

# Interaction between kink and radiation in $\phi^4$ Model

Negative radiation pressure and fractals

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# Outline

- 1 Introduction
  - Model and linearization
  - Spectral structure
  - Interactions
- 2 Kink in monochromatic wave
  - Radiation pressure
  - Negative radiation pressure - first order
  - second order
- 3 Excitation of internal degrees
  - Oscillating mode and creation
  - Creation of pairs and fractal boundary
  - Simplified theory
- 4 Another fractal example
  - Annihilation and scattering of kinks
  - n-bounce windows

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Let us consider real scalar field theory in 1+1 dim:

$$\ddot{\phi} - \phi'' + 2\phi(\phi^2 - 1) = 0.$$

The well known kink solution:

$$\phi_s(x) = \tanh x$$

can be perturbed with a small field  $\phi(x, t) = \phi_s(x) + \xi(x, t)$ .

We can write an equation for  $\xi$ :

$$\ddot{\xi} + \hat{L}\xi + N(\xi) = 0,$$

where

$$\hat{L} = -\frac{\partial^2}{\partial x^2} + 2(3 \tanh^2 x - 1)$$

and

$$N(\xi) = 6\phi_s\xi^2 + 2\xi^3.$$

Neglecting the nonlinear part we seek solutions in the form  $\xi = e^{i\omega_k t} \eta_k(x)$ , where  $\omega_k^2 = k^2 + 4$ .

### Spectral structure of the solutions

- one **translational** mode  $\phi_s(x + \delta x) = \phi_s(x) + \delta x \eta_t(x) + \mathcal{O}(\delta x^2)$

$$\eta_t(x) = \frac{1}{\cosh^2 x}, \quad \omega = 0, \quad (k = 2i)$$

- one **discrete** mode

$$\eta_d(x) = \frac{\tanh x}{\cosh x}, \quad \omega_d = \sqrt{3}, \quad (k = i)$$

- continues spectrum of **scattering** modes (radiation):

$$\eta_k(x) = e^{ikx} \left( 3 \tanh^2 x - 3ik \tanh x - 3 - k^2 \right)$$



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## What is interesting

- We can excite the oscillational mode causing
  - radiation and decay of this mode for small amplitude [Manton]
    - creation of two kinks and radiation
- We can reverse the process and create radiation far away and in that way excite
  - the translational mode
  - or oscillational mode (asymmetric) and then creation of kinks

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Let the perturbation  $\xi$  has a form:

$$\xi(x, t) = \frac{1}{2} A \eta_q(x) e^{i\omega_q t} + c.c.$$

### Question

How will the kink behave?

*Answer (?)* The kink will be pushed by the radiation

We simulate numerically the partial equation with conditions:

$$\phi(x, t=0) = \phi_s(x), \quad \dot{\phi}(x, t=0) = 0$$

$$\phi(-L, t) = 0, \quad \phi(L, t) = A \sin \omega_q t$$



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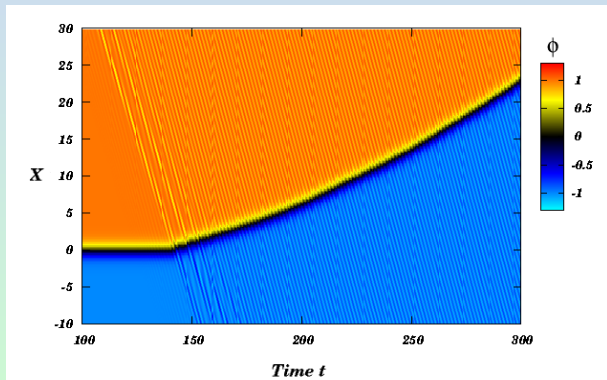
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# Kink motion in the field of radiation

$$A = 0.20, \omega_q = 4.0$$



Surprise!

The kink is going toward the source of radiation!

We have "negative radiation pressure".

The same happens to sine-Gordon model.

