

QCD and dynamical phase transitions, II

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- QGP and Bjorken Hydrodynamics

QGP: From Free-Streaming to Perfect-Fluid

- AdS/CFT and QCD Hydrodynamics

4 → 5d: Holographic renormalization

- AdS/CFT ⇒ asymptotic perfect fluid dynamics

Horizons vs. singularities

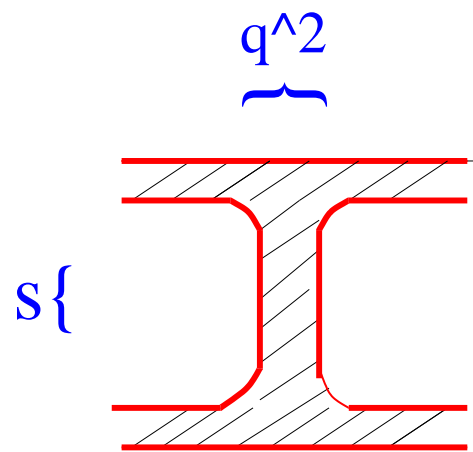
- Fast Thermalization to a perfect fluid

*AdS/CFT: Quasi-Normal Modes of a Moving
Black Hole*

^a(with **Romuald Janik**, Cracow U.) hep-th/0512162, hep-th/06.....
to appear

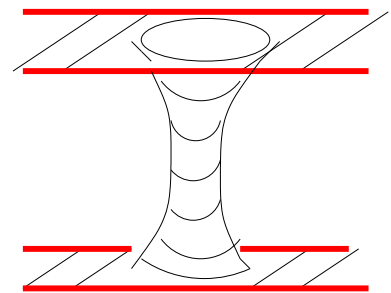
AdS/CFT Correspondence (1)

