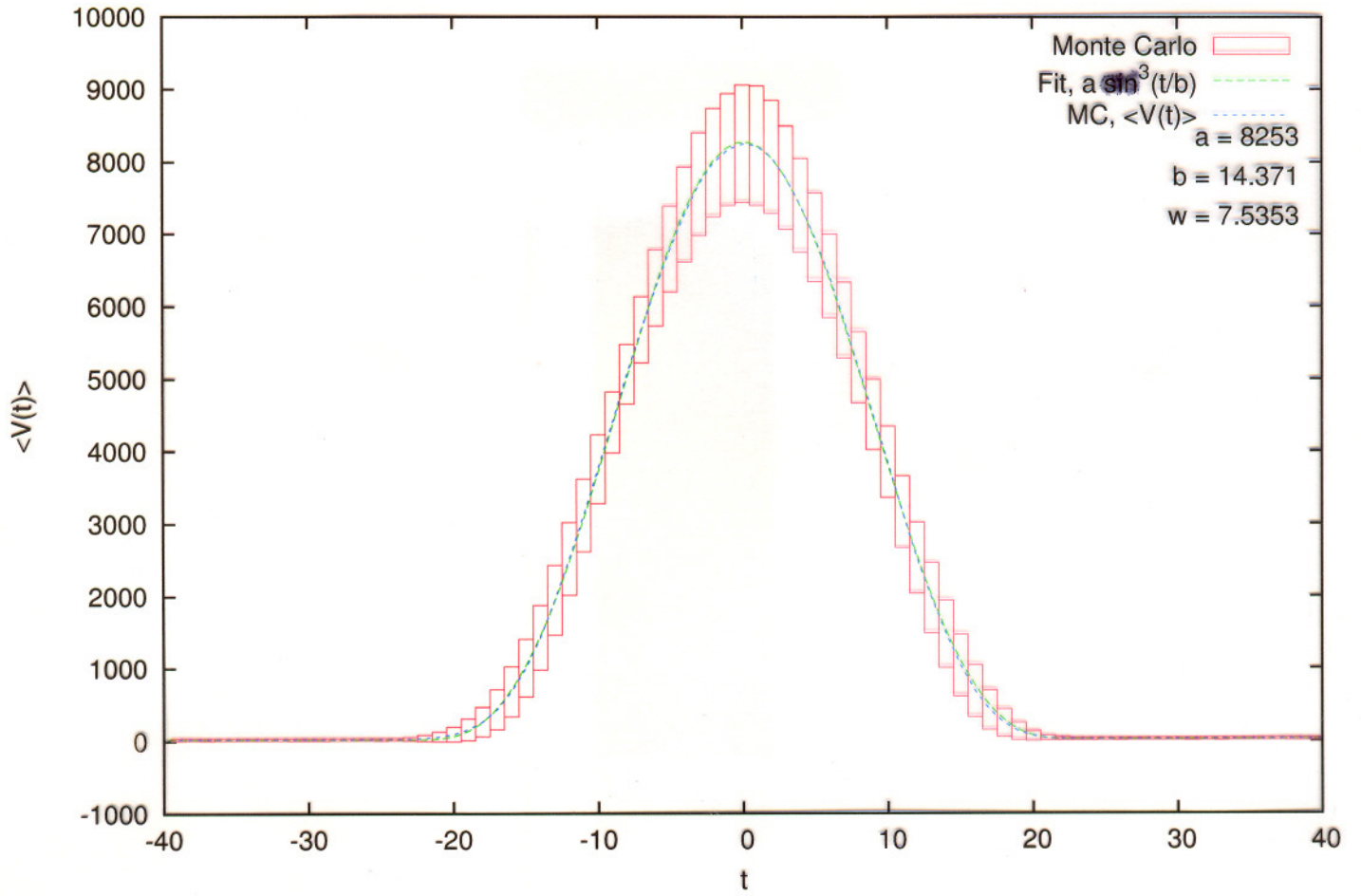
$$N_3(i) = N_4^{\frac{3}{4}} \frac{1}{S_0 N_4^{1/4}} \Omega^3 \frac{\dot{c}}{S_0 N_4^{1/4}}$$

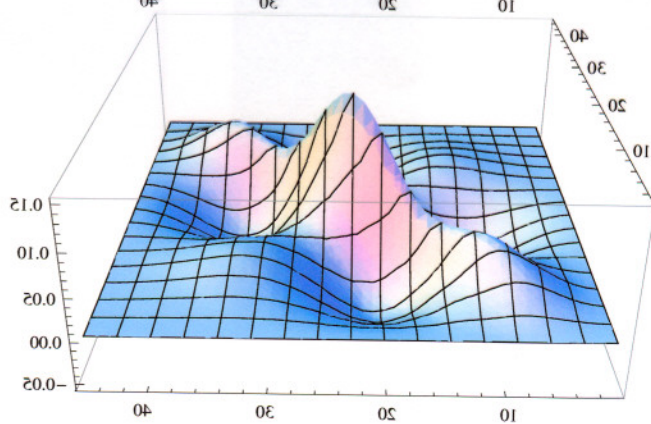
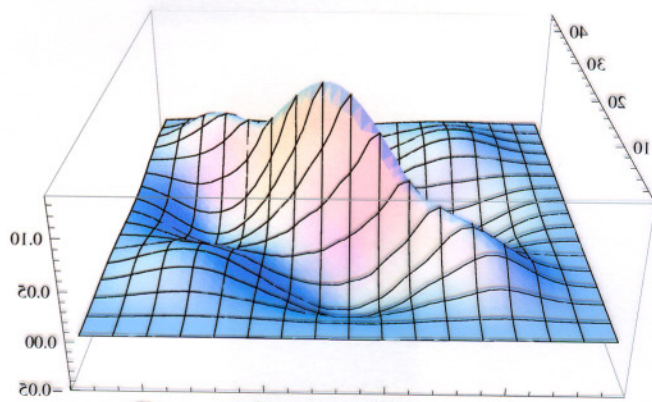
$$S_0 = 0.59 \text{ for } (K_0, \Delta) = (2.2, 0.6)$$

To obtain this curve we have a
average over different universes and
to compare different universes (comp. universes)
we have to fix the center of mass.