But this is not a generic feature. Most models behave "properly":

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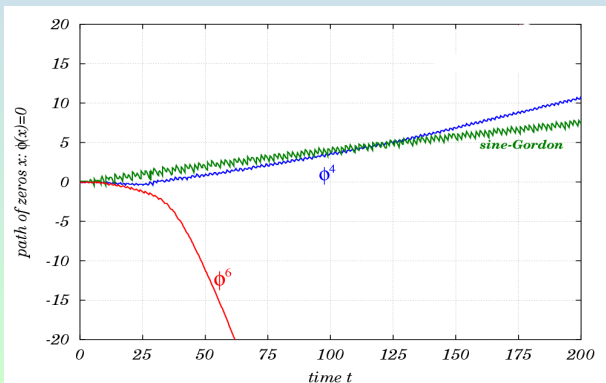
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## Question

What is so special in these theories?

The scattering modes have no reflection part

$$\eta_k(x) = e^{ikx} \left( 3 \tanh^2 x - 3ik \tanh x - 3 - k^2 \right) \approx Be^{ikx} \quad (x \rightarrow \pm\infty)$$

kinks are **transparent** in linear approximation.

The same for sine-Gordon equation (but there is no discrete oscillational mode).

## Reflectionless [Bordag]

All reflectionless spectra for potentials

$$V(x) = N(N+1) \tanh^2 x.$$

$N = 2$  for  $\phi^4$  and  $N = 1$  for s-G

$N - 1$  - number of discrete modes

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We move up to the second order and solve the equation for  $\xi$ .  
We seek the solution in the form of a series:

$$\xi(\mathbf{x}, t) = A\xi^{(1)}(\mathbf{x}, t) + A^2\xi^{(2)} + \dots,$$

where  $\xi^{(1)} = e^{i\omega_q t} \eta_q(\mathbf{x}) + \text{c.c.}$ . The equation for  $\xi^{(2)}$  is

$$\ddot{\xi}^{(2)} + \hat{L}\xi^{(2)} + \frac{3}{4}\phi_s \left( \eta_q^2 e^{2i\omega_q t} + 2\eta_q \eta_{-q} + \eta_{-q} e^{-2i\omega_q t} \right) = 0.$$

This is inhomogeneous linear partial equation. We can find a solution in a form

$$\xi^{(2)} = \xi_{+2}^{(2)}(\mathbf{x}) e^{2i\omega_q t} + \xi_{-2}^{(2)}(\mathbf{x}) e^{-2i\omega_q t} + \xi_0^{(2)}(\mathbf{x}, t).$$

## Note that

$$\langle \phi_s \eta_q \eta_{-q} | \eta_t \rangle = \langle \phi_s \eta_q^2 | \eta_t \rangle = \langle \phi_s \eta_{-q}^2 | \eta_t \rangle = 0$$

(because of reflectionless spectrum).

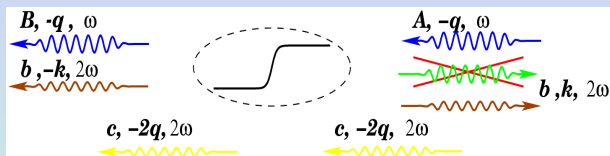
We are interested only in a time dependent part, and since we already know the solutions of homogeneous equation we can construct the Green's function.

It is quite easy to get the asymptotic form for large  $|x|$ :

$$\xi_{\pm\infty}^{(2)}(x, t) = b_{\pm}(\omega_q) \cos(2\omega_q t \mp kx \pm \delta_1) + c(2\omega_q) \cos(2\omega_q t + 2qx \pm \delta_2),$$

where  $k = k(q) = \sqrt{(2\omega_q)^2 - 4}$ .

The expressions for  $b$  and  $c$  are complicated but can be calculated analytically.