The Origin of String Theory



Veneziano Amplitude

$$A_{\mathbf{R}}(s, q^2)$$

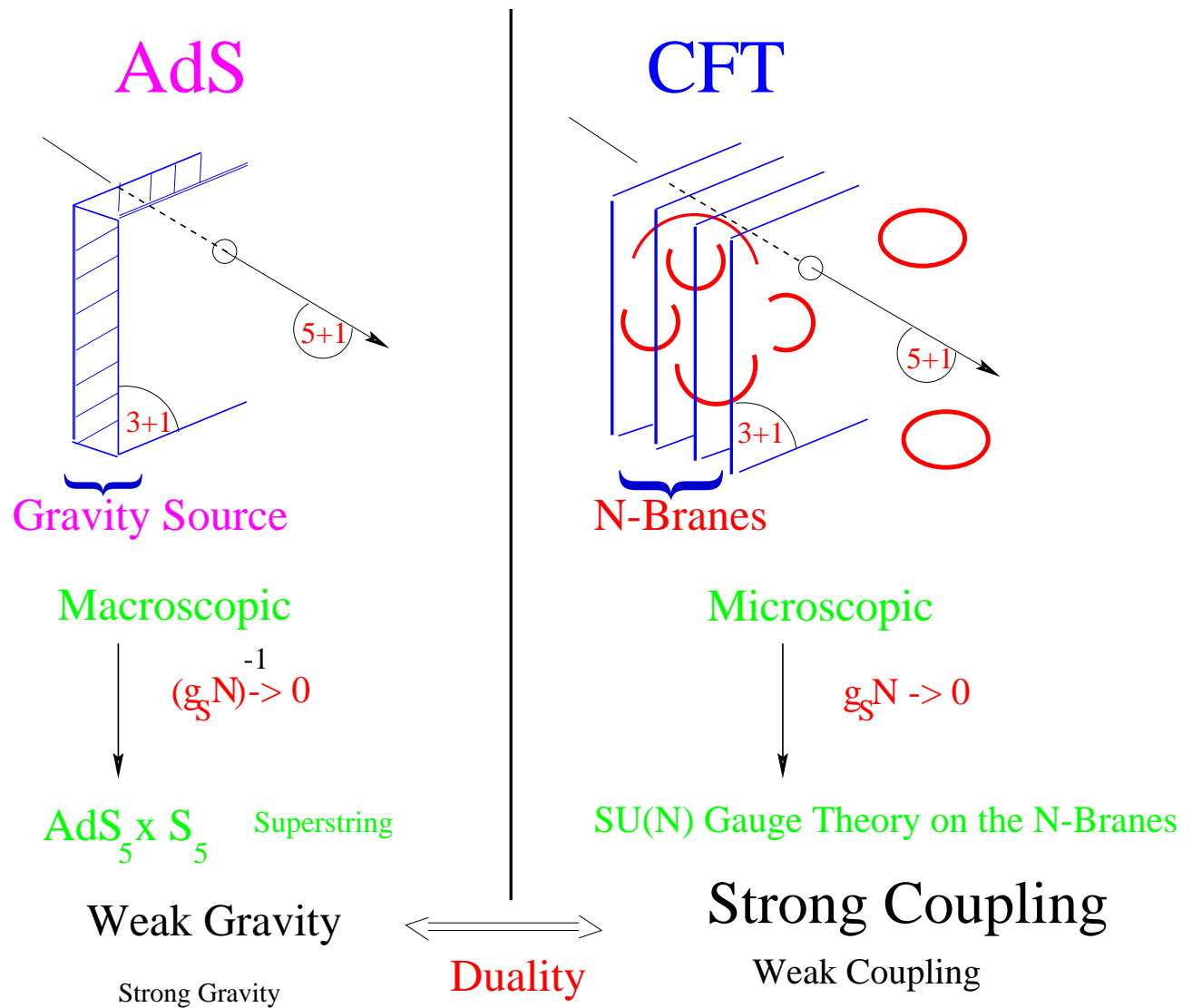


Shapiro-Virasoro Amplitude

$$A_{\mathbf{P}}(s, q^2)$$

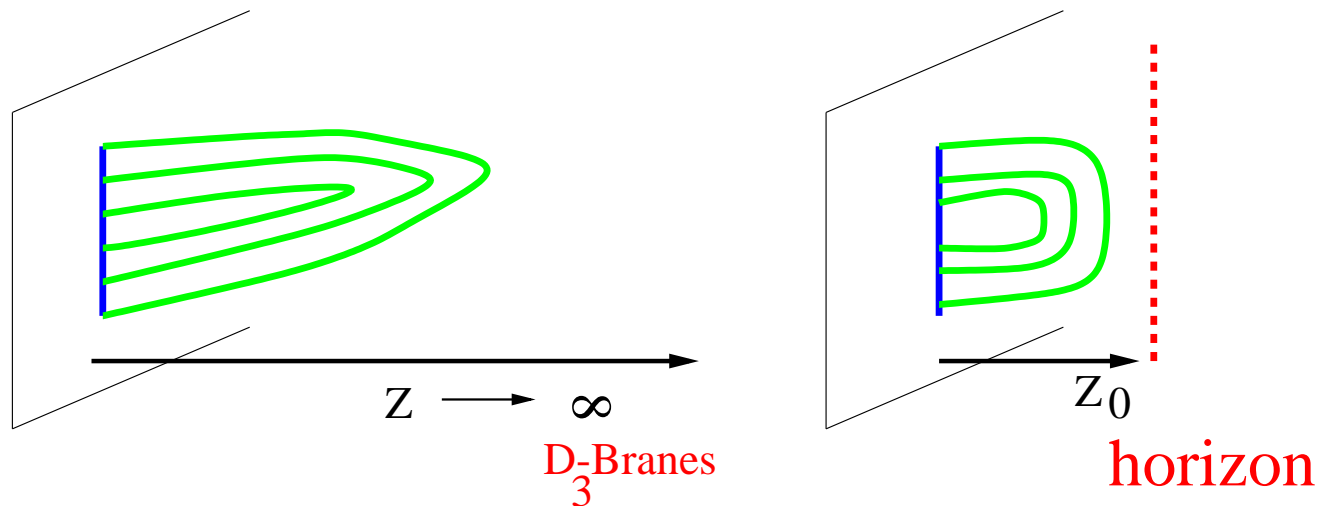
AdS/CFT Correspondence (2)

J.Maldacena (1998)



AdS/CFT Correspondence (3)

Holography



More on AdS₅

- D₃-brane Solution of Super Gravity:
≡ low energy IIB SuperString

$$ds^2 = f^{-1/2}(-dt^2 + \sum_{1-3} dx_i^2) + f^{1/2}(dr^2 + r^2 d\Omega_5)$$

“On-Branes × Out-Branes”

$$f = 1 + \frac{R^4}{r^4} ; R = 4\pi g_{YM}^2 \alpha'^2 N$$

- “Maldacena limit”:

$$\frac{\alpha'(\rightarrow 0)}{r(\rightarrow 0)} \rightarrow z , R \text{ fixed} \Rightarrow g_{YM}^2 N \rightarrow \infty$$

