# The Euclidean contour rotation in quantum gravity 

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Abstract: The talk will discuss the rotation of the contour of functional integration in quantum gravity from Lorentzian geometries to Euclidean geometries. In the usual framework of metric tensors the functional integral does not have a good definition and so the formulas are necessarily heuristic. However it is hoped that these formulas will provide exact mathematical results when applied to theories that are constructed with a fundamental Planck scale cut-off.

Euclidean "Physics"
Euclidean signature ++++

- Euclidean QG (Hawking)
- Heat Kernel Expansion
- NCG \& Connes-Chamseddine spectral action
$J: H \rightarrow H$ Dirac
N: $\boldsymbol{N} \rightarrow \partial C$ charge conj.


$$
\begin{aligned}
& \delta_{c c}=\operatorname{tr}\left(D^{2} / \Lambda^{2}\right)+\langle J \psi, D \psi\rangle \quad \Lambda=p l_{\text {anck mall }} \\
& \underset{n \rightarrow \infty}{\sim} c \Lambda^{4} \int d v+c^{\prime} \Lambda^{2} \int R d v+\ldots+S M
\end{aligned}
$$

Formula to "prove"

$$
\int_{G_{L}, M_{L}} e^{i S_{L}(l, \phi)} D l D \phi=\int_{G_{E}, M_{E}} e^{-S_{E}(e, \Phi)} D e D \Phi
$$

$\begin{array}{lll}l \in G_{L} & \text { Lorentzian geometry } & e \in G_{E} \\ \phi \in M_{L} & \text { Euclidean gamete } \\ \text { Mater fields } & \Phi \in M_{E} & \text { Euclidean matter }\end{array}$
LHS: physics $S_{L}$
RHS: whatever emerges

- Boundary conditions
- Functional integrals heuristic
- Euclidean E-H action not bounded below
- Planck scale

$$
l_{p}=\sqrt{G \hbar}
$$

- cutott?
- NC?
- nee physics?
- Embed $G_{L} \subset C_{L} \quad$ Complex Lorentzian geometries
- Rotate to $G_{L}^{\prime} \subset C_{L} \quad$ Imaginary Lorentzian contour
- Euclideanisation $G_{L}^{\prime} \cong G_{E}$


$$
\int_{G_{L}} e^{i S_{L}} D e \quad \int_{G_{L}^{\prime}} e^{i S_{L}} D l \quad \stackrel{E n c}{=} \int_{G_{E}} e^{-S_{E}} D_{e}
$$

- Include mates fields. Scalars unchanged

Contour rotation

$$
\lim _{R \rightarrow \infty} \int_{0}^{R} e^{i k x} M(x) d x=i \int_{0}^{\infty} e^{-k y} M(i y) d y
$$

Locally,

$$
\int_{0}^{R} e^{i w(x)} d x=\int_{w(0)}^{w(R)} e^{i w} \frac{1}{w^{\prime}(x(w))} d w
$$

$$
\text { Multiple integrals - same } R \text {. }
$$

## Quantum gravity and path integrals

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In order to make sure that one registers this surface term correctly one has to join the initial and final spacelike surfaces by a timelike tube at some large radius $r_{0}$. It is convenient to rotate the time interval on this timelike tube between the two surfaces into the complex plane so that it becomes purely imaginary. This makes the metric on the boundary positive definite so that the path integral can be taken over all positive-definite metrics $g$ that induce the given metric for the boundary.


# PATH INTEGRALS AND THE INDEFINITENESS OF THE GRAVITATIONAL ACTION 

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The Euclidean action for gravity is not positive definite unlike those of scalar and Yang-Mills fields. Indefiniteness arises because conformal transformations can make the

The Euclidean action for gravity is not positive definite unlike those of scalar and Yang-Mills fields. Indefiniteness arises because conformal transformations can make the action arbitrarily negative. In order to make the path integral converge one has to take the contour of integration for the conformal factor to be parallel to the imaginary axis.
... unbounded below

$$
\begin{equation*}
\hat{I}[g]=\frac{-1}{16 \pi G} \int R(g)^{1 / 2} \mathrm{~d}^{4} x-\frac{1}{8 \pi G} \int[K](h)^{1 / 2} \mathrm{~d}^{3} x \tag{1.3}
\end{equation*}
$$

is not positive semi-definite. (The minus sign comes from the direction of the Wick rotation, which has to be chosen to be consistent with that for the matter fields.) Under conformal transformations of the metric $\tilde{g}_{a b}=\Omega^{2} g_{a b}, R$ transforms as

$$
\begin{equation*}
\widetilde{R}=\Omega^{-2} R-6 \Omega^{-3} \Omega \tag{1.4}
\end{equation*}
$$

and

$$
\begin{equation*}
\widetilde{K}=\Omega^{-1} K+3 \Omega^{-2} \Omega_{, a} n^{a}, \tag{1.5}
\end{equation*}
$$

where $n^{a}$ is the unit outward normal to the boundary $\partial \mathrm{M}$. Thus

$$
\begin{equation*}
\hat{I}[\tilde{g}]=\frac{-1}{16 \pi G} \int_{\mathrm{M}} \Omega^{2} R+6 \Omega_{, a} \Omega^{\prime}(g)^{1 / 2} \mathrm{~d}^{4} x-\frac{1}{8 \pi G} \int_{\partial \mathrm{M}}\left[\Omega^{2} K\right](h)^{1 / 2} \mathrm{~d}^{4} x \tag{1.6}
\end{equation*}
$$

One sees that $\hat{I}$ may be as negative, as one wants, by choosing a rapidly varying conformal factor $\Omega$.

$$
\begin{aligned}
\text { But: } & \text { Plank sate limits freperang } \Rightarrow \delta_{E} \text { bounded below } \\
& \text { Recall } \delta_{c c} \geq 0 \text { ! }
\end{aligned}
$$

This is not

Wick Rotation


Metrics physically ditterent!