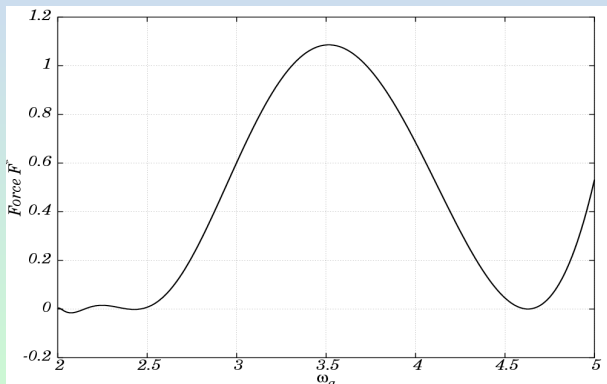
Having these we can write the conservation laws for energy and momentum inside the large segment and averaged over a period:

$$q\omega_q A^2 - q\omega_q B^2 - 2k\omega_q (b_-^2 + b_+^2) = \frac{dE}{dt} = 2M^* \frac{d\gamma}{dt}$$

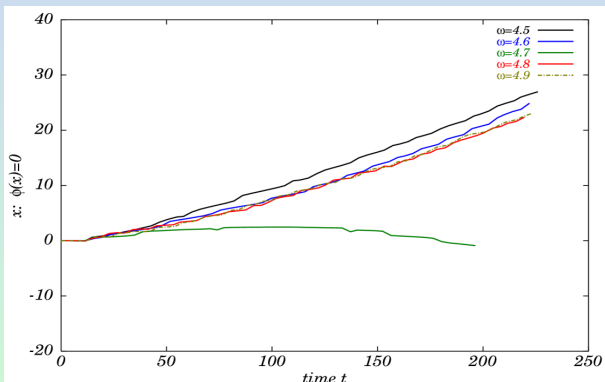
and

$$q^2 A^2 - q^2 B^2 - k^2 (b_-^2 - b_+^2) = -\frac{dP}{dt} = -F^*$$

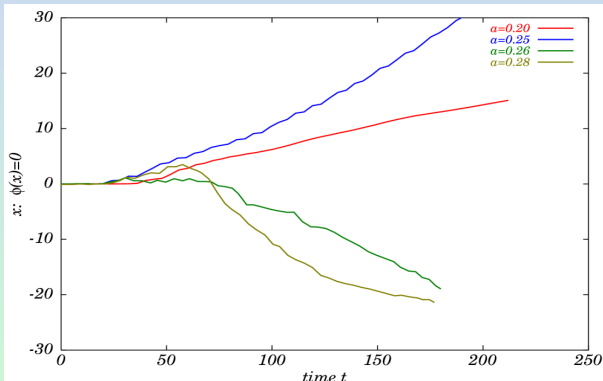
As a solution of this system of equations we obtain the force with which the kink is being pulled by this radiation:



We can test our predictions for the minimum:



But for large amplitudes the higher orders are more important:



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Let us take a closer look at the second possible process:



We can simulate this process numerically.

In order to measure the excitation of the osc. mode we calculate the projection onto our mode:

$$A_d(t) = \langle \phi - \phi_s | \eta_d \rangle$$

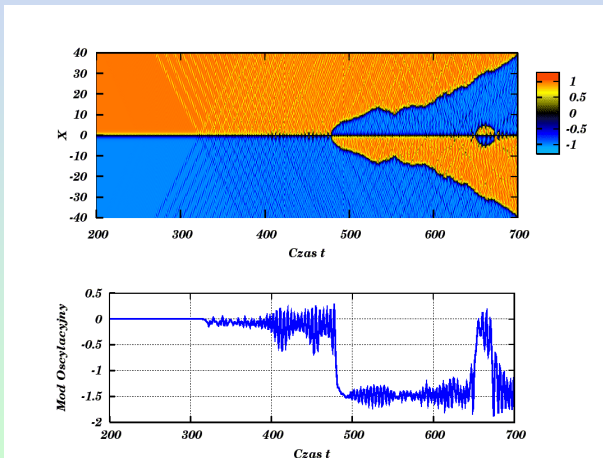
The radiation in linear approx is orthogonal to  $\eta_d$ .