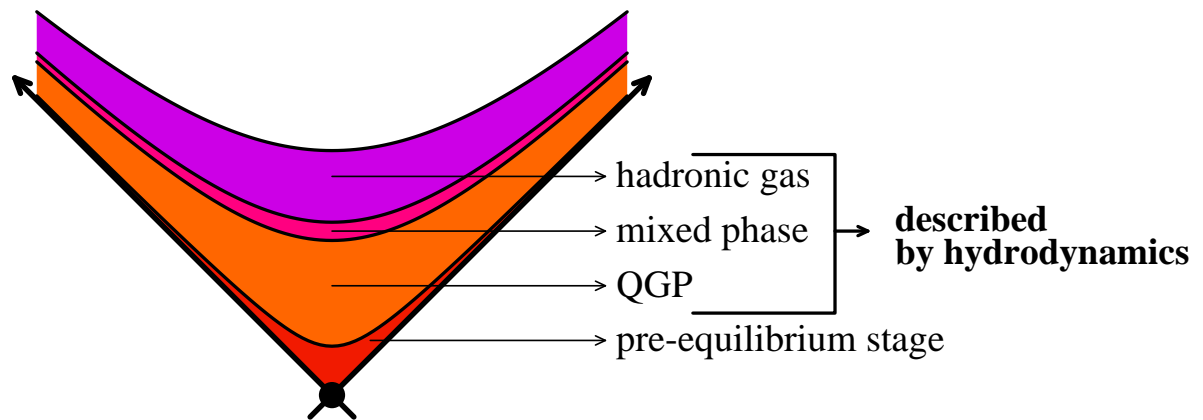
Strong coupling limit

$$ds^2 = \frac{1}{z^2}(-dt^2 + \sum_{1-3} dx_i^2 + dz^2) + R^2 d\Omega_5$$

Background Structure: AdS₅ × S₅

QGP formation and Relativistic Hydrodynamics

J.D.Bjorken (1982)



- Boost Invariance

$$\tau = \sqrt{x_0^2 + x_1^2} ; y = \frac{1}{2} \log \frac{x_0 + x_1}{x_0 - x_1} ; x_T \rightarrow x_1, x_2$$

- QGP: Perfect fluid behaviour
- Pre-equilibrium stage: Fast

The 4d Energy-Momentum Tensor

- Constraint Equations

$$\begin{aligned} T^\mu{}_\mu &\equiv -T_{\tau\tau} + \frac{1}{\tau^2}T_{yy} + 2T_{xx} = 0 \\ D_\nu T^{\mu\nu} &\equiv \tau \frac{d}{d\tau} T_{\tau\tau} + T_{\tau\tau} + \frac{1}{\tau^2}T_{yy} = 0 \end{aligned}$$

- Boost-invariant Tensor

$$T_{\mu\nu} = \begin{pmatrix} f(\tau) & 0 & 0 & 0 \\ 0 & -\tau^3 \frac{d}{d\tau} f(\tau) - \tau^2 f(\tau) & 0 & 0 \\ 0 & 0 & f(\tau) + \frac{1}{2}\tau \frac{d}{d\tau} f(\tau) & 0 \\ 0 & 0 & 0 & \dots \end{pmatrix}$$