Ambiguity up to one lattice spacing (at least)

$K0 = 2.200000$, $\Delta = 0.600000$, $K4 = 0.925000$, $Vol = 160k$



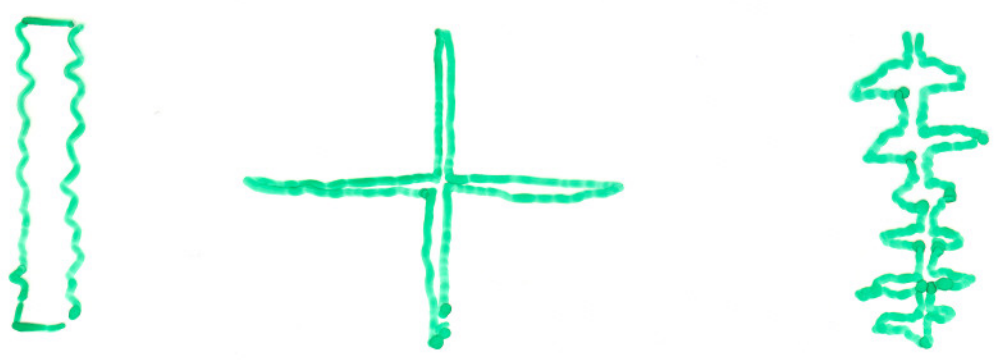


The scaling shows:

- ① Time extent of blob scales like $N_4^{1/4}$
- ② Spatial extent of blob scales as $N_4^{3/4}$

Thus: canonical time and space dimensions.

Nothing put in by hand: No background
 No reason it should not look like:



In fact we see all these "perverse" configurations for different choices of coupling constants.

Lesson from Euclidean approach, going all the way back to DT used in non-critical strings:

Just because you stack d-dimensional BB
 you do not necessarily obtain d-dim extended objects

Recall mini-superspace cosmology

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from the 80ies: (Hartle-Hawking

$$ds^2 = g_{tt}(dt)^2 + a^2(t) d\Omega_3$$

$$V_3(t) = 2\pi^2 a^3(t)$$

$$S = -\frac{1}{24\pi} \frac{1}{G} \int dt \sqrt{g} \left[\frac{g^{tt} \dot{V}_3^2}{V_3^4} + \kappa_2 V_3^{1/3} \right]$$

$$\int dt \sqrt{g_{tt}} V_3(t) = V_4$$

$$\sqrt{g_{tt}} V_3(t) = V_4 \frac{3}{4B} \cos^3 \frac{t}{B}, \quad \sqrt{g_{tt}} = \frac{R}{B}$$

$$V_4 = \frac{8\pi^2}{3} R^4$$

Natural discretization of this action:

$$(1) S_{Dis} = k_0 \sum_i \left[\frac{(N_3(i+1) - N_3(i))^2}{N_3(i)} + \tilde{\kappa}_3 N_3^{1/3}(i) \right]$$

$$(2) S_{Dis} = k_0 \sum_i \left[\frac{(N_3(i+1) - N_3(i))^2}{F(N_3(i))} - U(N_3(i)) \right]$$

While (1) by construction has a "classical" solution of the form observed, it is of interest to check to what extent the data allow us to deduce F and U

We want to derive MSF from first principles!

$$\bar{N}_3(i) = \frac{1}{K} \sum_{k=1}^K N_3^{(k)}(i) = \langle N_3(i) \rangle$$

$$C(i,j) = \langle (N_3(i) - \bar{N}_3(i)) (N_3(j) - \bar{N}_3(j)) \rangle$$

$$= \frac{1}{K} \sum_{k=1}^K (N_3^{(k)}(i) - \bar{N}_3(i)) (N_3^{(k)}(j) - \bar{N}_3(j))$$

$$n_{(i)} \equiv N_3(i) - \bar{N}_3(i)$$

$$C(i,j) = \langle n_{(i)} n_{(j)} \rangle$$

$$S_{\text{Dis}}(N) = S_{\text{Dis}}(\bar{N} + n) = S_{\text{Dis}}(\bar{N}) + \frac{1}{2} \sum_{i,j} n_i \hat{P}_{ij} n_j + O(n^3)$$

$$C(i,j) \approx \frac{\int \pi dn_i \ n_i n_j \ e^{-\frac{1}{2} \sum_{i,j} n_i \hat{P}_{ij} n_j + O(n^3)}}{\int \pi dn_i \ e^{-\frac{1}{2} \sum_{i,j} n_i \hat{P}_{ij} n_j + O(n^3)}}$$

In the quadratic approx:

$$\hat{C}(i,j) = (\hat{P}^{-1})_{ij}, \quad \hat{P}_{ij} = (C^{-1})_{ij}$$