When the creation process occurs instead of a kink an antikink remains and two kinks are radiated out:

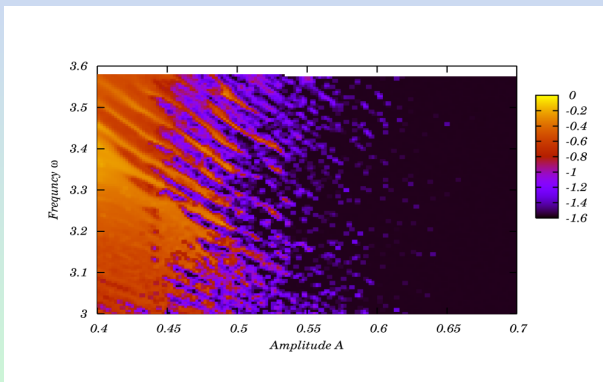
What we see is

$$A_d(t) = \langle -2\phi_s | \eta_d \rangle = -\frac{\pi}{2}$$

Example figure:



Finally we can present the figure:

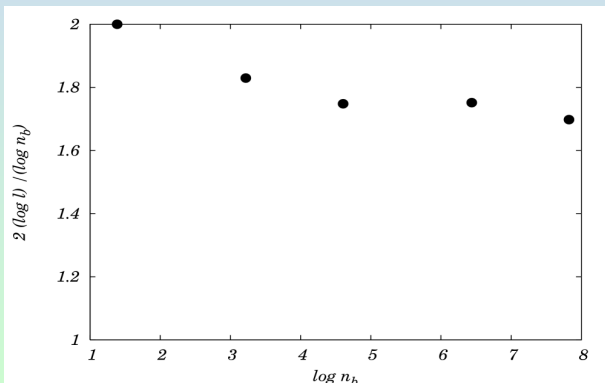


Notice a very complicated (fractal?) critical line for production of defects.

We can measure the fractal dimension:

$$D = \lim_{l \rightarrow 0} \frac{2 \log l}{\log n_b},$$

where  $l$  - length of a box,  $n_b$  - number of boxes containing boundary.



Again let us use the second order equation:

$$\ddot{\xi} - \hat{L}\xi + 6\phi_s\xi^2 = 0$$

but now  $\xi(x, t) = Ah_k(x) \cos \omega t + A_d(t)\eta_d(x) + (\eta)$ .

We substitute the above to our eq. and project onto the oscillational mode (integrate with  $\eta_d$ ):

$$\ddot{A}_d + \omega_d^2 A_d + \alpha(k)A^2 \cos^2 \omega t + \beta(k)AA_d \cos \omega t + \gamma A_d^2 = 0.$$

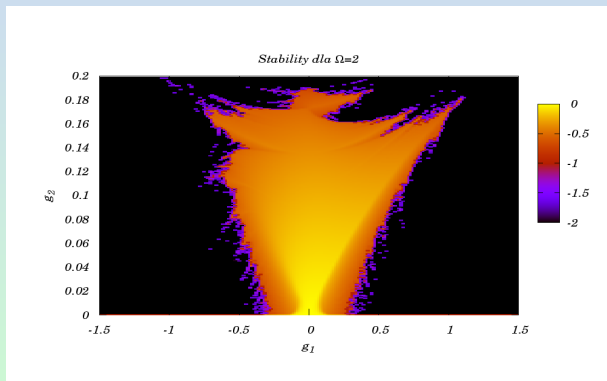
By rescaling we can see there are only three important coefficients:

$$\ddot{u} + u + u^2 + g_1 u \cos \Omega t + g_2 \cos^2 \Omega t = 0.$$

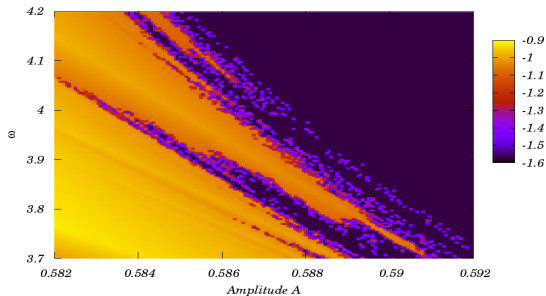
The above eq. is similar to the Mathieu's eq. but with nonlinear term and external force.