- 1-parameter family

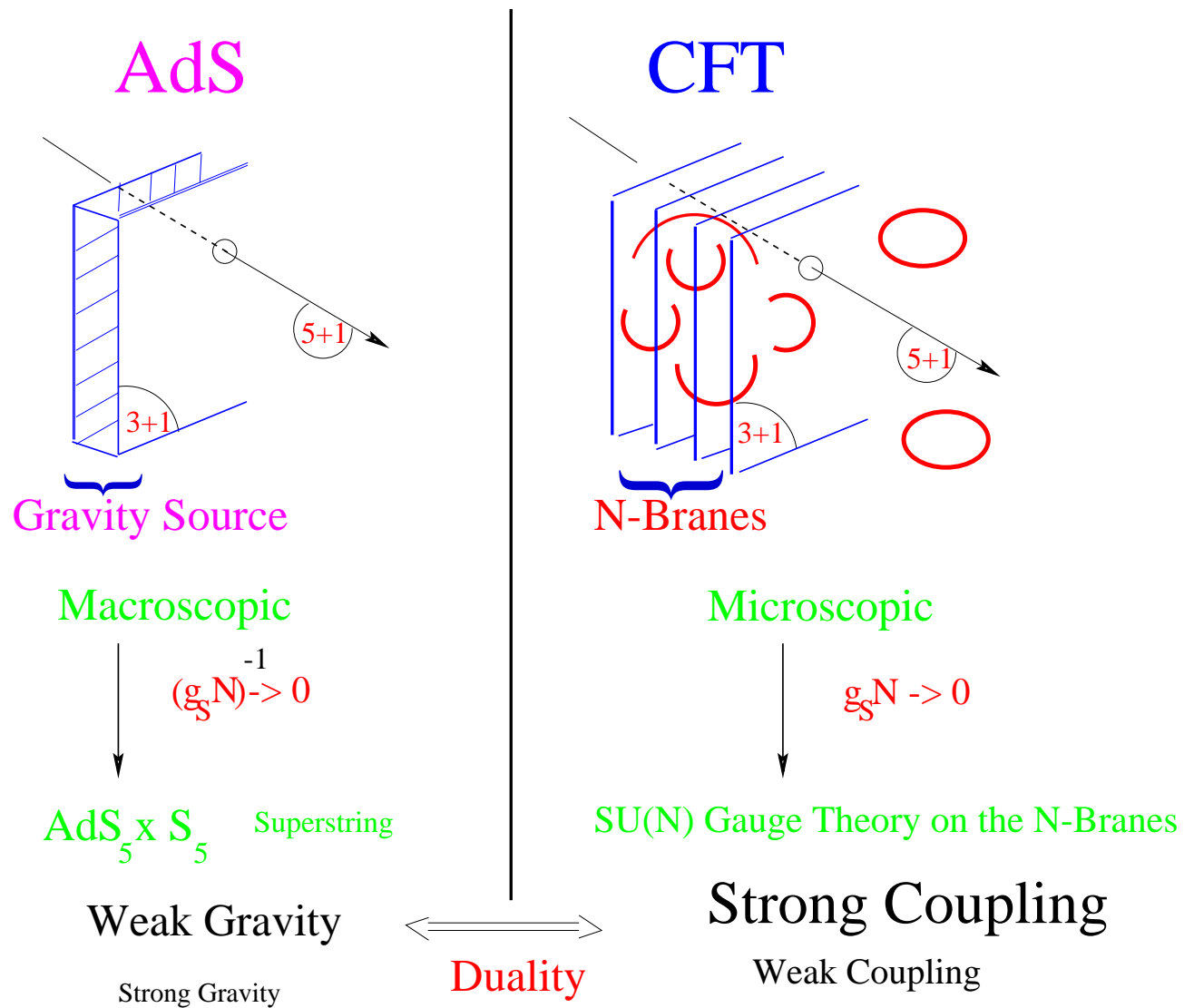
$$f(\tau) \propto \tau^{-s} : (0 < s < 4 ; T_{\mu\nu} t^\mu t^\nu \geq 0)$$

$$f(\tau) \propto \tau^{-\frac{4}{3}} : \text{Perfect Fluid}$$

$$f(\tau) \propto \tau^{-1} : \text{Free streaming}$$

AdS/CFT Correspondence

J. Maldacena (1998)



4 \rightarrow 5d:
Holographic Renormalization

K. Skenderis (2002)

- Fefferman-Graham Coordinates:

$$ds^2 = \frac{g_{\mu\nu} dx^\mu dx^\nu + dz^2}{z^2}$$

- 4d \Leftrightarrow 5d metric:

$$g_{\mu\nu} = g_{\mu\nu}^{(0)} (= \eta_{\mu\nu}) + z^2 g_{\mu\nu}^{(2)} (= 0) + z^4 g_{\mu\nu}^{(4)} (\propto \langle T_{\mu\nu} \rangle) + \dots$$

+ ...: from Einstein Eqs.