Hawking, Cargese lectures 1978

## EUCLIDEAN QUANTUM GRAVITY

I feel that one should adopt a similar Euclidean approach in quantum gravity and supergravity. Of course one cannot simply replace the time coordinates by imaginary quantities because there is no preferred set of time coordinates in general relativity. Instead I think one should perform the path integrals over all positive definite metrics, most of which will not admit a section on which the metric is real and Lorentzian, and then analytically continue the result of the path integral, if necessary. In order to restrict the path integral to positive definfte metrics and to exclude integration over metrics with Lorentzian or ultra hyperbolic signatures, one should probably integrate not over the components of the metric $g_{a b}$ but over the components $e^{a}{ }_{m}$ of a tetrad. This can be regarded as the square root of the metric

$$
\begin{equation*}
g_{a b}=e_{a}^{m} e_{b m} \tag{1.4}
\end{equation*}
$$

## Frame field

| 4-manifold | $M$ |
| :--- | :--- |
| Vector fields | $l_{0}, l_{1}, l_{2}, l_{3}$ |
| Spin connection | $\nabla$ |

Minkowski metric $\eta=\left(\begin{array}{cccc}-1 & 0 & 0 & 0 \\ 0 & 1 & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1\end{array}\right)$
$\rightarrow$ Functions

$$
\begin{array}{ll}
\sigma_{a b}^{c}: & \nabla_{l_{l}} l_{b}=\sigma_{a b}^{c} l_{c} \\
c_{a b}^{c}: & {\left[l_{a}, l_{b}\right]=c_{a b}^{c} l_{c}}
\end{array}
$$



Curvature $\quad R\left(l_{a}, l_{b}\right) l_{c}=\left(l_{a}\left(\sigma_{b c}^{e}\right)+\sigma_{b c}^{d} \sigma_{a d}^{e}-l_{b}\left(\sigma_{a c}^{e}\right)-\sigma_{a c}^{d} \sigma_{b d}^{e}-c_{a b}^{d} \sigma_{d c}^{e}\right) l_{e}$

On $G_{L}^{\prime}$, fields $l_{0}=i e_{o}, l_{1}=e_{1}, l_{2}=e_{2}, l_{3}=e_{3}, \quad e$ real

$$
\sigma_{b c}^{a}=i^{n} \omega_{b c}^{a}, n=\delta_{b 0}+\delta_{c 0}-\delta_{d 0}, \quad \omega \text { real }
$$

Actions

$$
\begin{aligned}
& S_{L}=\int-2 \Lambda+\frac{R}{16 \pi G} d V+S_{L}^{\text {scalar }} \\
& S_{E}=\int 2 \Lambda-\frac{R_{E}}{16 \pi G} d V_{E}+S_{E}^{\text {scalar }}
\end{aligned}
$$

Euclideanisation
(c.f. Samuel 2015, D'Andrea, Kurkov, Lizzi 2016)

On $G_{L}^{\prime}: \quad i S_{L}=S_{E}$
$\psi \in C^{4}$
Euclidean inner product $\left(\psi^{\prime}, \psi\right)_{E}$
Lorentzian $\left(\psi^{\prime}, \psi\right)=\left(\gamma^{0} \psi^{\prime}, \psi\right)_{E}$
Gamma matrices $\gamma^{a} \gamma^{b}+\gamma^{b} \gamma^{a}=-2 \eta^{a b}$
Chirality $\quad \gamma, \quad \gamma^{2}=1$
Charge conjugation $J: C^{4} \rightarrow C^{4}$ antilinear

$$
\gamma= \pm 1 L / R
$$

$$
v^{2}=1
$$

$$
J \gamma=-\gamma J
$$

Spinors on manifold

$$
\begin{aligned}
& \left\langle\psi^{\prime}, \psi\right\rangle=\int_{M}\left(\psi^{\prime}, \psi\right) d V \\
& D=i \gamma^{a} \nabla_{l_{a}}
\end{aligned}
$$

Euclidean spinors

$$
\begin{aligned}
& \psi(x) \text { same } \\
& \gamma \quad \text { same } \\
& J_{E}=\gamma^{0} J
\end{aligned}
$$

## bit rousts files tiny

$$
D_{E}=\gamma_{E}^{a} \nabla_{E e_{a}}
$$

no i!
$D_{E}^{*}=D_{E}$
$u \sin y(,)_{E}$

On $G_{L}^{\prime}: \quad l_{0}=i e_{0}$, etc.
Calculation: $D=i D_{E}$

$$
\begin{aligned}
&\langle\psi, D \psi\rangle=\int(J \bar{\psi}, D \psi) d V \quad d e t \\
&=\int\left(\gamma^{0} J \bar{\psi}, i D_{E} \psi_{E}\right)_{E},(-i) d V_{E} \\
&=\int\left(J_{E} \bar{\psi}, D_{E} \psi\right) d V_{E} \\
& \text { Sones } f_{0, m n} l_{a}
\end{aligned}
$$