The above equations are correct for all double well field th. in 1+1d (the difference is in  $\alpha$  and  $\beta$ ) where we can expand the potentials into the Taylor's series (exception: compactons).

Therefore it is interesting to investigate this equation for different  $g_1$  and  $g_2$ :



Or finally we can reproduce the figure:



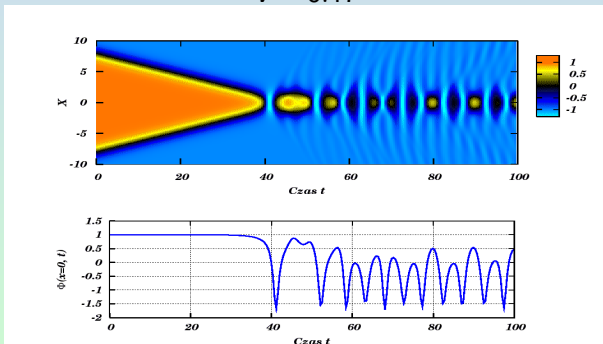


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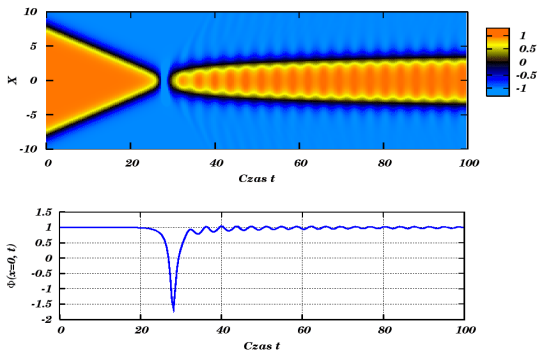
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Matzner studied collision of two kinks.  
The kinks can either annihilate creating pseudo-breather (bion, oscillon) or can be scattered back

$$v = 0.17$$



$$v = 0.26$$



When  $v$  is between 0.18 and 0.26 for certain velocities the kinks can come back and reflect once more.

We can observe a narrowing series of **two-bounce windows**

Each two-bounce window is surrounded by series of narrowing **three-bounce windows**.

Each three-bounce window is...

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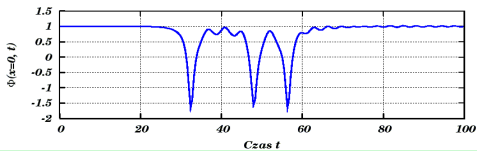
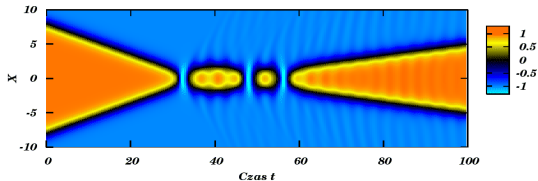
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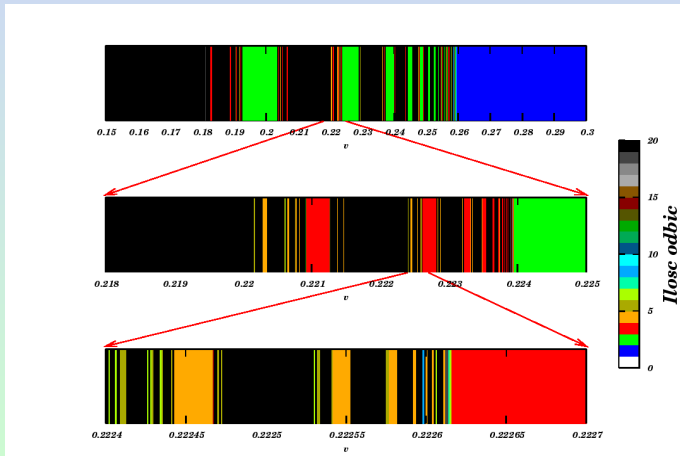
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# Three-bounce window

$$v = 0.221$$








# nice fractal structure















# Literature

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*Nonlinearity*, **12**, 851 (1997).
-  T.R.  
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