Test on Statics: 4d Perfect Fluid \Leftrightarrow 5d Black Hole

- 4d Perfect Fluid

$$\langle T_{\mu\nu} \rangle \propto g_{\mu\nu}^{(4)} = \begin{pmatrix} 3/z_0^4 & 0 & 0 & 0 \\ 0 & 1/z_0^4 & 0 & 0 \\ 0 & 0 & 1/z_0^4 & 0 \\ 0 & 0 & 0 & 1/z_0^4 \end{pmatrix}$$

- Derivation of the Fefferman-Graham metrics

$$ds^2 = -\frac{(1 - z^4/z_0^4)^2}{(1 + z^4/z_0^4)z^2} dt^2 + (1 + z^4/z_0^4) \frac{dx^2}{z^2} + \frac{dz^2}{z^2}$$

- It is indeed the Static Black Hole

$$\tilde{z} = \frac{z}{\sqrt{1 + \frac{z^4}{z_0^4}}}$$

$$ds^2 = -\frac{1 - \tilde{z}^4/\tilde{z}_0^4}{\tilde{z}^2} dt^2 + \frac{dx^2}{\tilde{z}^2} + \frac{1}{1 - \tilde{z}^4/\tilde{z}_0^4} \frac{d\tilde{z}^2}{\tilde{z}^2}$$

Boost-Invariant Geometries

- **Boost-Invariant 5-d F-G metric:**

$$ds^2 = \frac{-e^{a(\tau,z)} d\tau^2 + \tau^2 e^{b(\tau,z)} dy^2 + e^{c(\tau,z)} dx_{\perp}^2}{z^2} + \frac{dz^2}{z^2}$$

- **Einstein Equation(s):**

$$[a(\tau, z), b(\tau, z), c(\tau, z)] = [a(v), b(v), c(v)] + \mathcal{O}\left(\frac{1}{\tau^{\#}}\right)$$

$$v = \frac{z}{\tau^{s/4}}$$

$$\begin{aligned} v(2a'(v)c'(v) + a'(v)b'(v) + 2b'(v)c'(v)) - 6a'(v) - 6b'(v) - 12c'(v) + vc'(v)^2 &= 0 \\ 3vc'(v)^2 + vb'(v)^2 + 2vb''(v) + 4vc''(v) - 6b'(v) - 12c'(v) + 2vb'(v)c'(v) &= 0 \\ 2vzb''(v) + 2sb'(v) + 8a'(v) - vsa'(v)b'(v) - 8b'(v) + vsb'(v)^2 + \\ 4vsc''(v) + 4sc'(v) - 2vsa'(v)c'(v) + 2vsc'(v)^2 &= 0 . \end{aligned}$$

- **Asymptotic Solution**

$$\begin{aligned} a(v) &= A(v) - 2m(v) \\ b(v) &= A(v) + (2s - 2)m(v) \\ c(v) &= A(v) + (2 - s)m(v) \end{aligned}$$

Dual of a Perfect Relativistic fluid

$$v = \frac{z}{\tau^{1/3}}$$

- Asymptotic metric

$$ds^2 = \frac{1}{z^2} \left[-\frac{\left(1 - \frac{e_0}{3} \frac{z^4}{\tau^{4/3}}\right)^2}{1 + \frac{e_0}{3} \frac{z^4}{\tau^{4/3}}} d\tau^2 + \left(1 + \frac{e_0}{3} \frac{z^4}{\tau^{4/3}}\right) (\tau^2 dy^2 + dx_{\perp}^2) \right] + \frac{dz^2}{z^2}$$

- Black Hole off in the 5th dimension

$$\text{Horizon : } z_0 = \left(\frac{3}{e_0}\right)^{\frac{1}{4}} \cdot \tau^{\frac{1}{3}} .$$

$$\text{Temperature : } T(\tau) \sim \frac{1}{z_0} \sim \tau^{-\frac{1}{3}}$$

$$\text{Entropy : } S(\tau) \sim \text{Area} \sim \tau \cdot \frac{1}{z_0^3} \sim \text{const}$$

Free Streaming and general cases

$$v = \frac{z}{\tau^{1/4}}$$

- **Asymptotic metric**

$$ds^2 = \frac{\left(-\left(1 + \frac{v^4}{\sqrt{8}}\right) \frac{1-2\sqrt{2}}{2} \left(1 - \frac{v^4}{\sqrt{8}}\right) \frac{1+2\sqrt{2}}{2} dt^2 + \left(1 + \frac{v^4}{\sqrt{8}}\right)^{\frac{1}{2}} \left(1 - \frac{v^4}{\sqrt{8}}\right)^{\frac{1}{2}} \tau^2 dy^2 + \right.}{z^2} \\ \left. + \frac{\left(1 + \frac{v^4}{\sqrt{8}}\right) \frac{1+\sqrt{2}}{2} \left(1 - \frac{v^4}{\sqrt{8}}\right) \frac{1-\sqrt{2}}{2} dx_{\perp}^2 \right)}{z^2} + \frac{dz^2}{z^2}$$

- **True Singularities ?**

Ricci scalar:

$$R = -20 + \mathcal{O}\left(\frac{1}{\tau^{\#}}\right) .$$

Riemann tensor squared:

$$\mathfrak{R}^2 = R^{\mu\nu\alpha\beta} R_{\mu\nu\alpha\beta}$$

AdS/CFT: Selection of the Perfect Fluid

- Curvature invariant:

$$\mathfrak{R}^2 = R^{\mu\nu\alpha\beta} R_{\mu\nu\alpha\beta}$$

$$\mathfrak{R}^2 = \frac{4}{(1 - \Delta(s)^2 v^8)^4} \cdot \left[\begin{aligned} &10 \Delta(s)^8 v^{32} - 88 \Delta(s)^6 v^{24} + 42 v^{24} s^2 \Delta(s)^4 + \\ &+ 112 v^{24} \Delta(s)^4 - 112 v^{24} \Delta(s)^4 s + 36 v^{20} s^3 \Delta(s)^2 - 72 v^{20} s^2 \Delta(s)^2 + \\ &+ 828 \Delta(s)^4 v^{16} + 288 v^{16} \Delta(s)^2 s - 288 v^{16} \Delta(s)^2 - 108 v^{16} s^2 \Delta(s)^2 + \\ &- 136 v^{16} s^3 + 27 v^{16} s^4 - 320 v^{16} s + 160 v^{16} + 296 v^{16} s^2 + 36 v^{12} s^3 + \\ &- 72 v^{12} s^2 - 88 \Delta(s)^2 v^8 + 42 v^8 s^2 + 112 v^8 - 112 v^8 s + 10 \end{aligned} \right] + \mathcal{O}\left(\frac{1}{\tau^\#}\right)$$

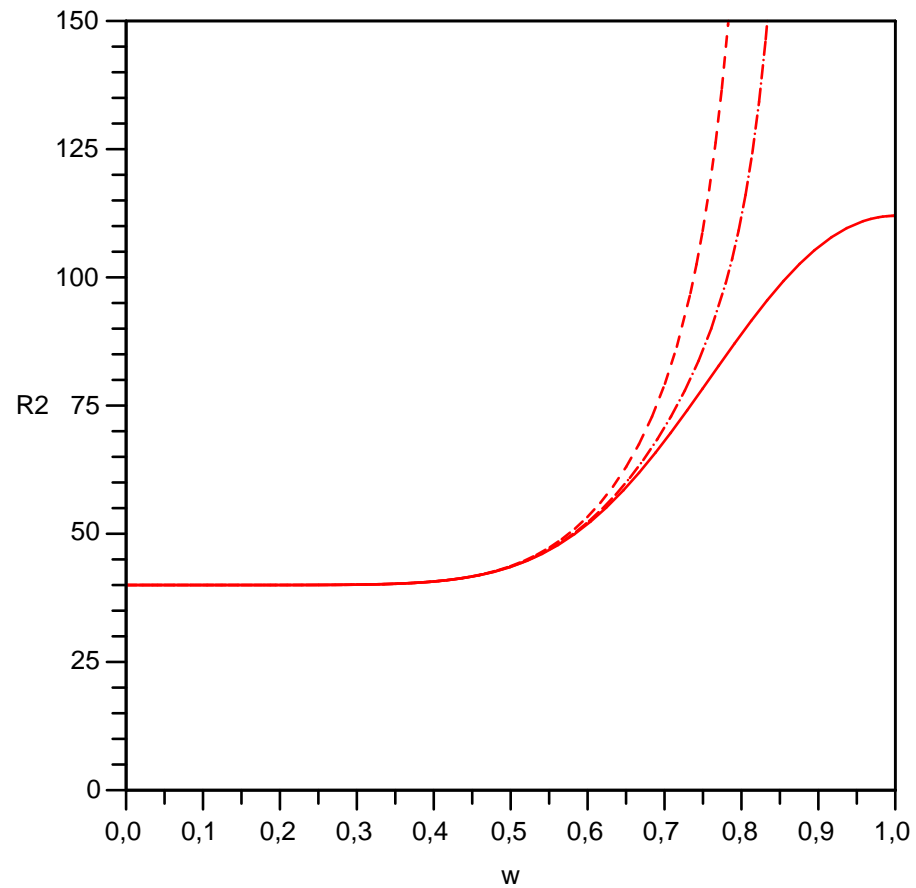
- \mathfrak{R}^2 for $s = \frac{4}{3}$:

$$\mathfrak{R}^2_{\text{perfect fluid}} = \frac{8(5w^{16} + 20w^{12} + 174w^8 + 20w^4 + 5)}{(1 + w^4)^4}$$

where $w = v/\Delta\left(\frac{4}{3}\right)^{\frac{1}{4}} \equiv \sqrt[4]{3} v$.

AdS/CFT: Selection of the Perfect Fluid

$$\mathcal{R}^2 \text{ for } s = \frac{4}{3} \pm .1$$



Thermalization response-time of a perfect fluid

- **Quasi-normal scalar modes of a static Black Hole in Fefferman-Graham coordinates**

$$\Delta\phi \equiv \frac{1}{\sqrt{-g}} \partial_i (\sqrt{-g} g^{ij} \partial_j \phi) = 0$$

$$-\frac{1}{z^3} \frac{\left(1 - \frac{z^4}{z_0^4}\right)^2}{\left(1 + \frac{z^4}{z_0^4}\right)} \partial_t^2 \phi(t, z) + \partial_z \left(\frac{1}{z^3} \left(1 - \frac{z^8}{z_0^8}\right) \partial_z \phi(t, z) \right) = 0.$$

- **Separation of variables** $\phi(t, z) = e^{i\omega t} \phi(z)$

$$\phi'' + \frac{1 - \tilde{z}^2}{\tilde{z}(1 - \tilde{z})(2 - \tilde{z})} \phi' + \left(\frac{\omega}{\pi T}\right)^2 \frac{1}{4\tilde{z}(1 - \tilde{z})(2 - \tilde{z})} \phi = 0$$

- **Dominant Decay Time**

$$\frac{\omega_c}{\pi T} \sim 3.1194 - 2.74667 i$$

- **A note on Viscosity vs. Q-n modes**

Thermalization response-time of a perfect fluid

- Quasinormal scalar modes for the boost-invariant Black Hole geometry

$$\Delta\phi \equiv \frac{1}{\sqrt{-g}} \partial_i (\sqrt{-g} g^{ij} \partial_j \phi) = 0$$

$$\left[\partial_z \rightarrow \tau^{-\frac{1}{3}} \partial_v ; \partial_\tau \rightarrow \partial_\tau - \frac{1}{3} \tau^{-\frac{4}{3}} \partial_v \right]$$

$$-\frac{1}{v^3} \frac{(1+v^4)^2}{1-v^4} \partial_\tau^2 \phi(\tau, v) + \tau^{-\frac{2}{3}} \partial_v \left(\frac{1}{v^3} (1-v^8) \partial_v \phi(\tau, v) \right) = 0$$

- Separation of variables $\phi(\tau, v) = f(\tau)\phi(v)$

$$\partial_\tau^2 f(\tau) = -\omega^2 \tau^{-\frac{2}{3}} f(\tau) \Rightarrow f(\tau) = \sqrt{\tau} J_{\pm\frac{3}{4}} \left(\frac{3}{2} \omega \tau^{\frac{2}{3}} \right) \sim \tau^{\frac{1}{6}} e^{\frac{3}{2} i \omega \tau^{\frac{2}{3}}}$$

$$\partial_v \left(\frac{1}{v^3} (1-v^8) \partial_v \phi(v) \right) + \omega^2 \frac{1}{v^3} \frac{(1+v^4)^2}{1-v^4} \phi(v) = 0$$

- Dominant Decay Proper-Time

$$\frac{\omega_c}{\pi T} \sim 3.1194 - 2.74667 i$$

Conclusions II

- **AdS/CFT and Hydrodynamics**
Construction of the “Dual” of a relativistic fluid
- **4d \rightarrow 5d : Holographic Renormalisation**
5-d Metric from the 4-d $T_{\mu\nu}$
- **5d Physical Criterium**
Non-singular 5d Horizon \Rightarrow Asymptotic
4d perfect fluid
- **Quasi-Normal mode analysis:**
The perfect fluid is thermally “very stable”

Unexpected AdS/CFT consequences for QCD at strong coupling?